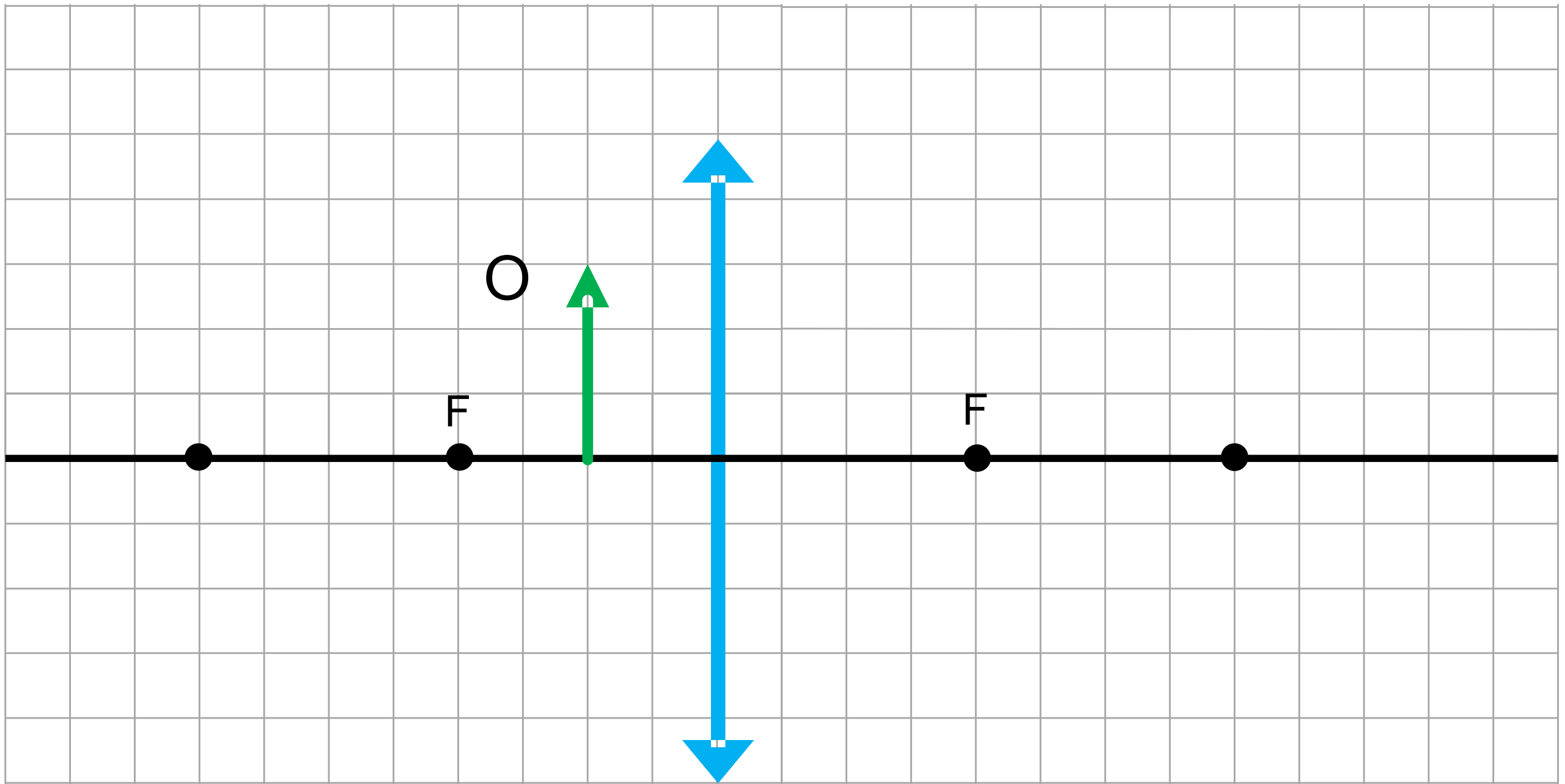


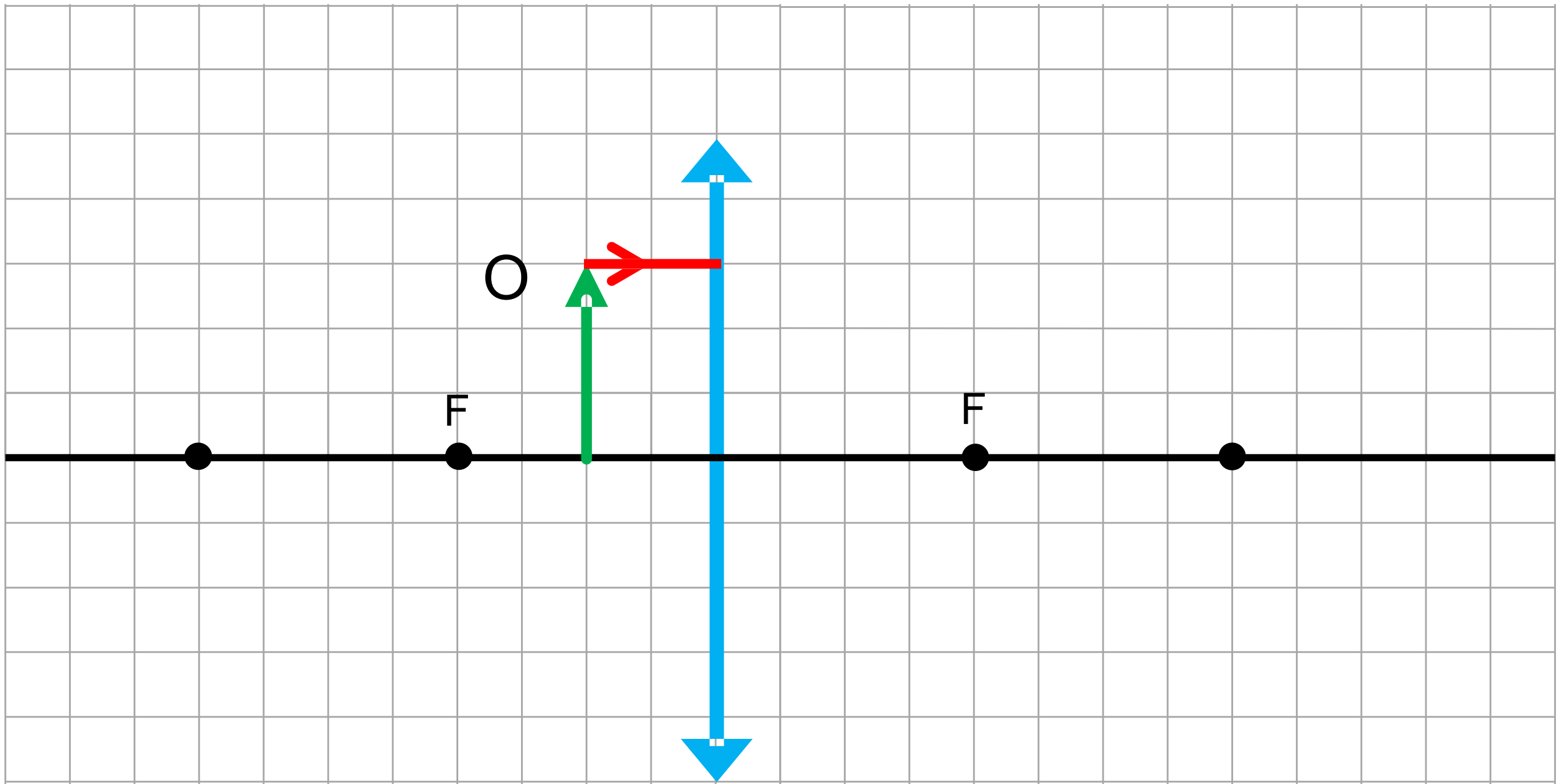
INSTRUMENTOS ÓTICOS

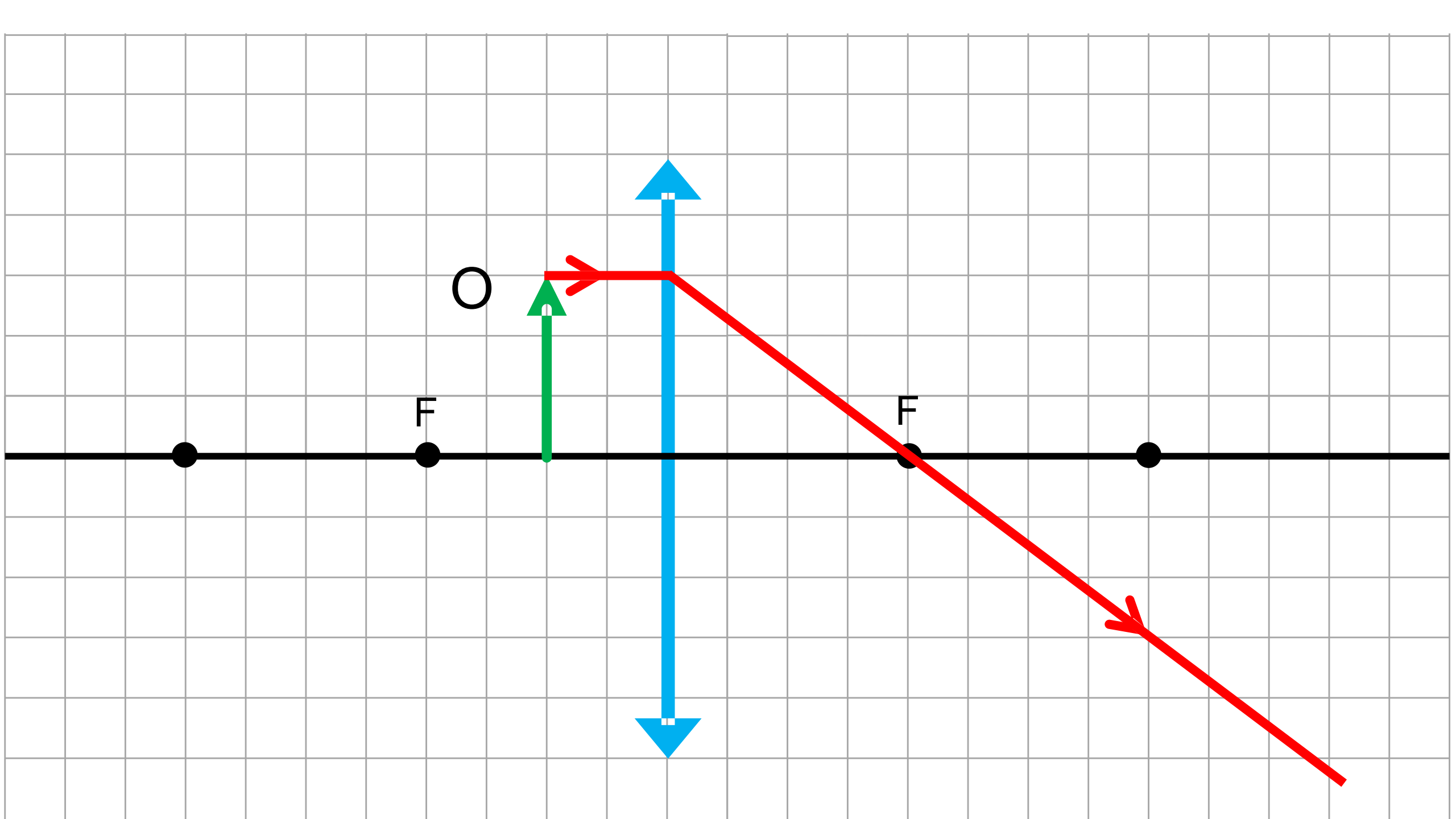
FÍSICA

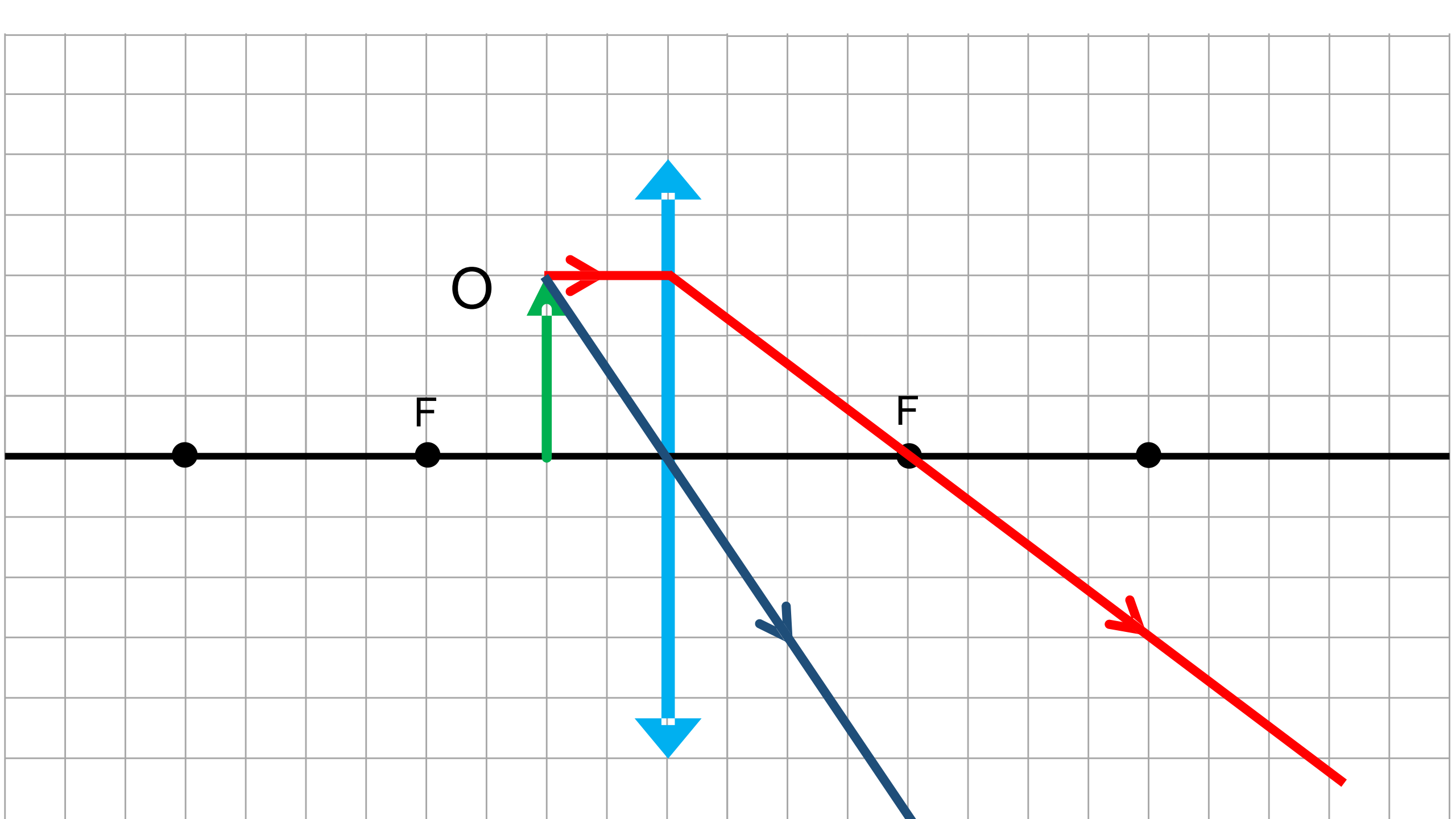
Professor Danilo

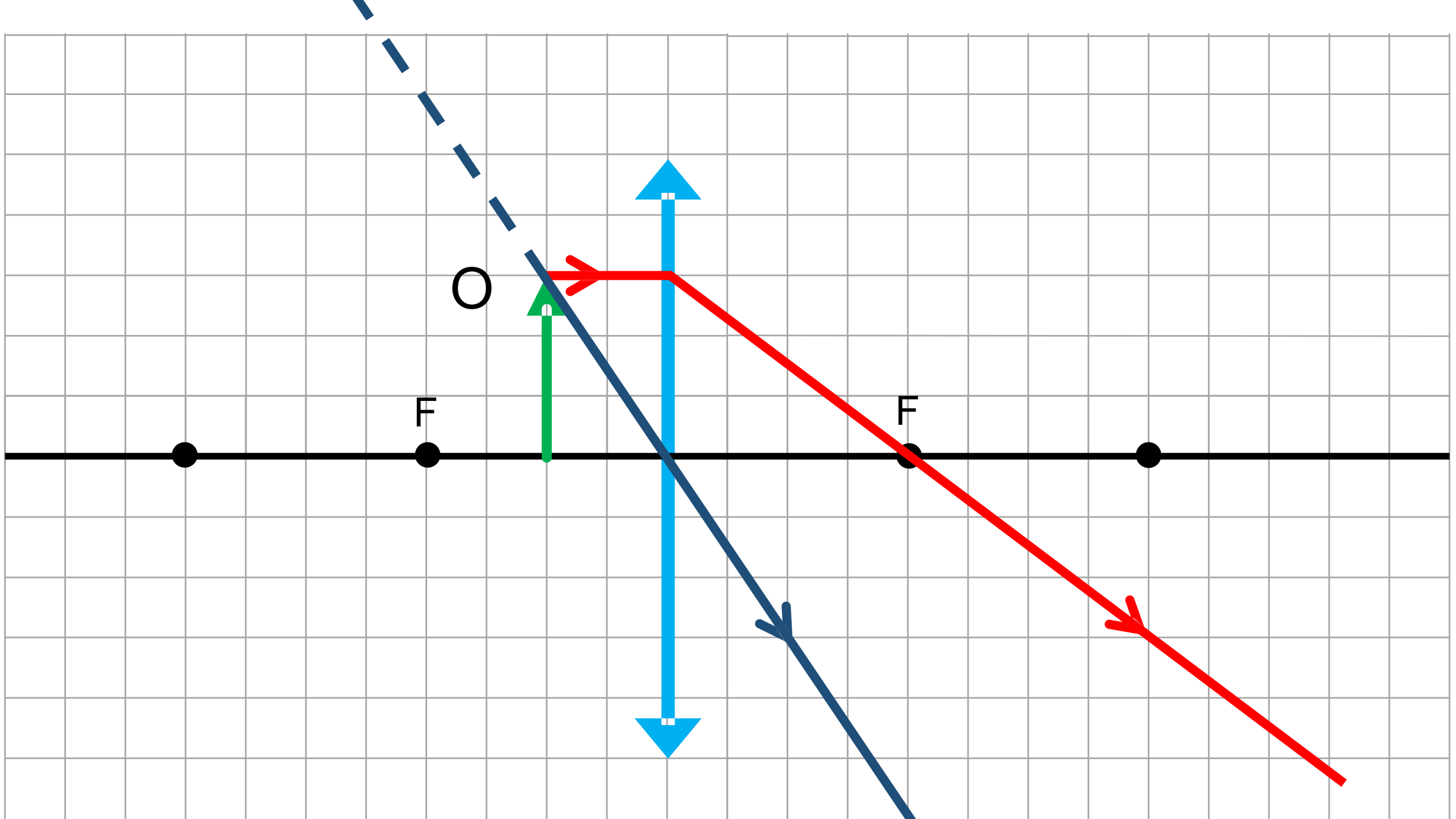
Q. 12 - Lupa

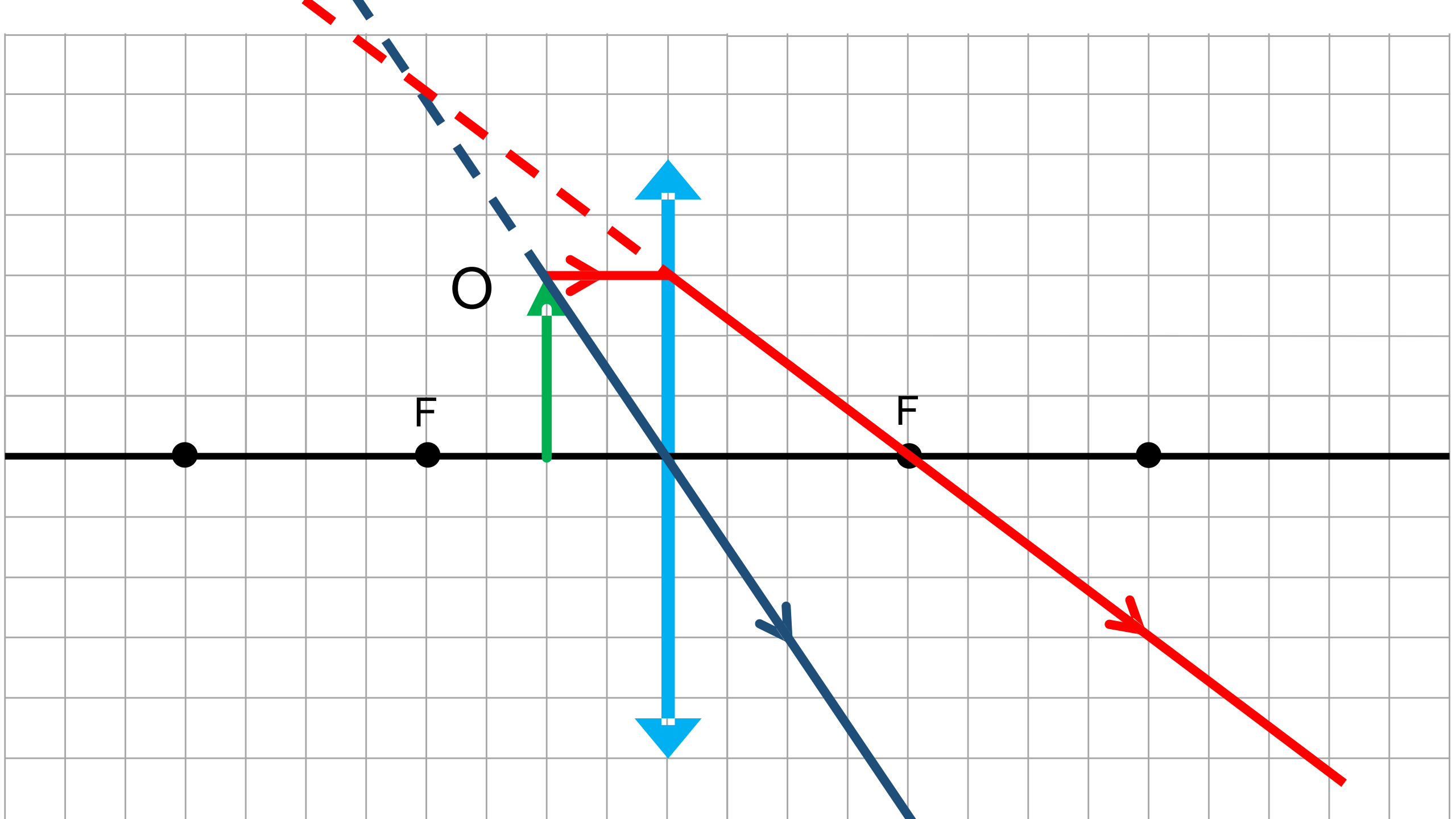


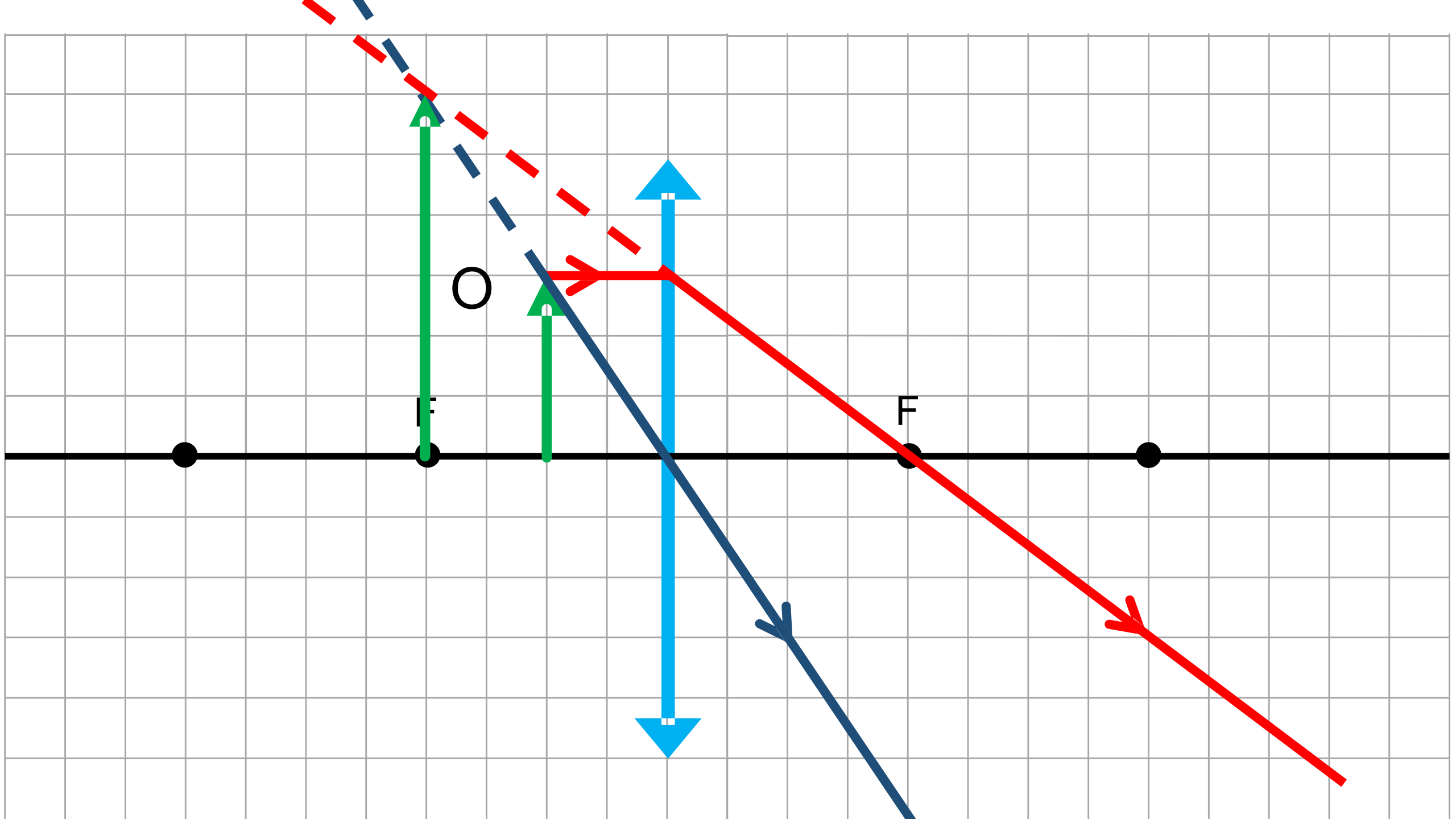












LUPA

LUPA

- IMAGEM:

LUPA

- IMAGEM:
 - VIRTUAL

LUPA

- IMAGEM:
 - VIRTUAL
 - DIREITA

LUPA

- IMAGEM:
 - VIRTUAL
 - DIREITA
 - MAIOR

LUPA

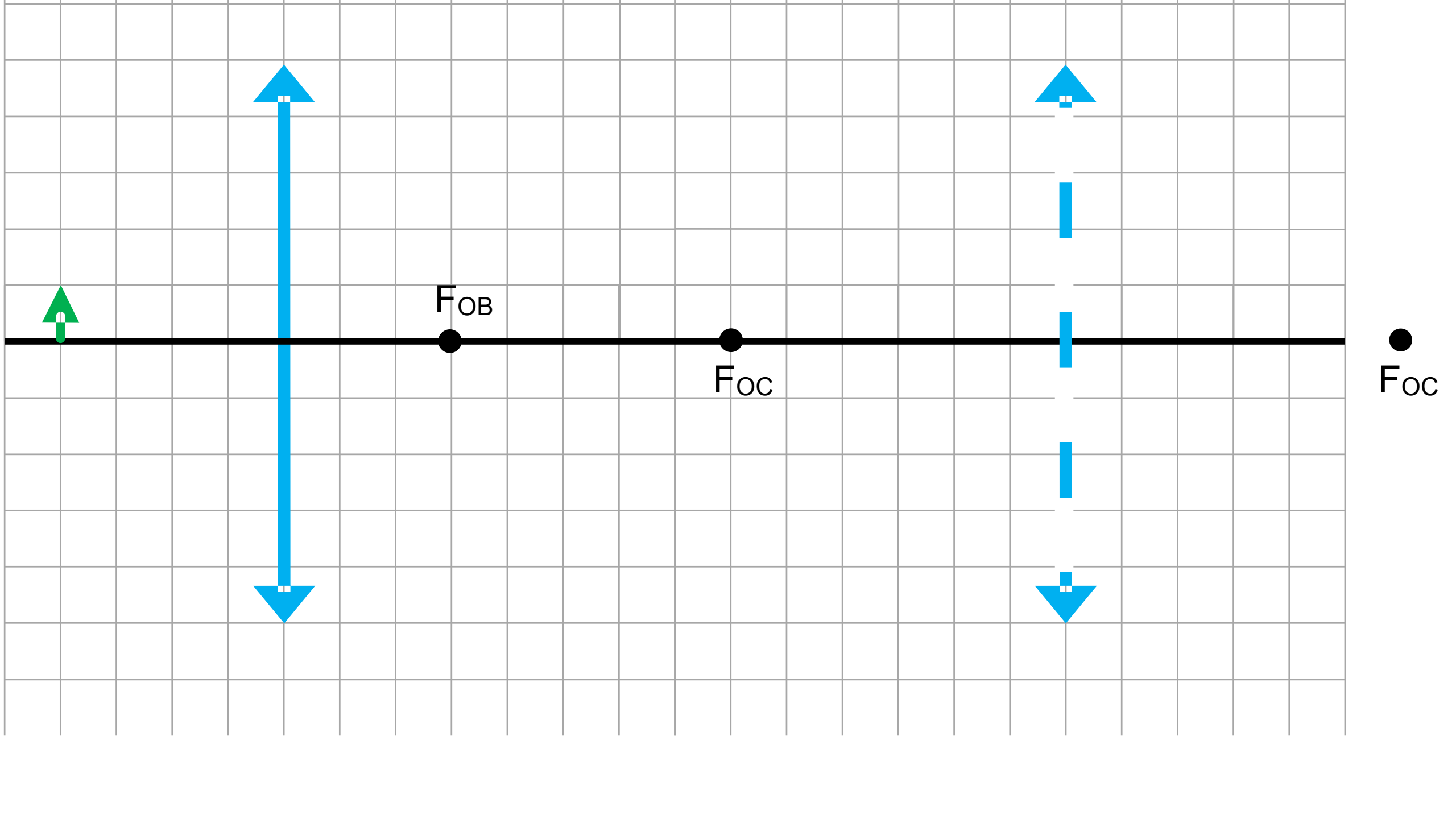
- IMAGEM:
 - VIRTUAL
 - DIREITA
 - MAIOR
 - Mais distante da lente que o objeto

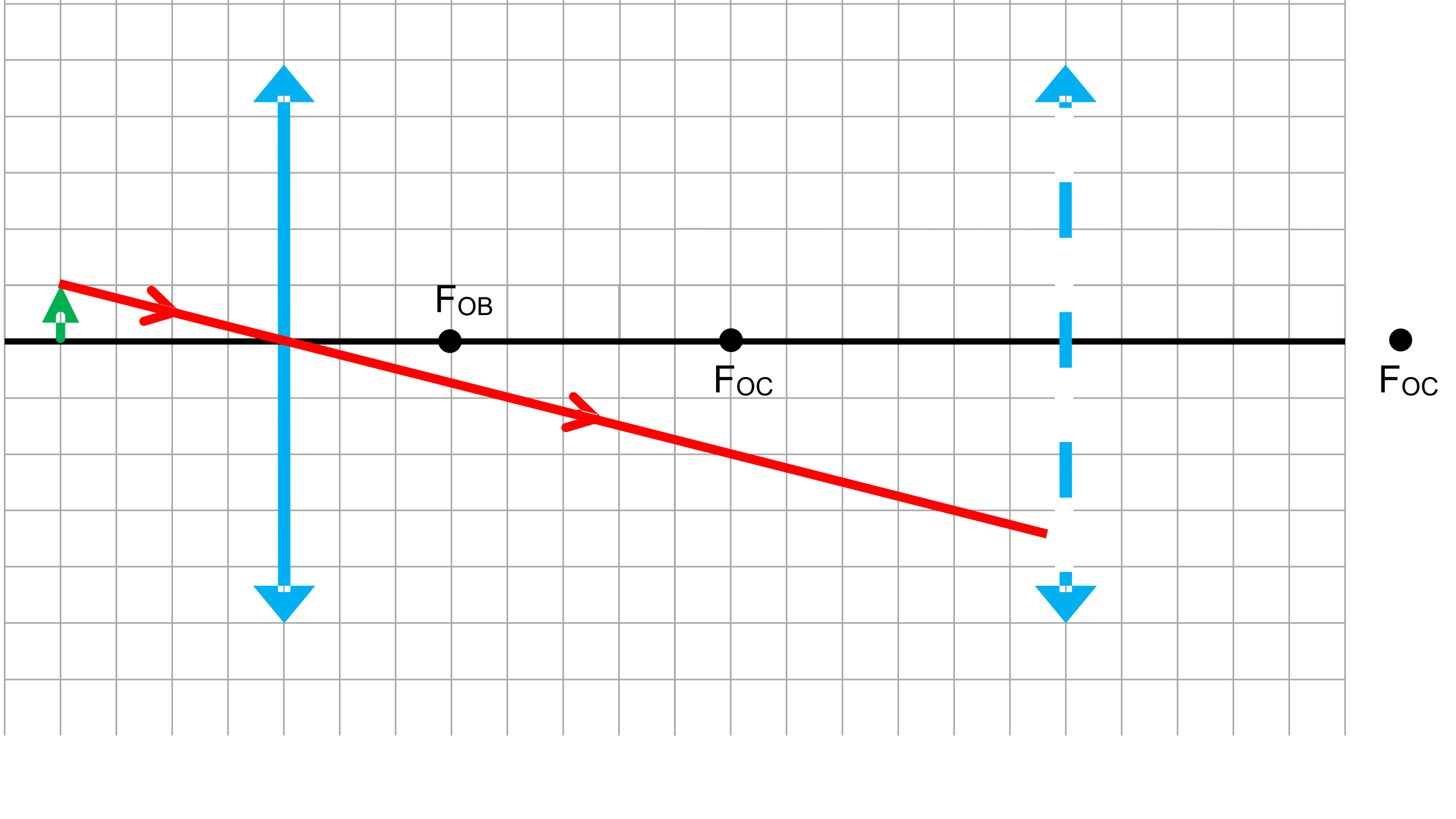
LUPA

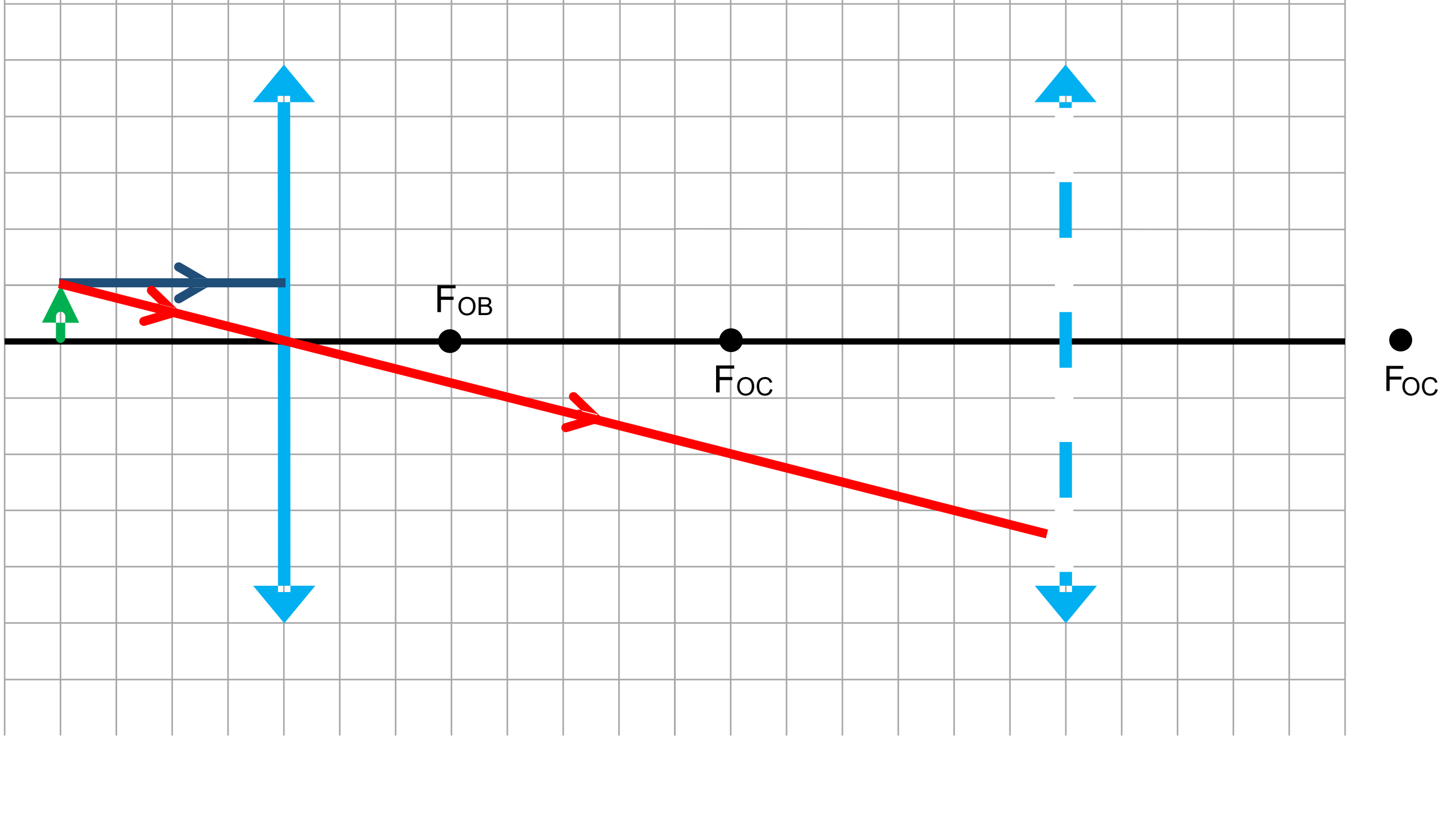
- IMAGEM:
 - VIRTUAL
 - DIREITA
 - MAIOR
 - Mais distante da lente que o objeto
- Qualquer lente convergente pode servir como lupa

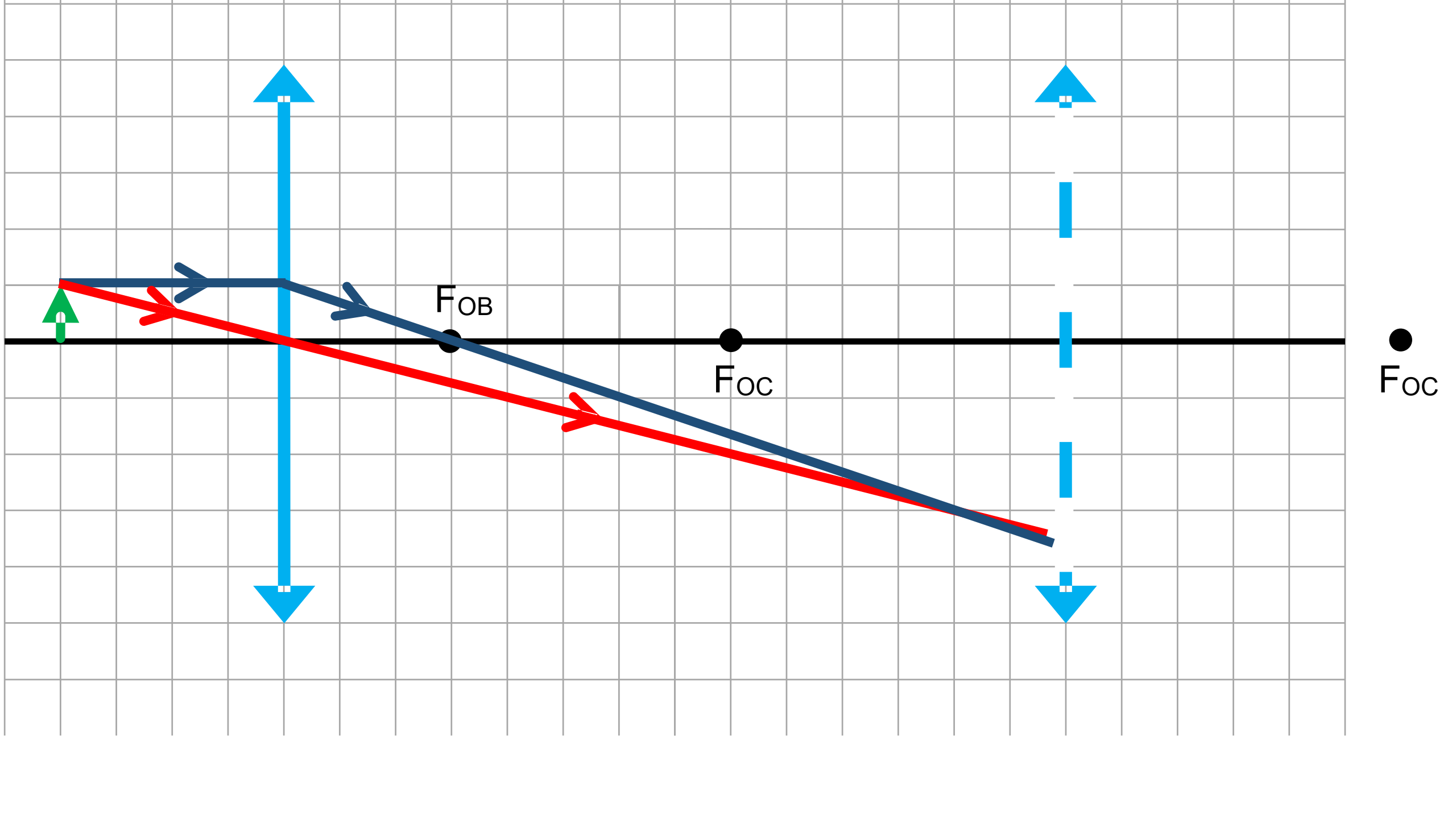


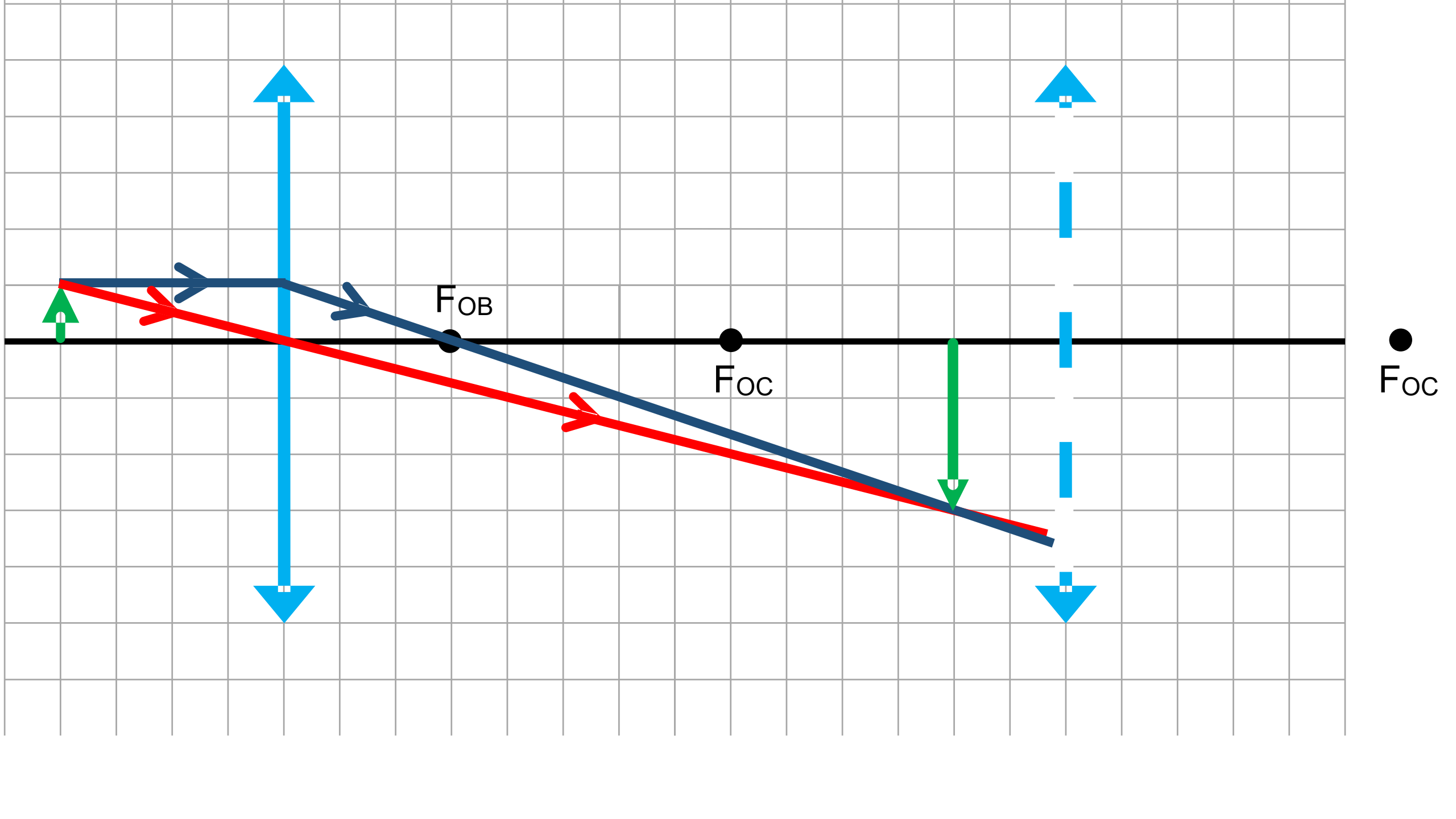
Q. 13 - MICROSCÓPIO

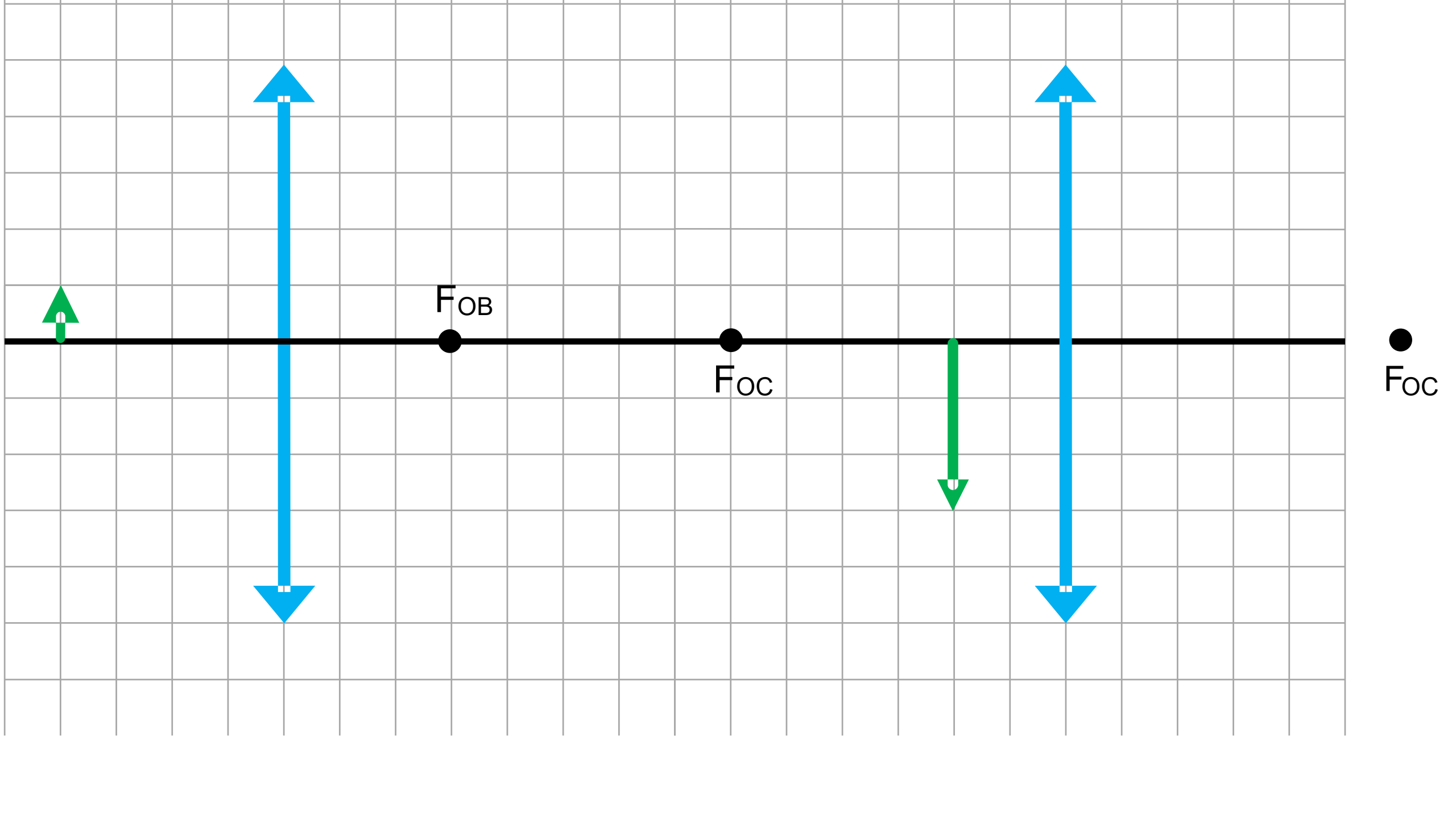


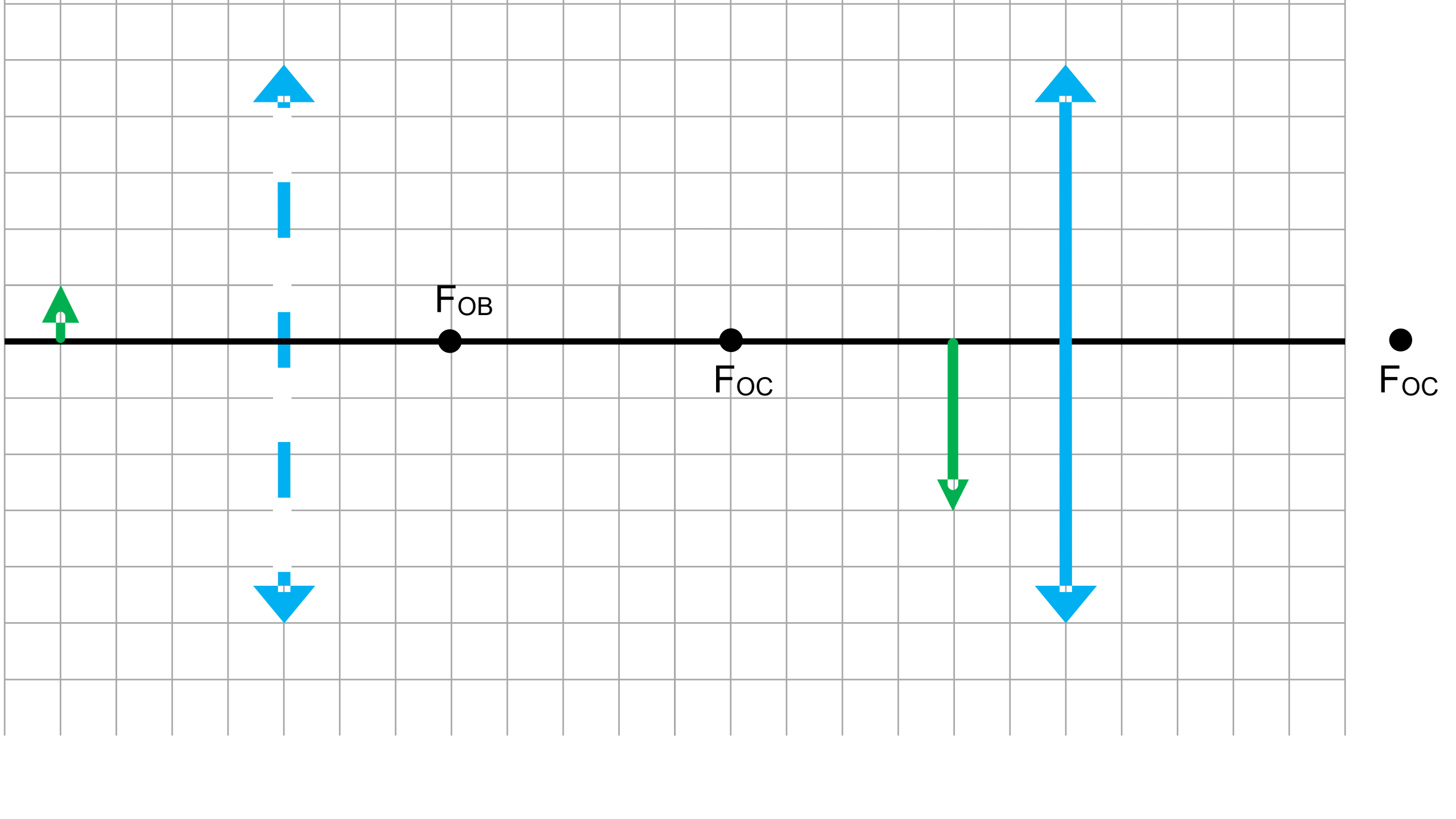


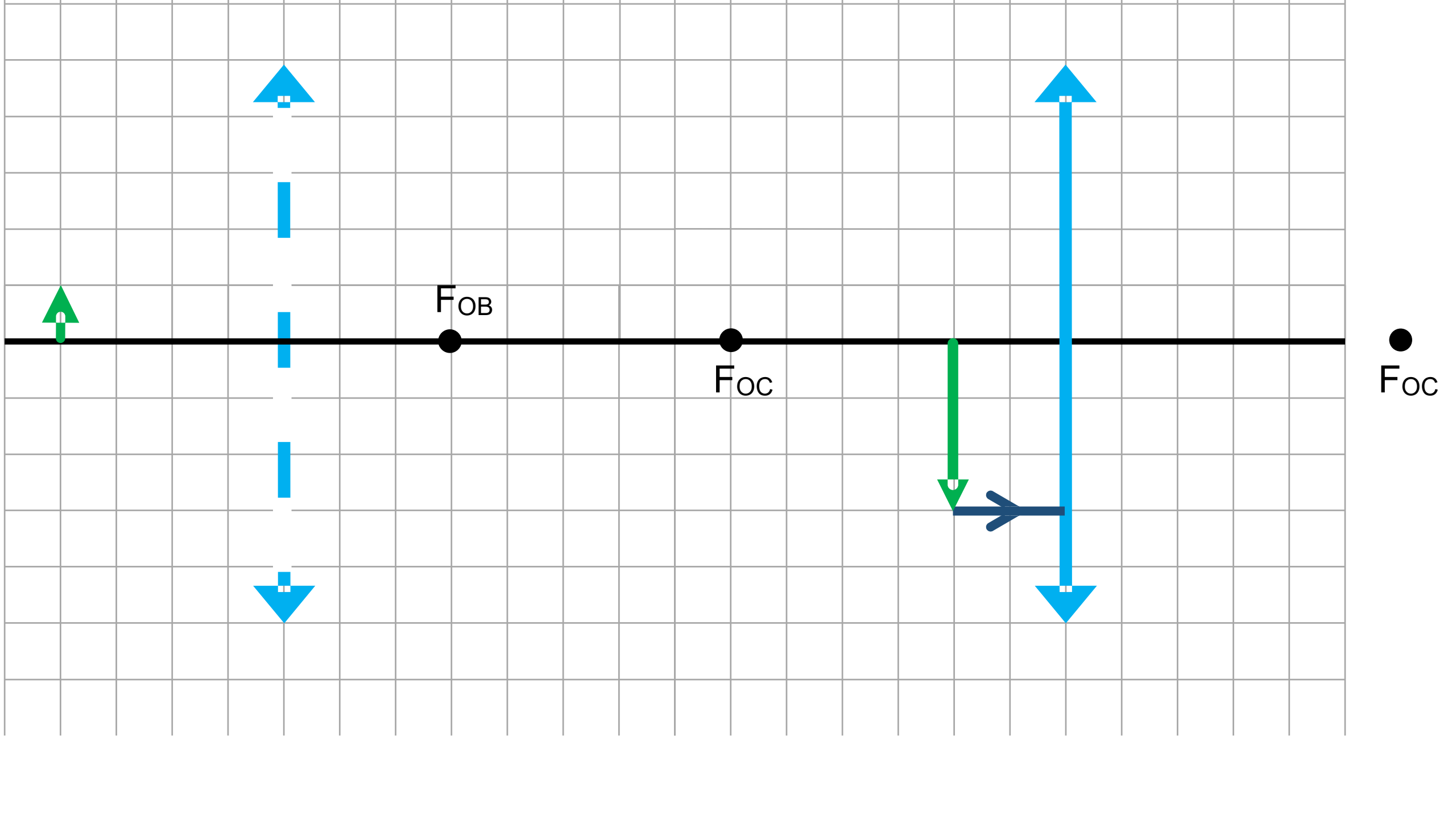


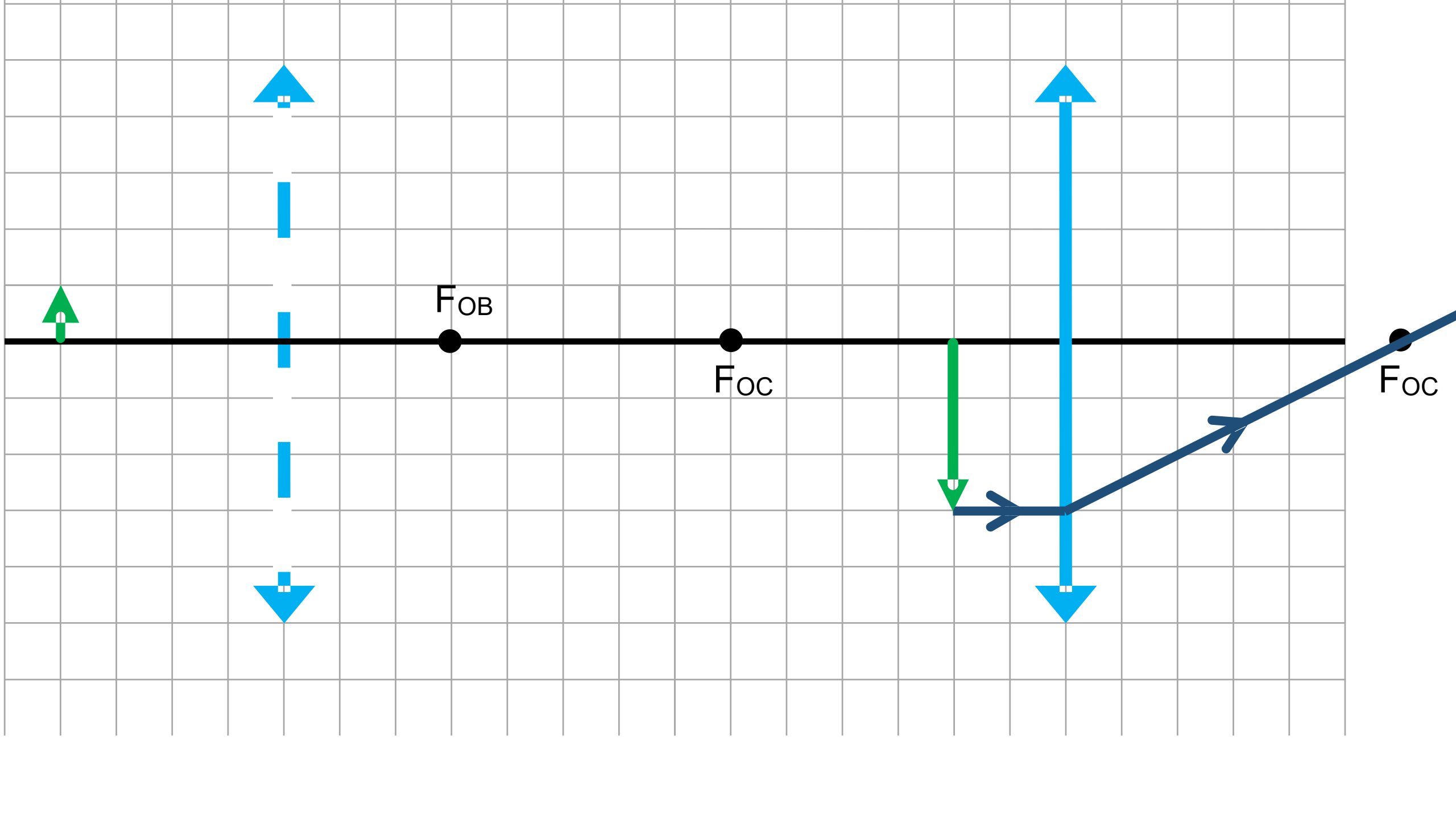


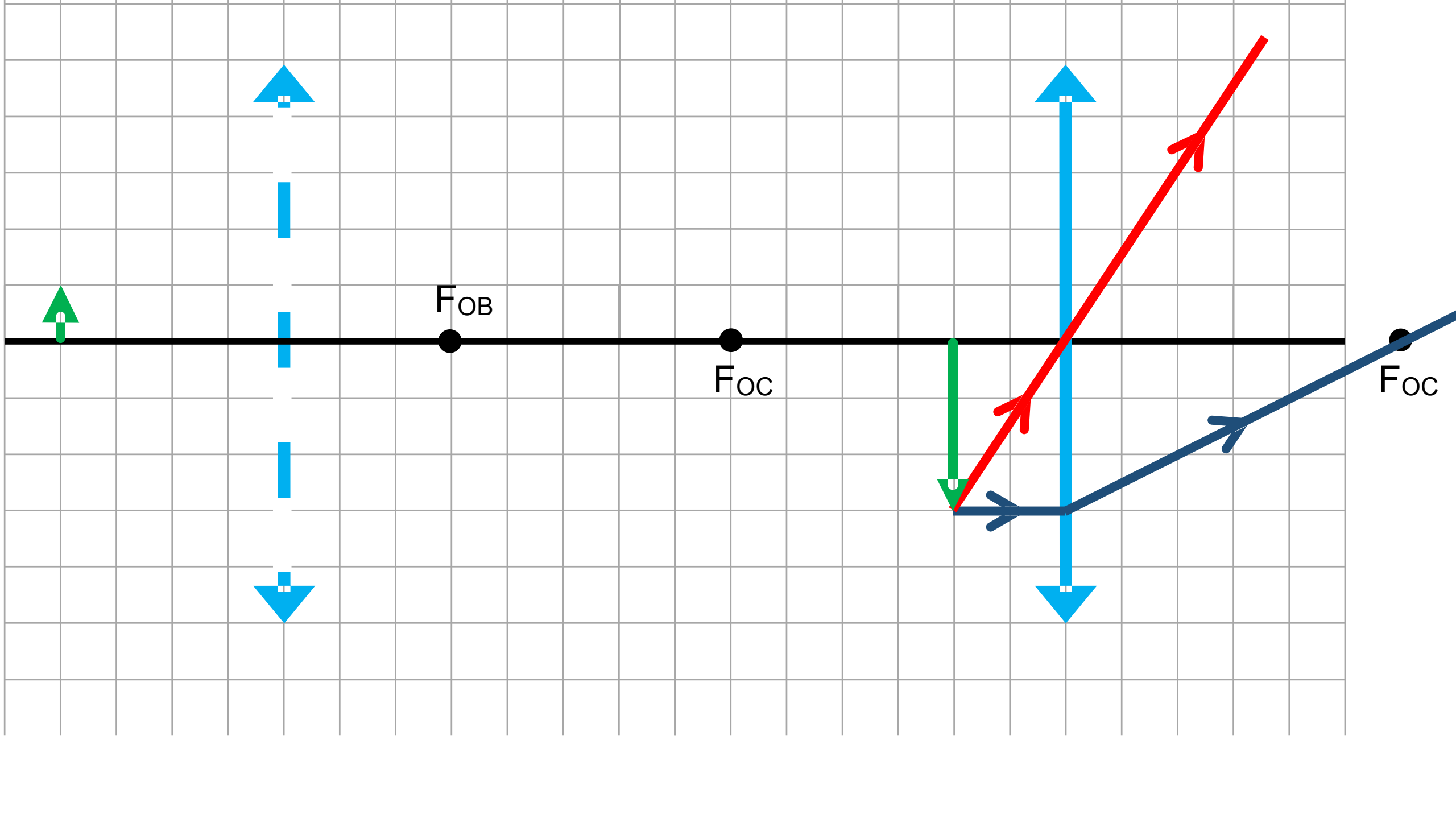


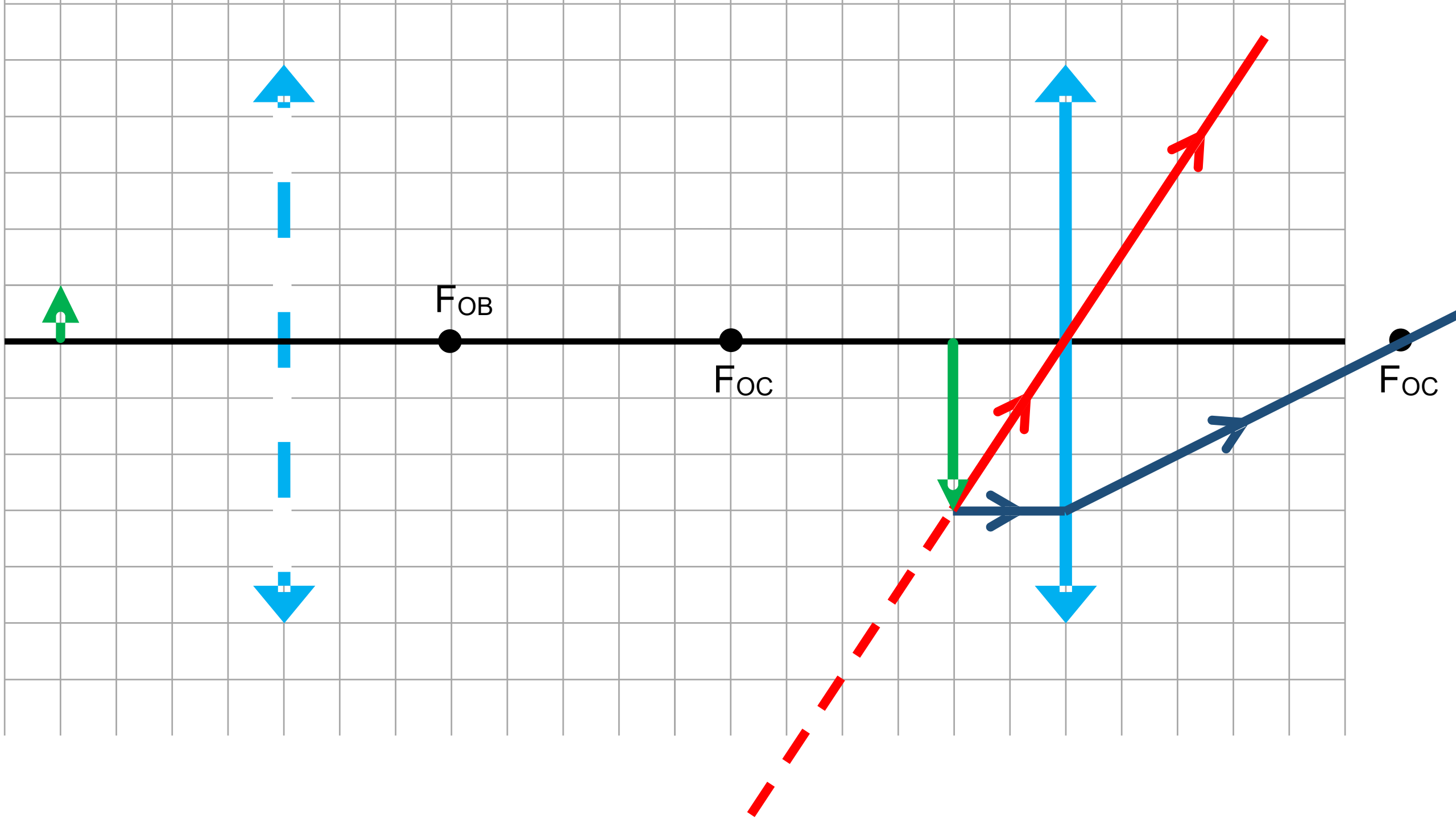


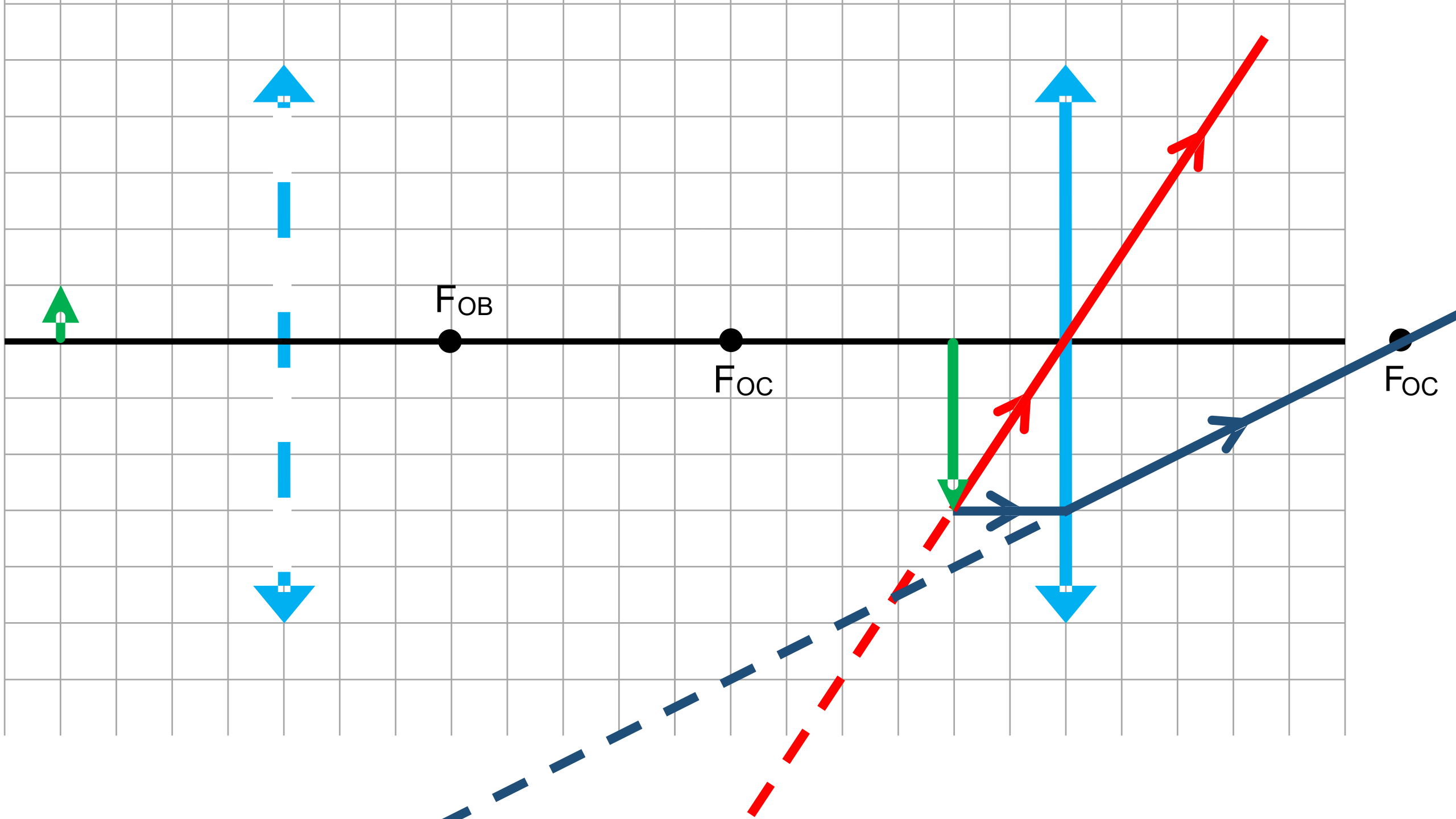


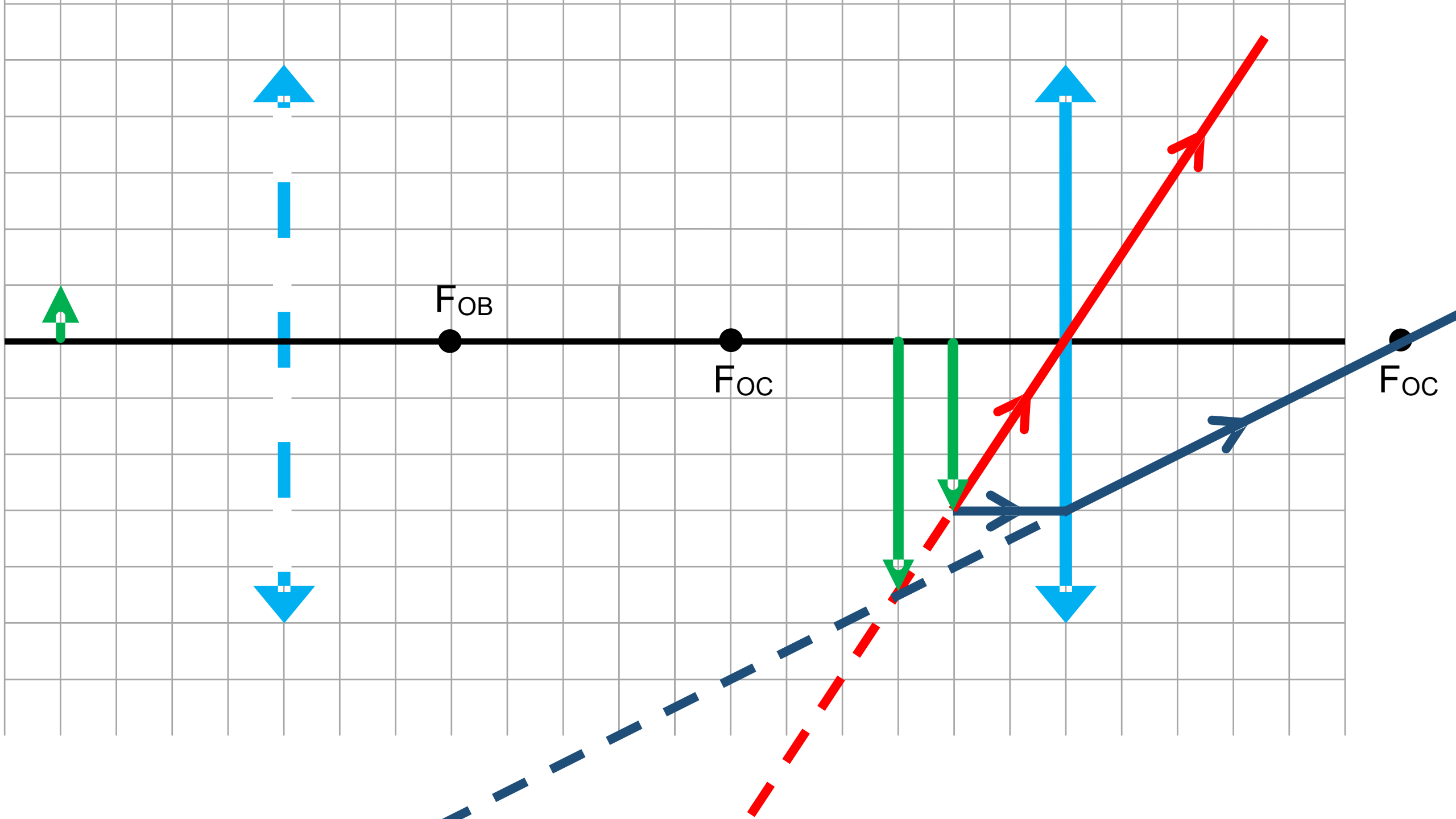


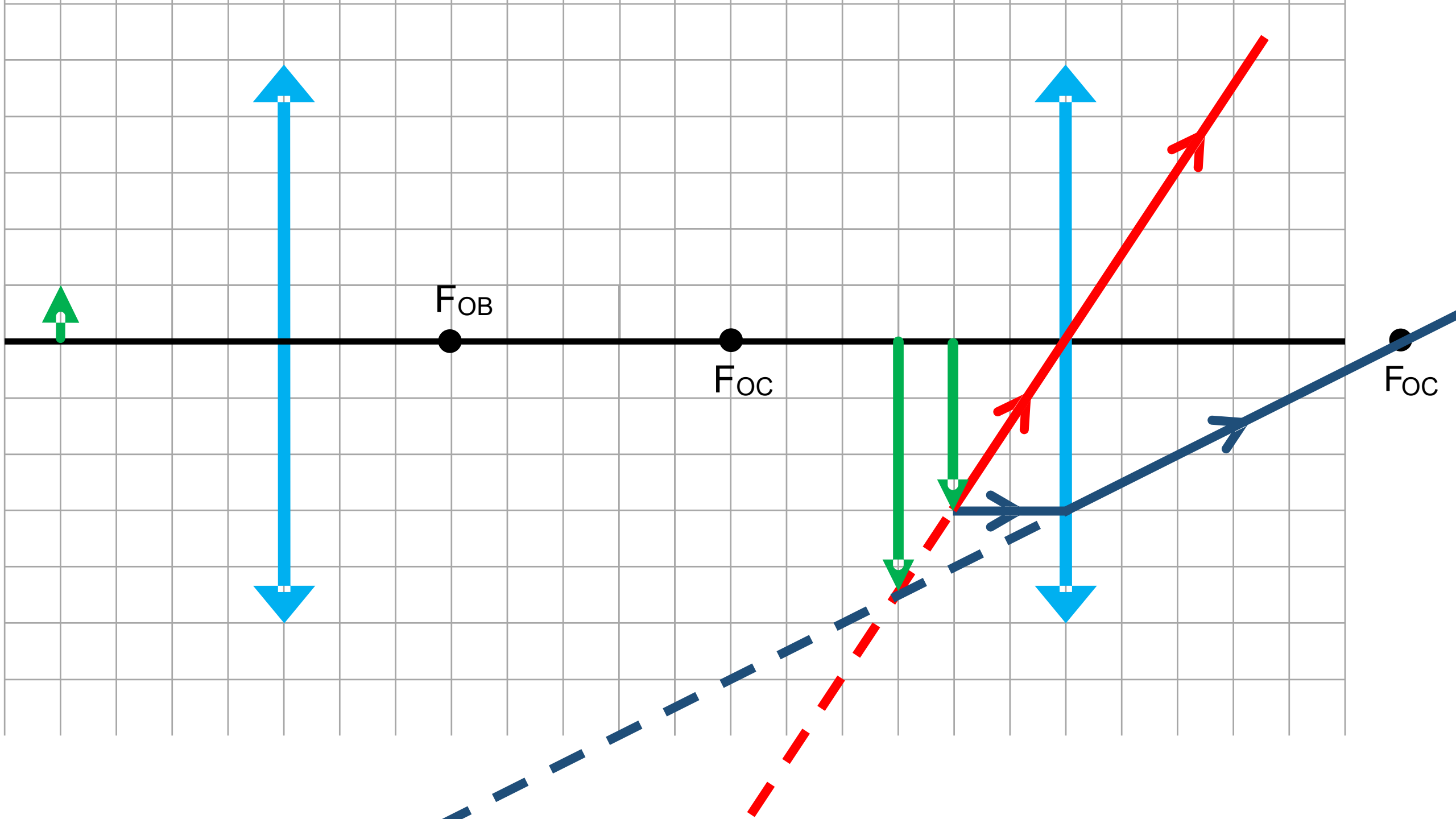


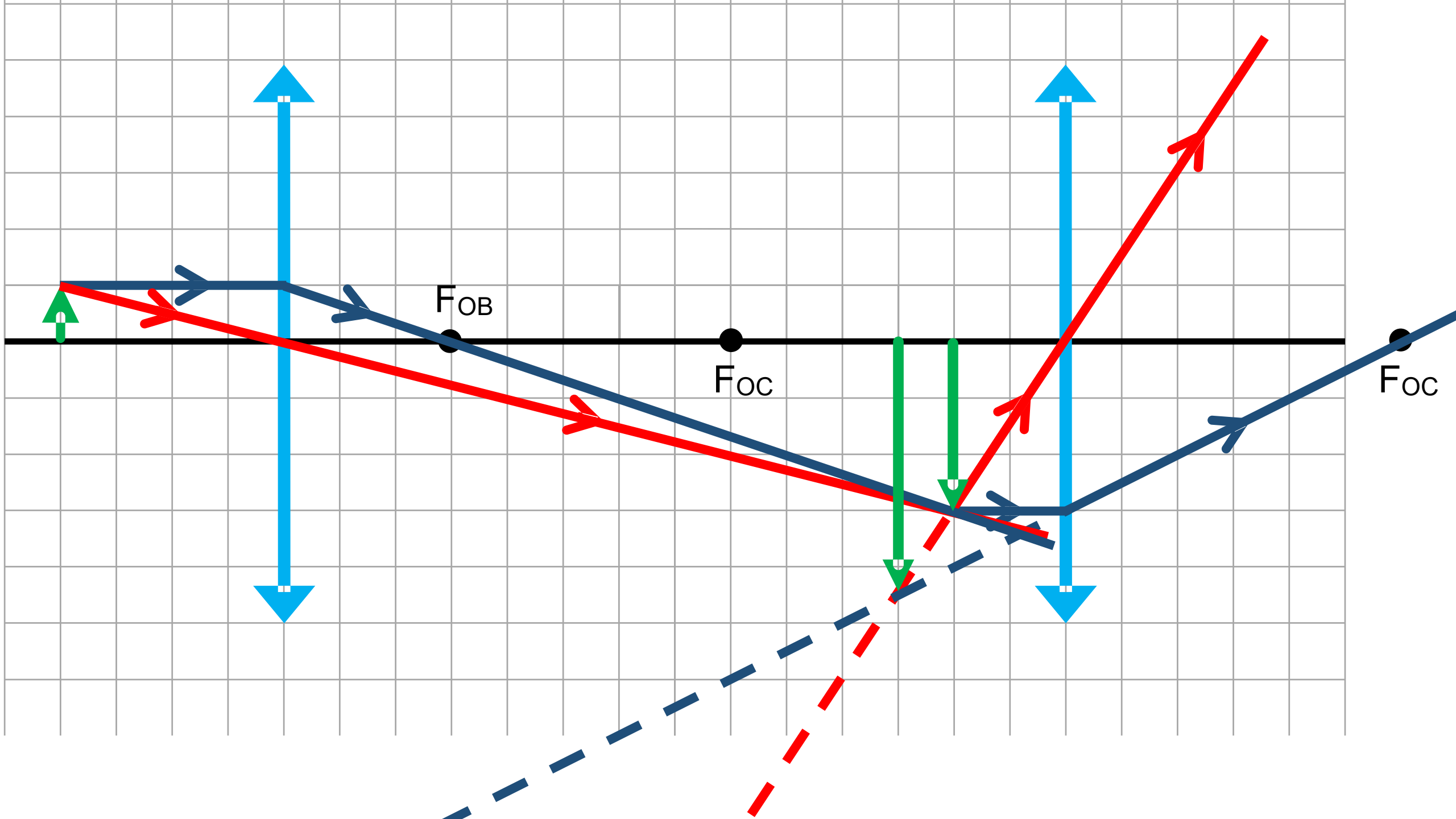












MICROSCÓPIO

MICROSCÓPIO

- IMAGEM:

MICROSCÓPIO

- IMAGEM:
 - VIRTUAL

MICROSCÓPIO

- IMAGEM:
 - VIRTUAL
 - INVERTIDA

MICROSCÓPIO

- IMAGEM:
 - VIRTUAL
 - INVERTIDA
 - MAIOR

MICROSCÓPIO

- IMAGEM:
 - VIRTUAL
 - INVERTIDA
 - MAIOR
 - AUMENTO:

MICROSCÓPIO

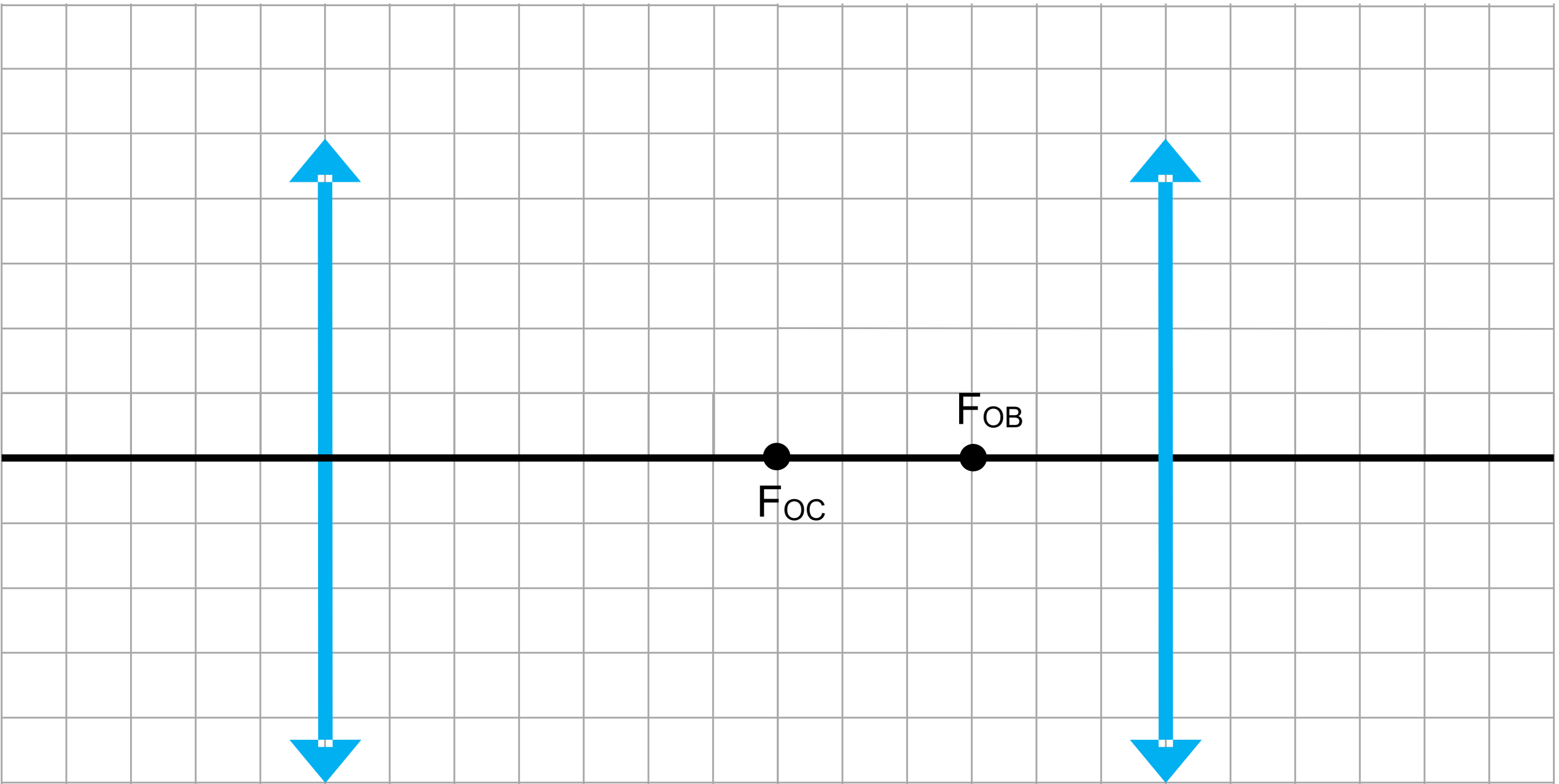
- IMAGEM:

- VIRTUAL
- INVERTIDA
- MAIOR

- AUMENTO: $A = A_{OB} \cdot A_{OC}$



Q. 14 - LUNETAS ASTRONÔMICAS



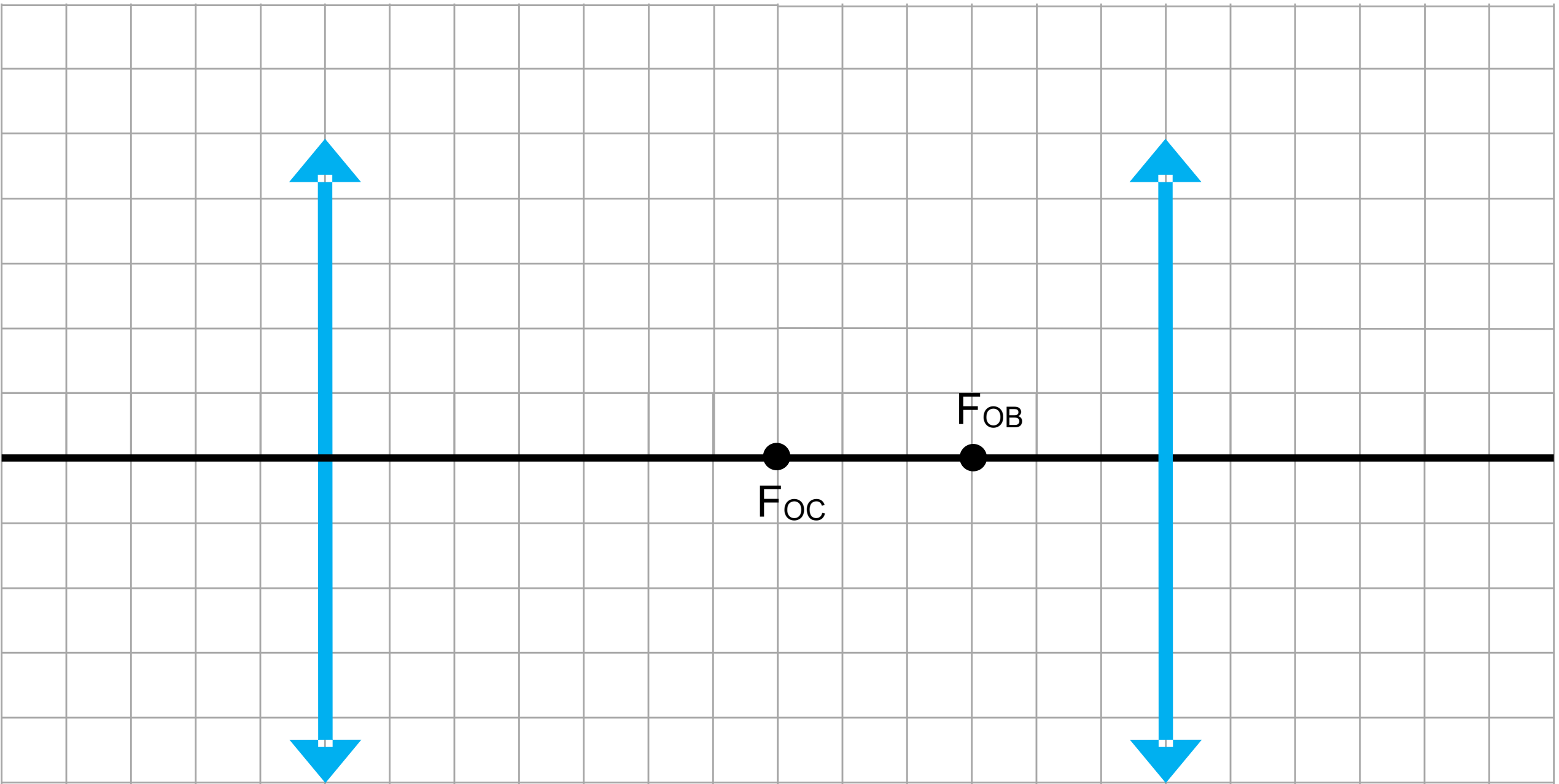
Objetiva

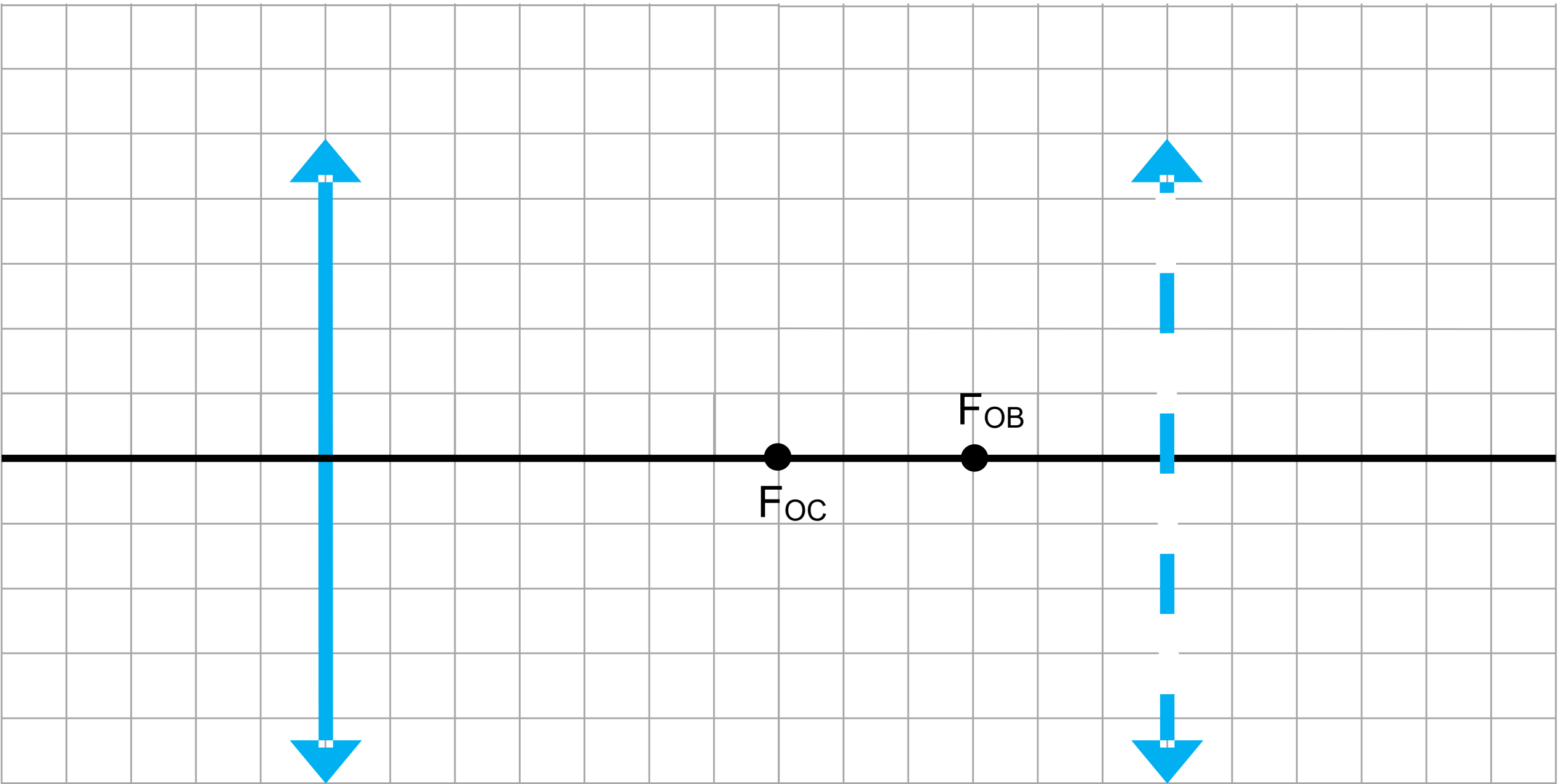
Ocular

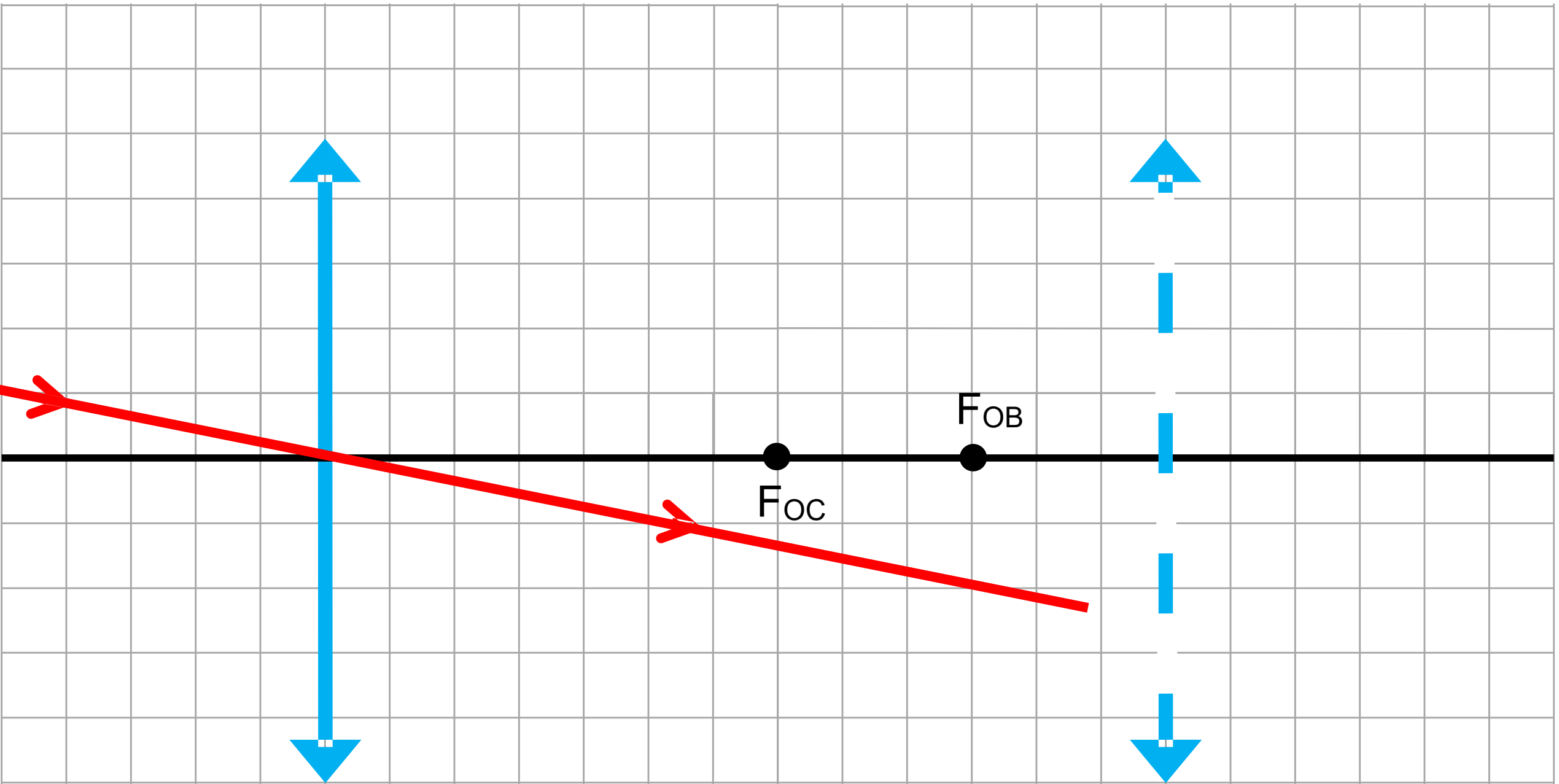
F_{oc}

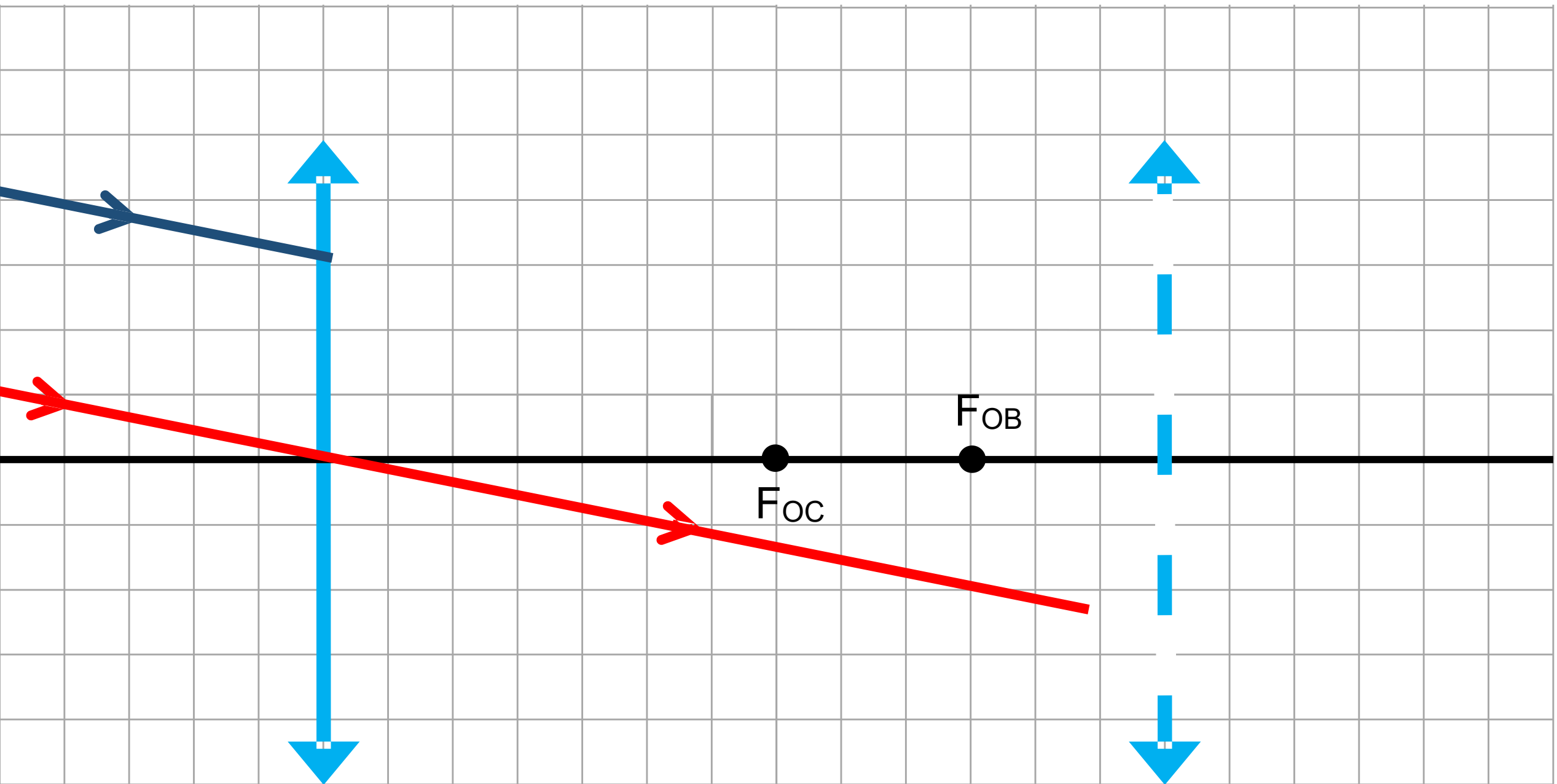
F_{OB}

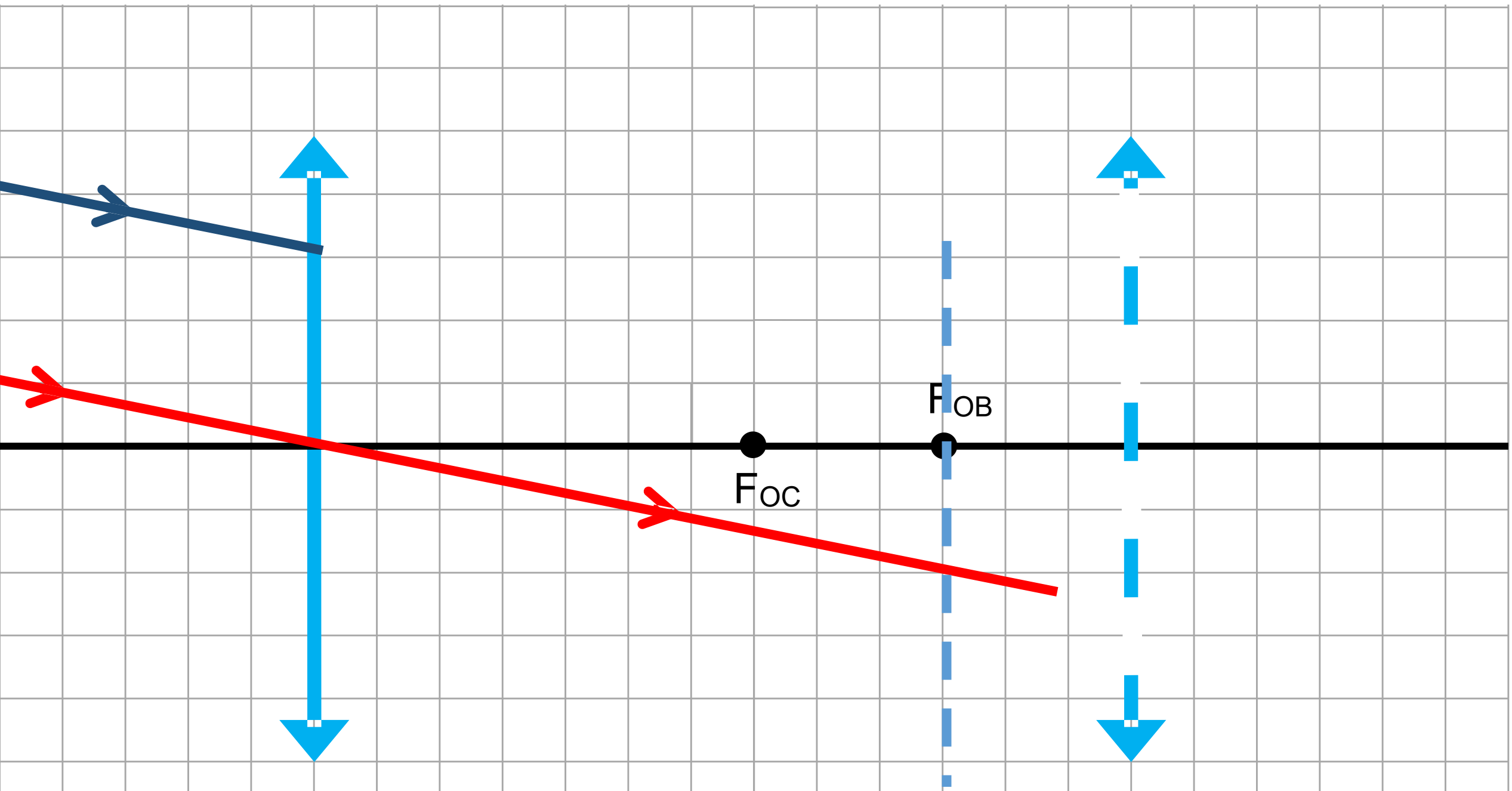


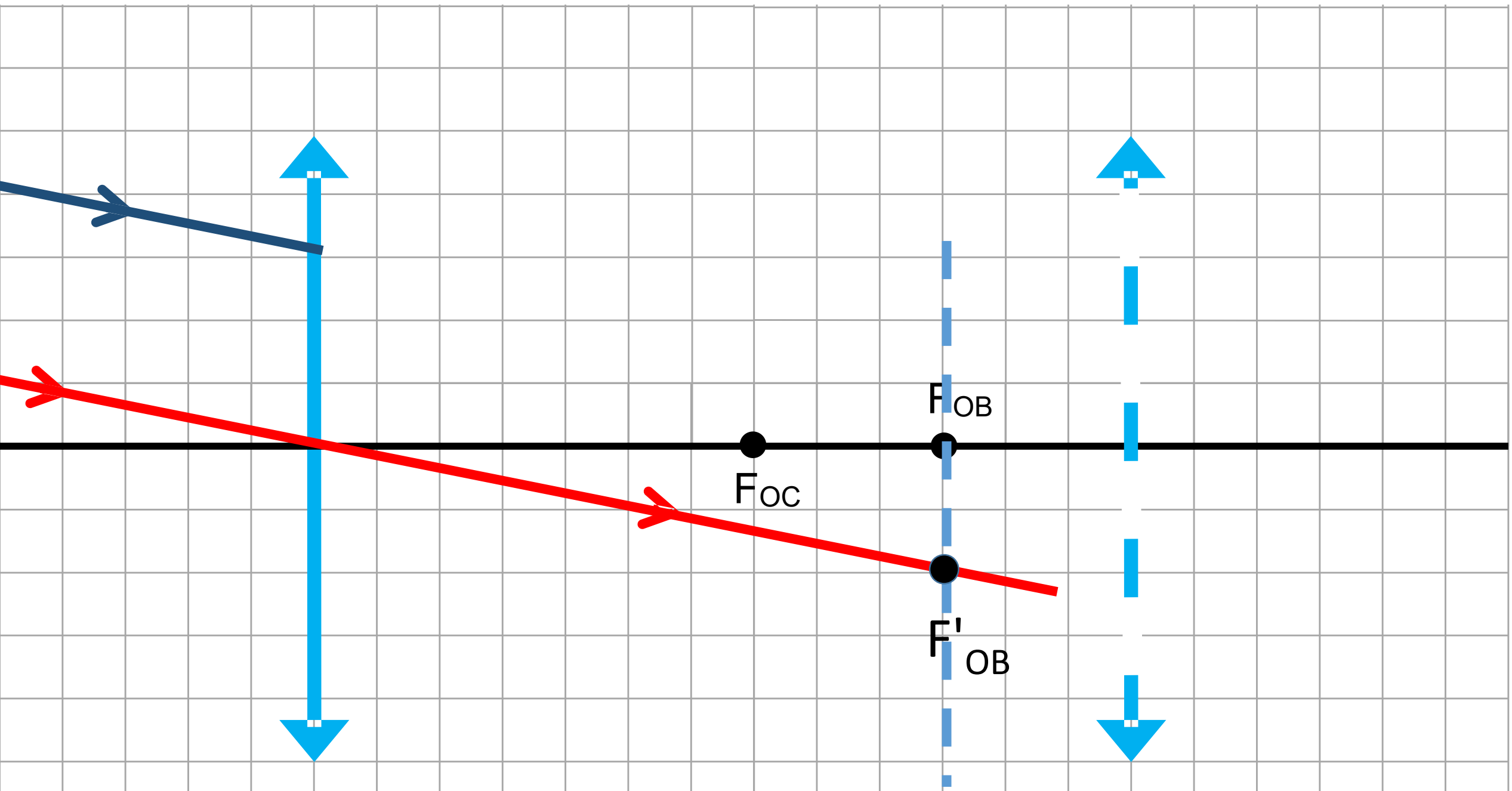


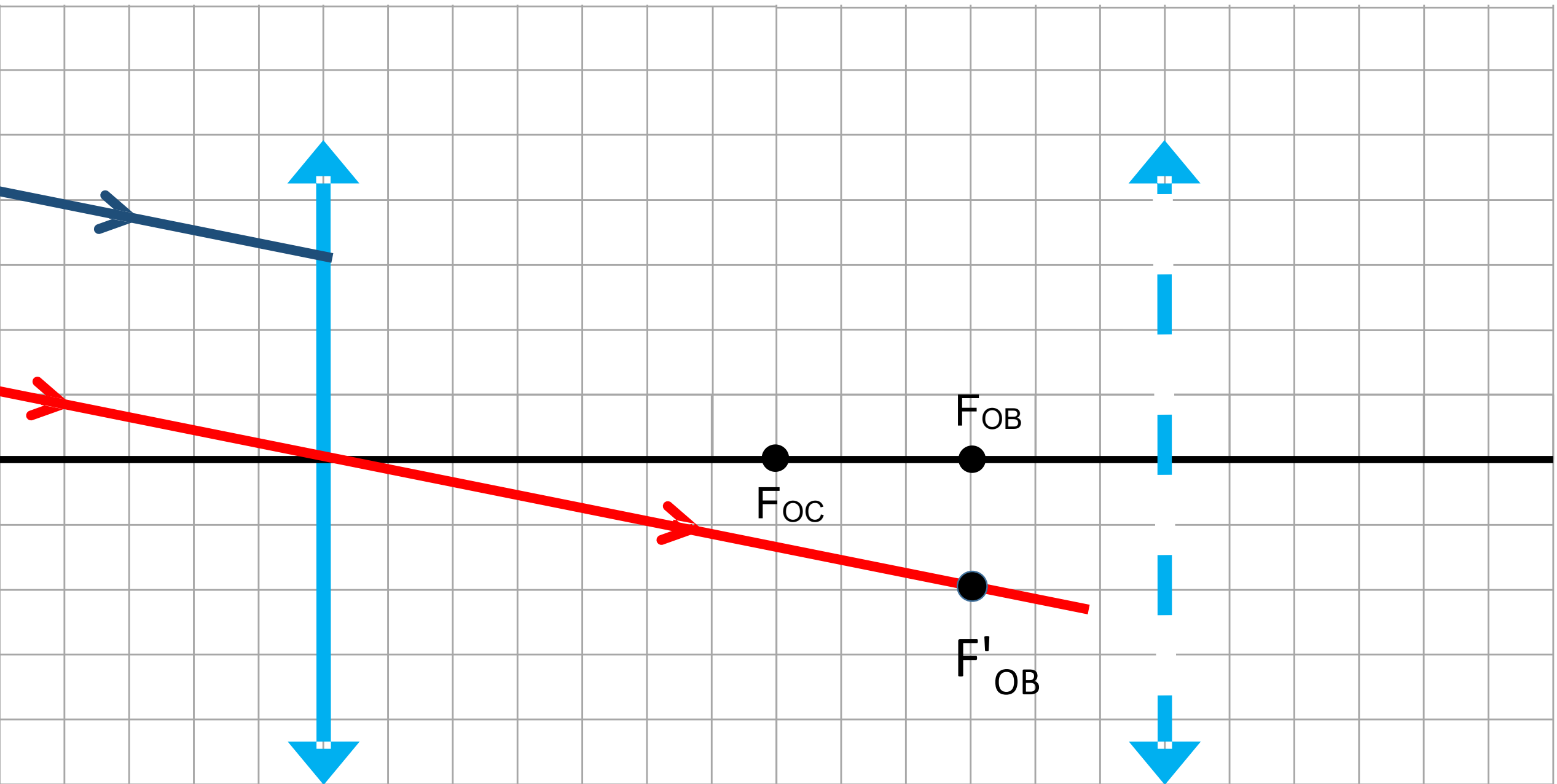


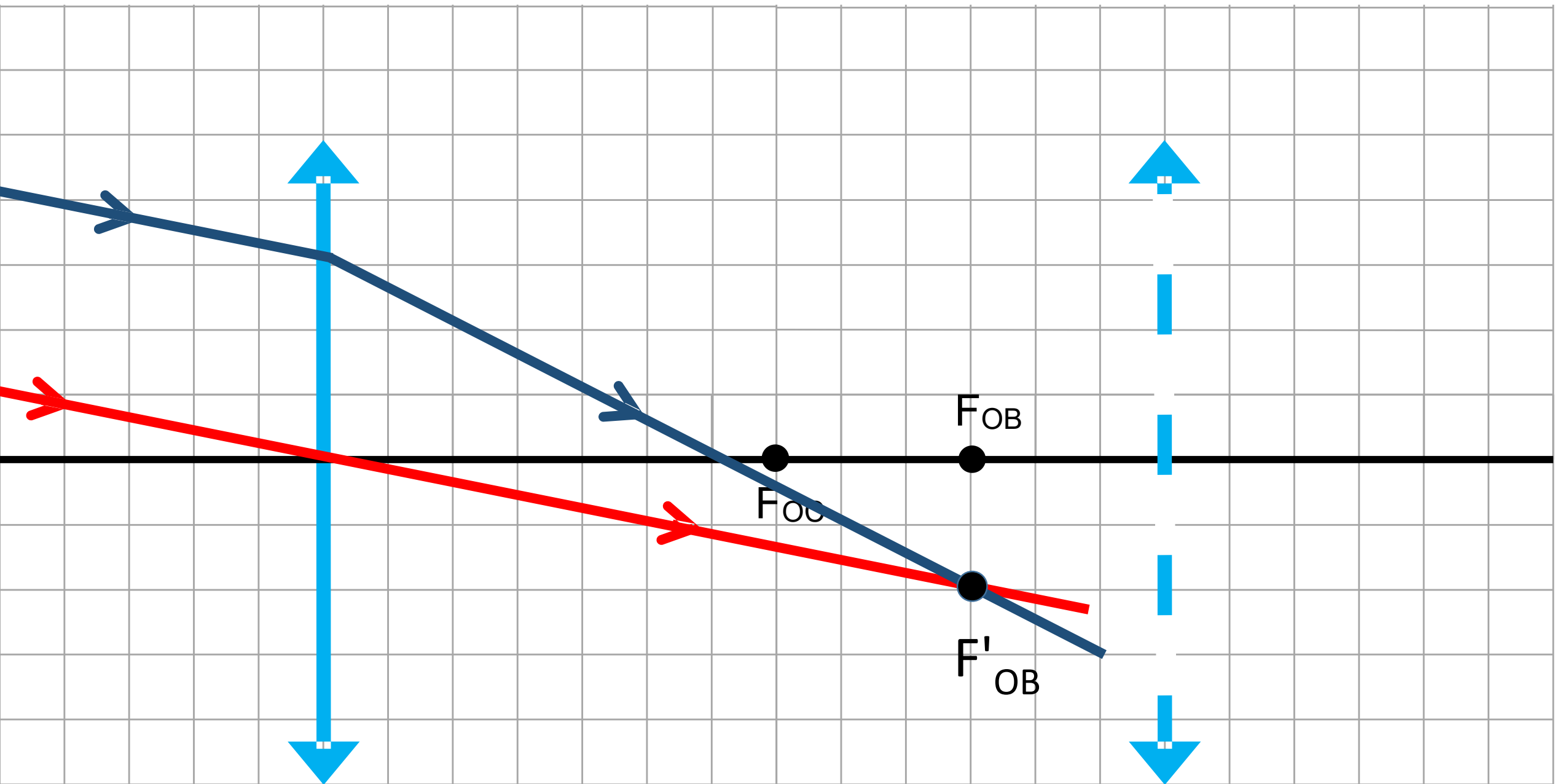


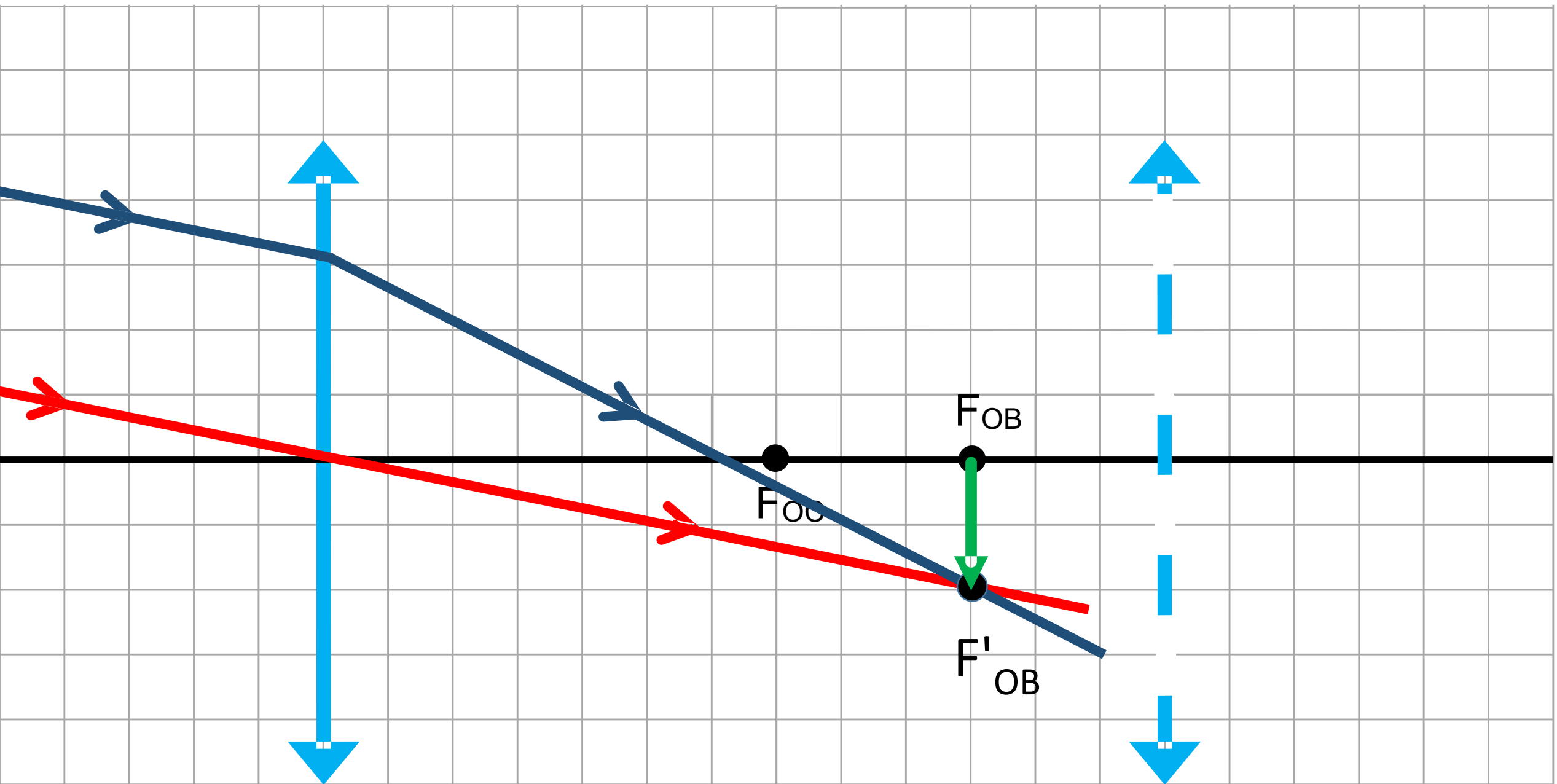


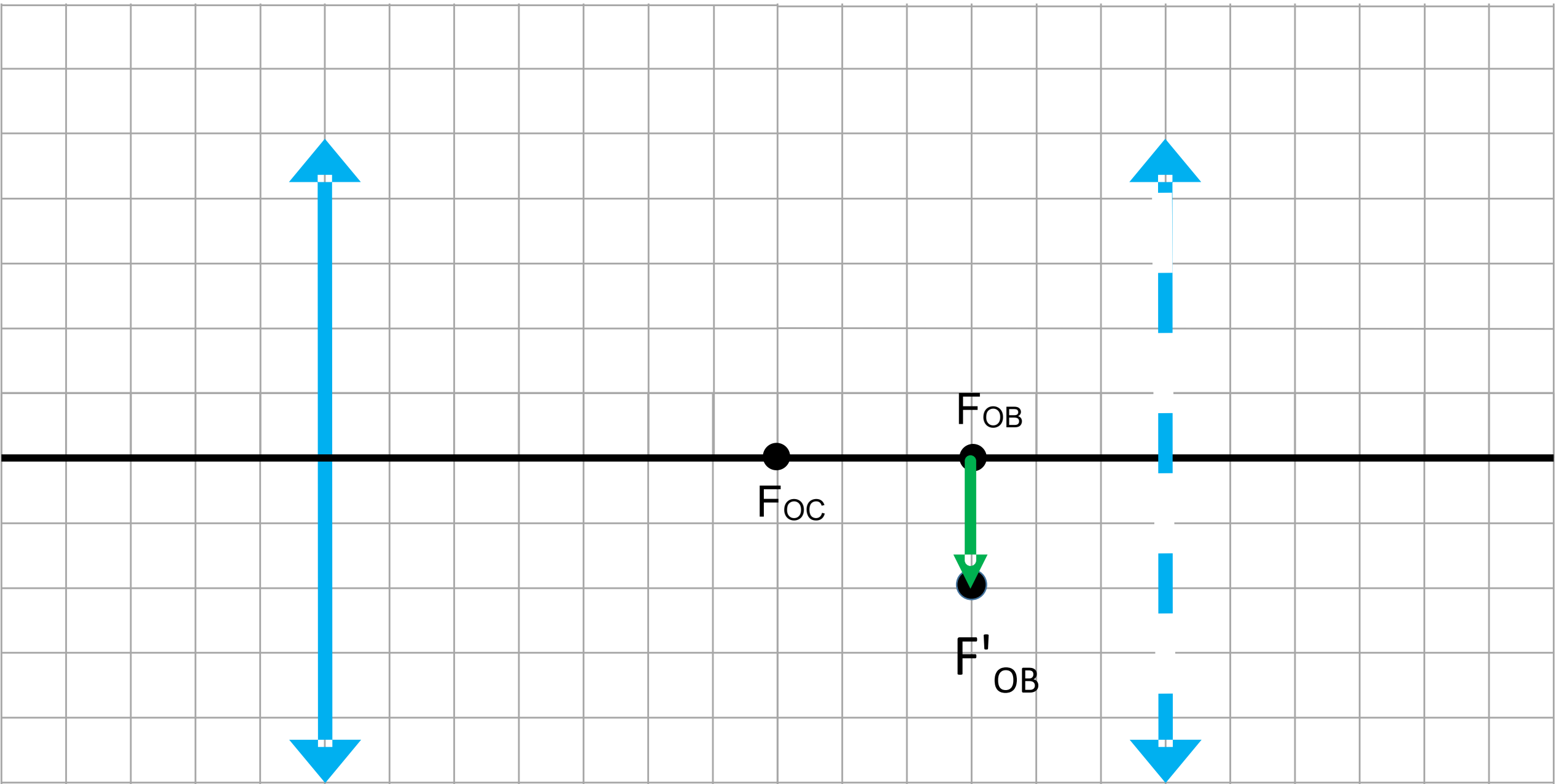


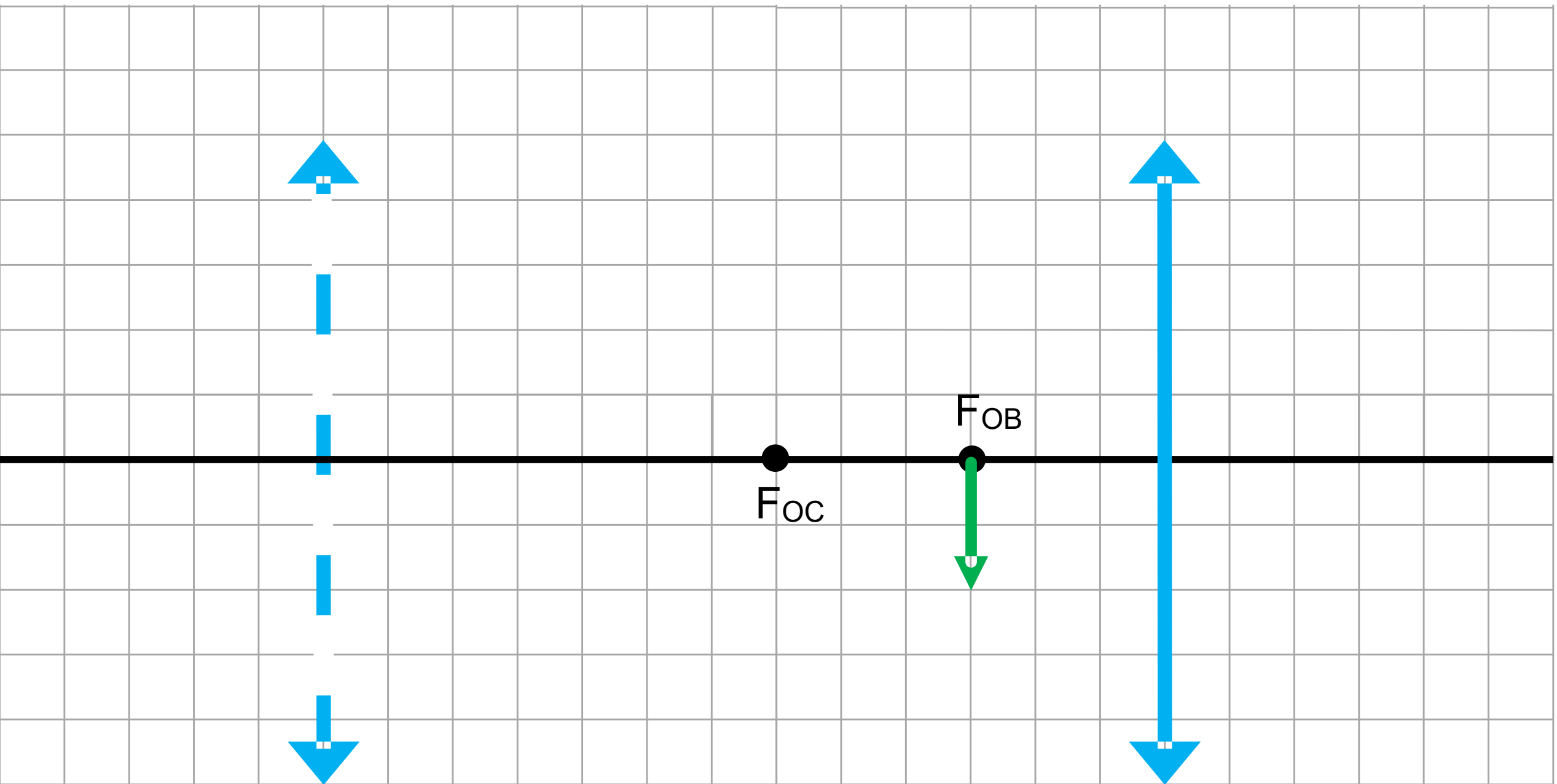


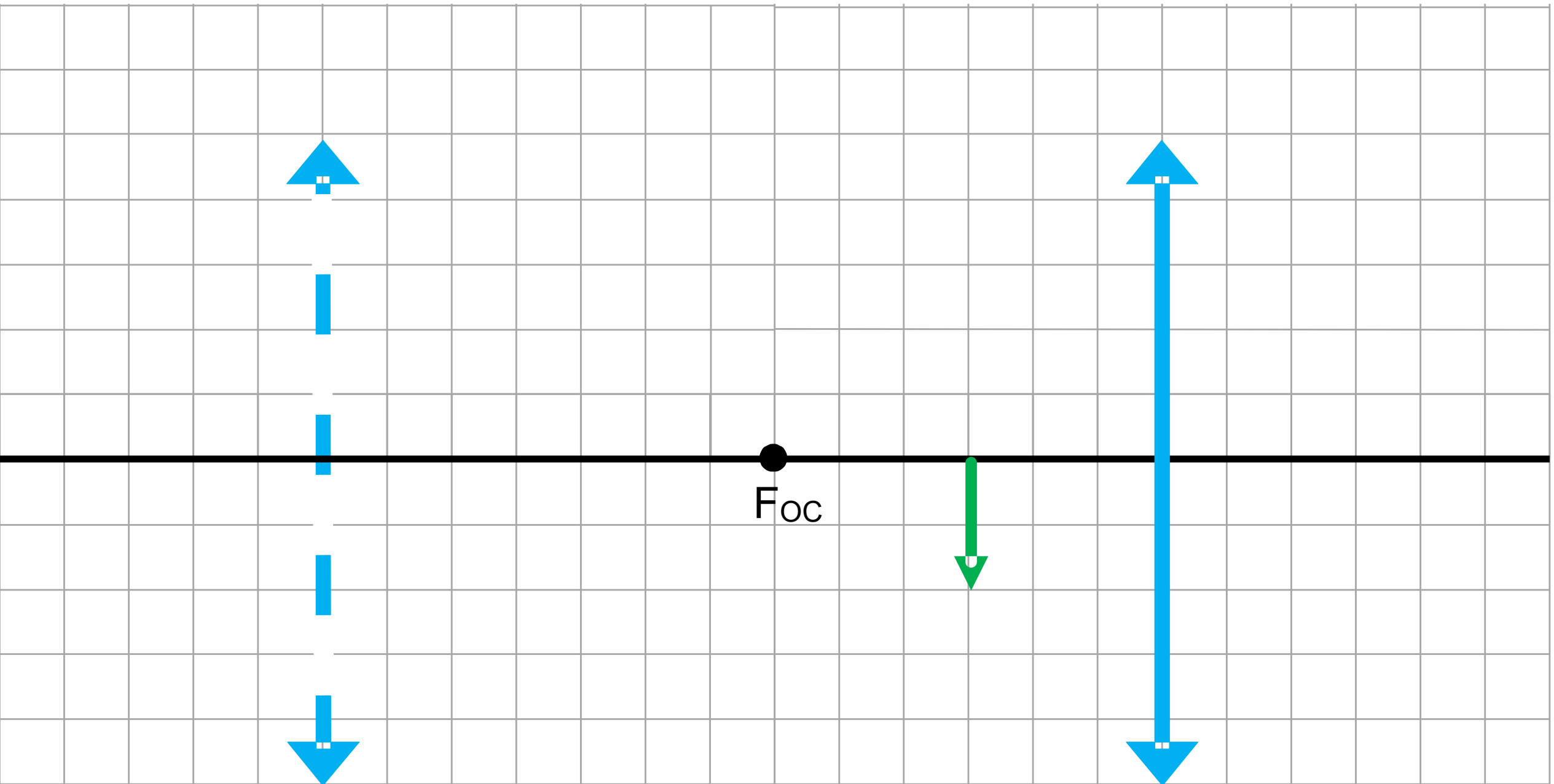


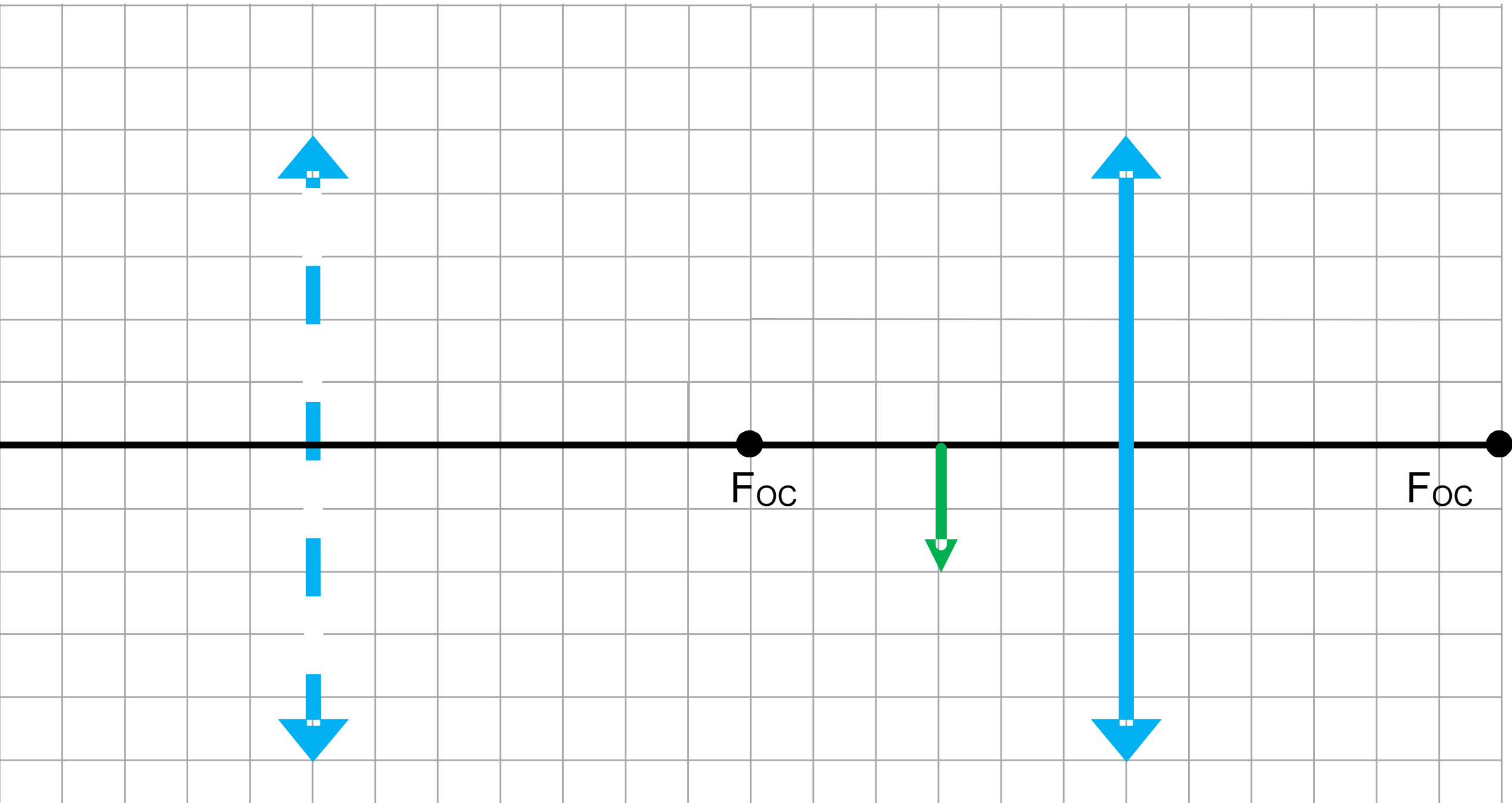


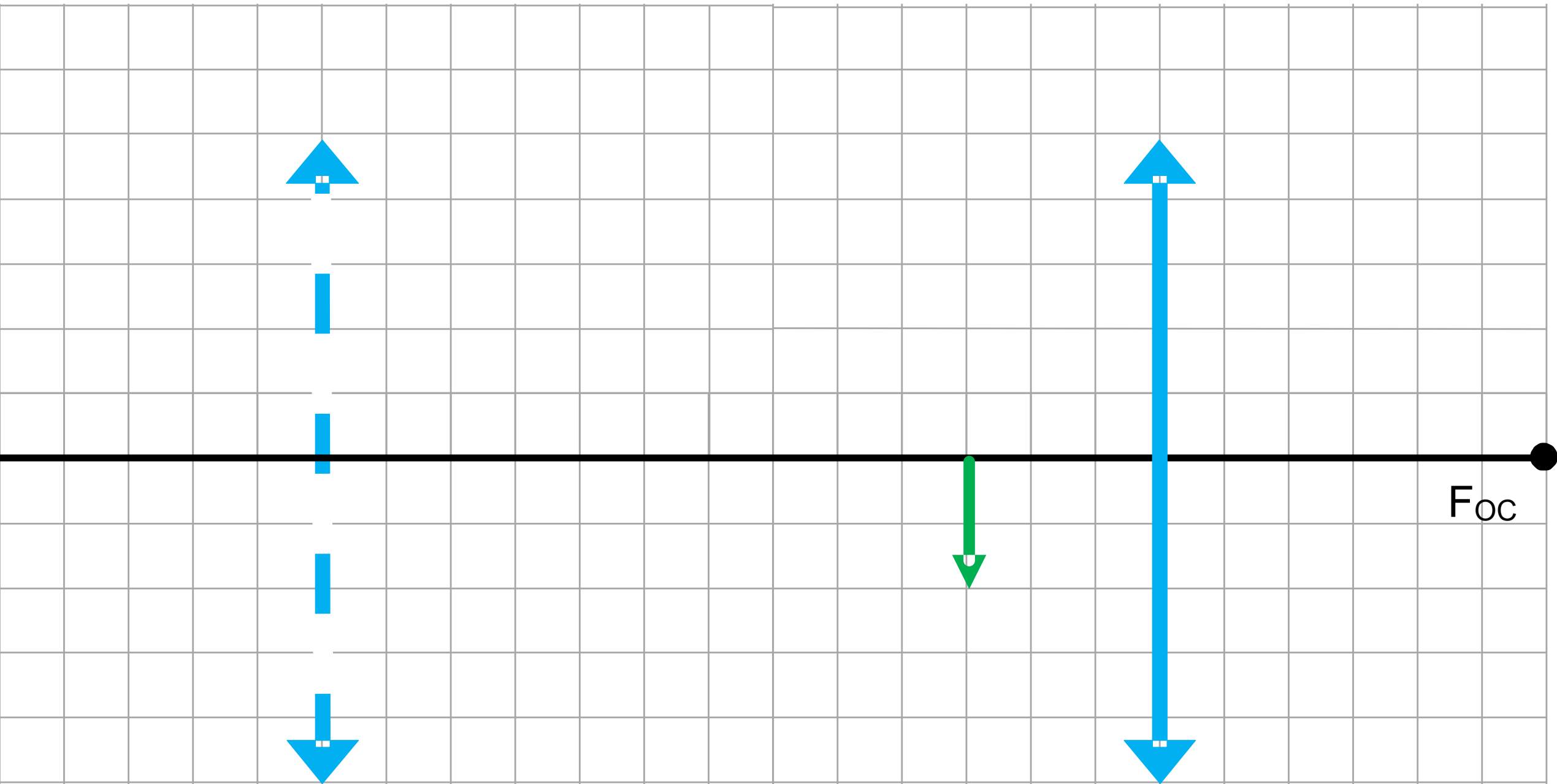


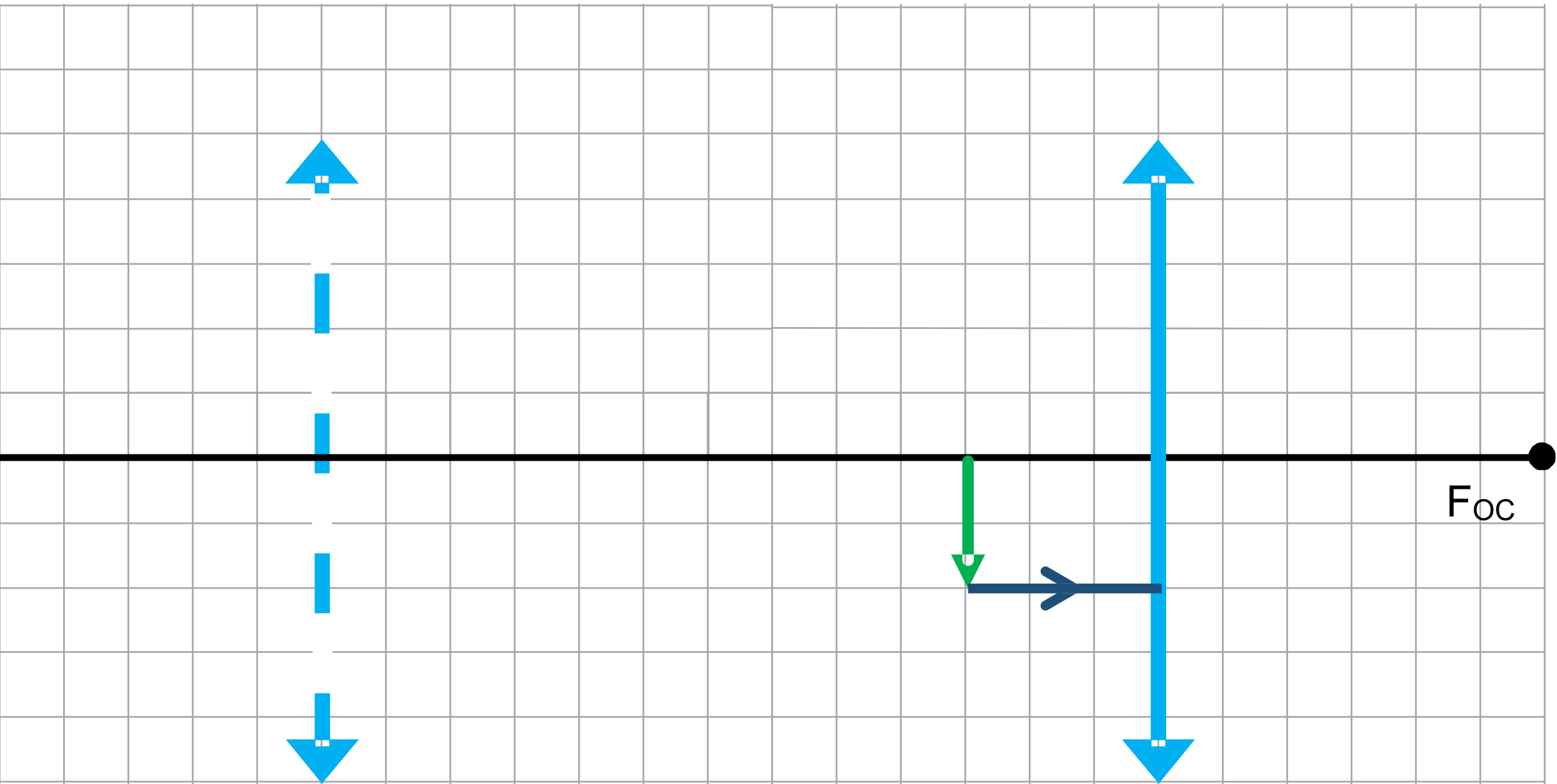


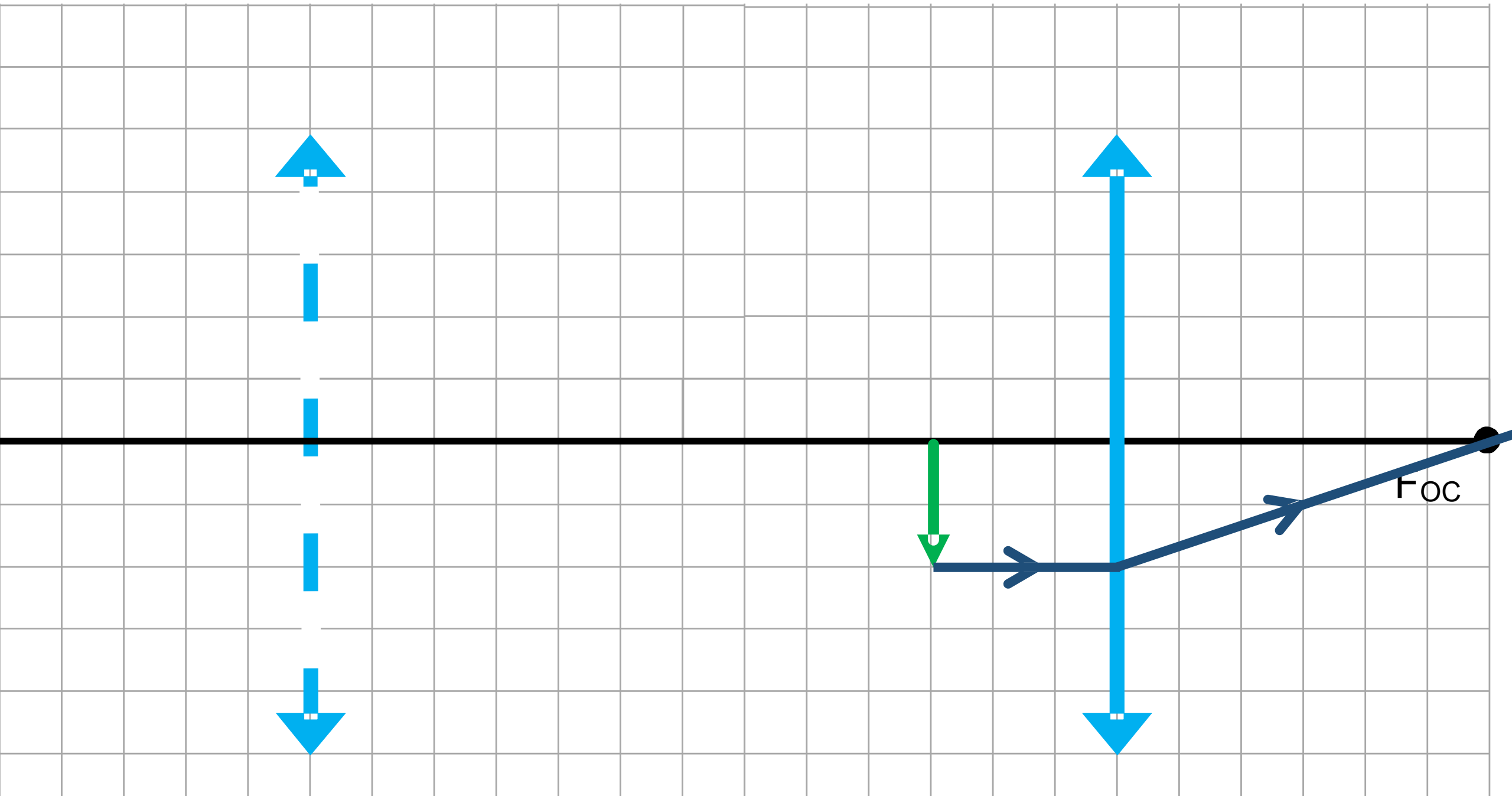


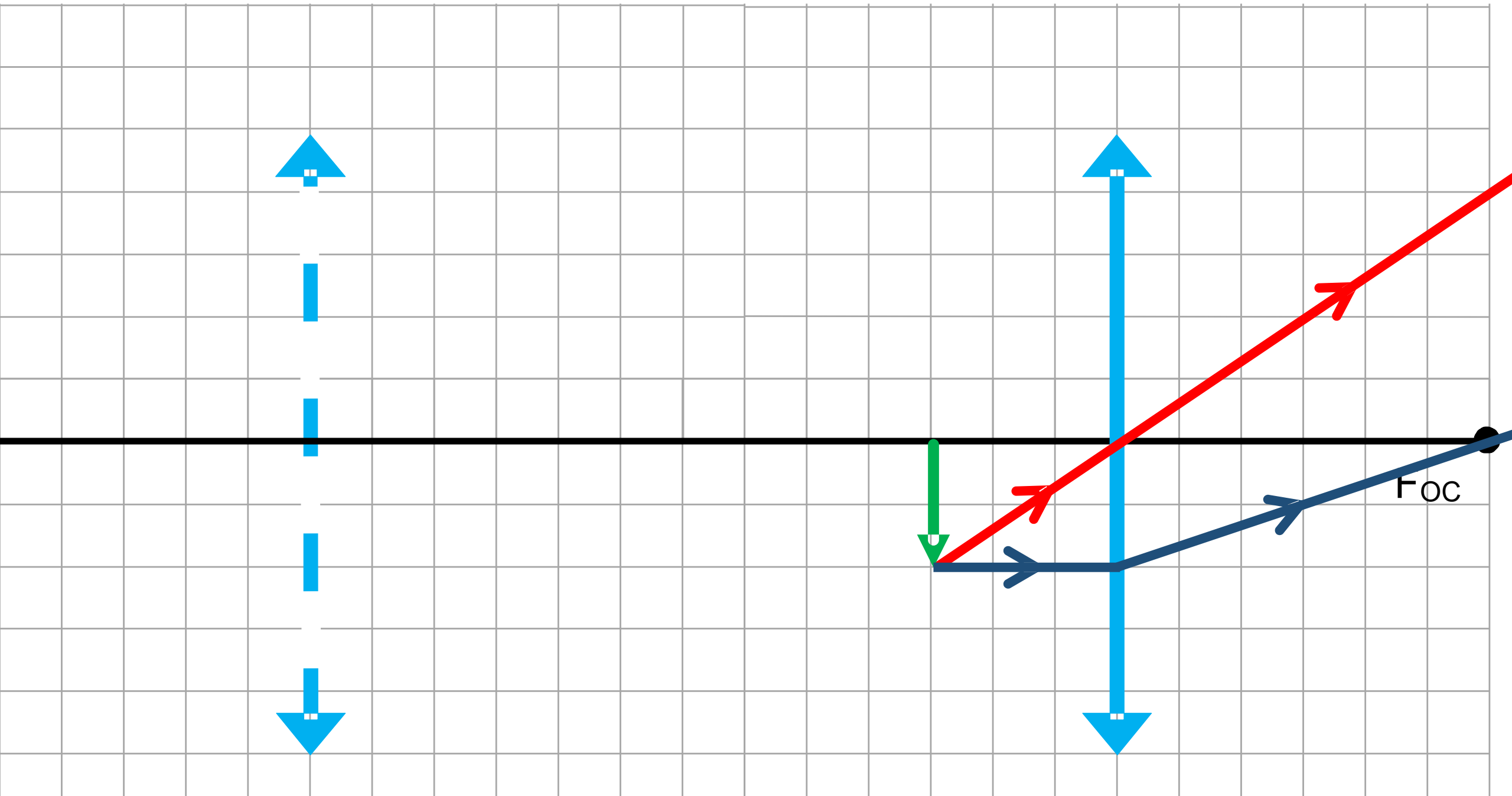


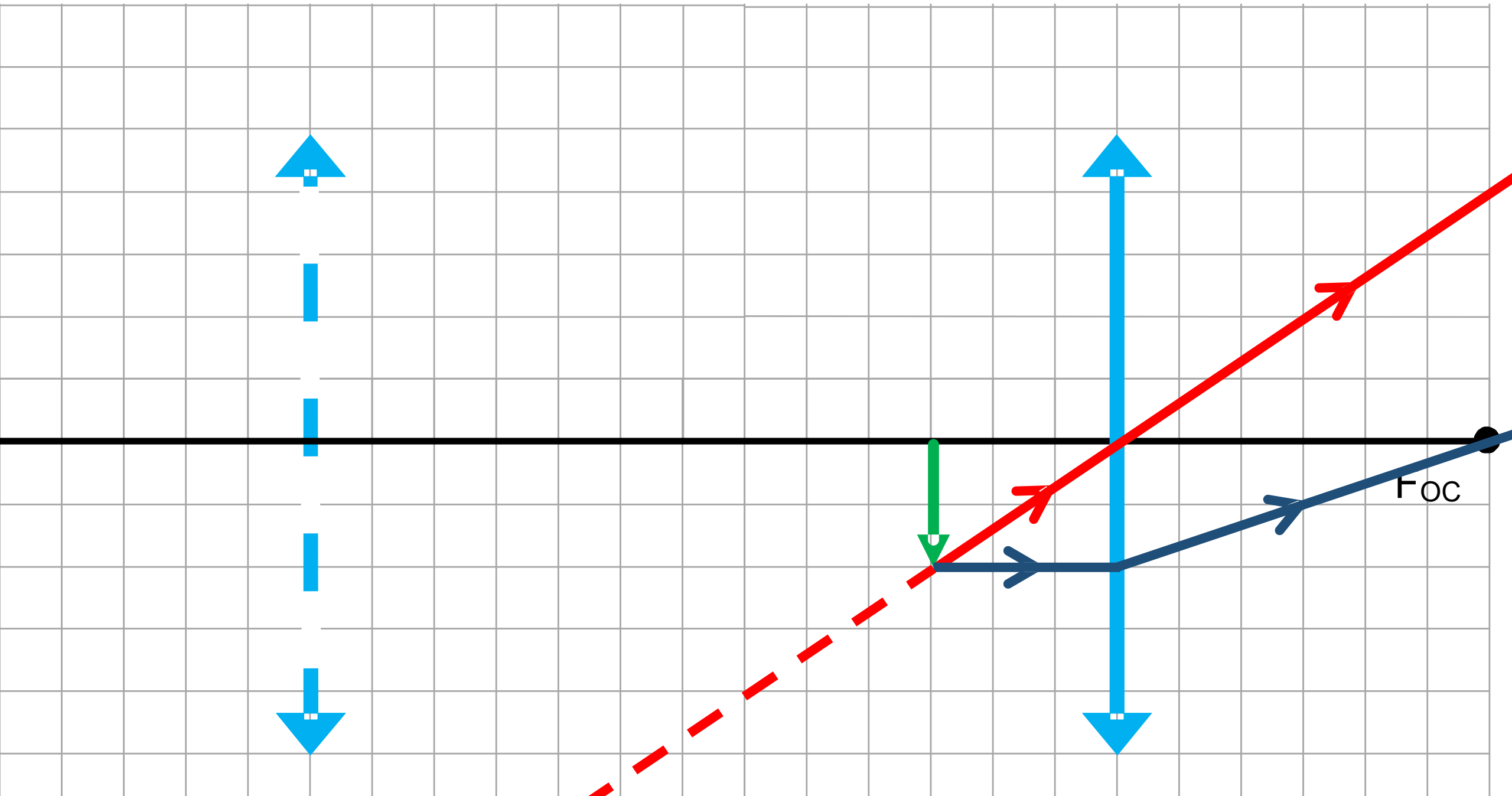


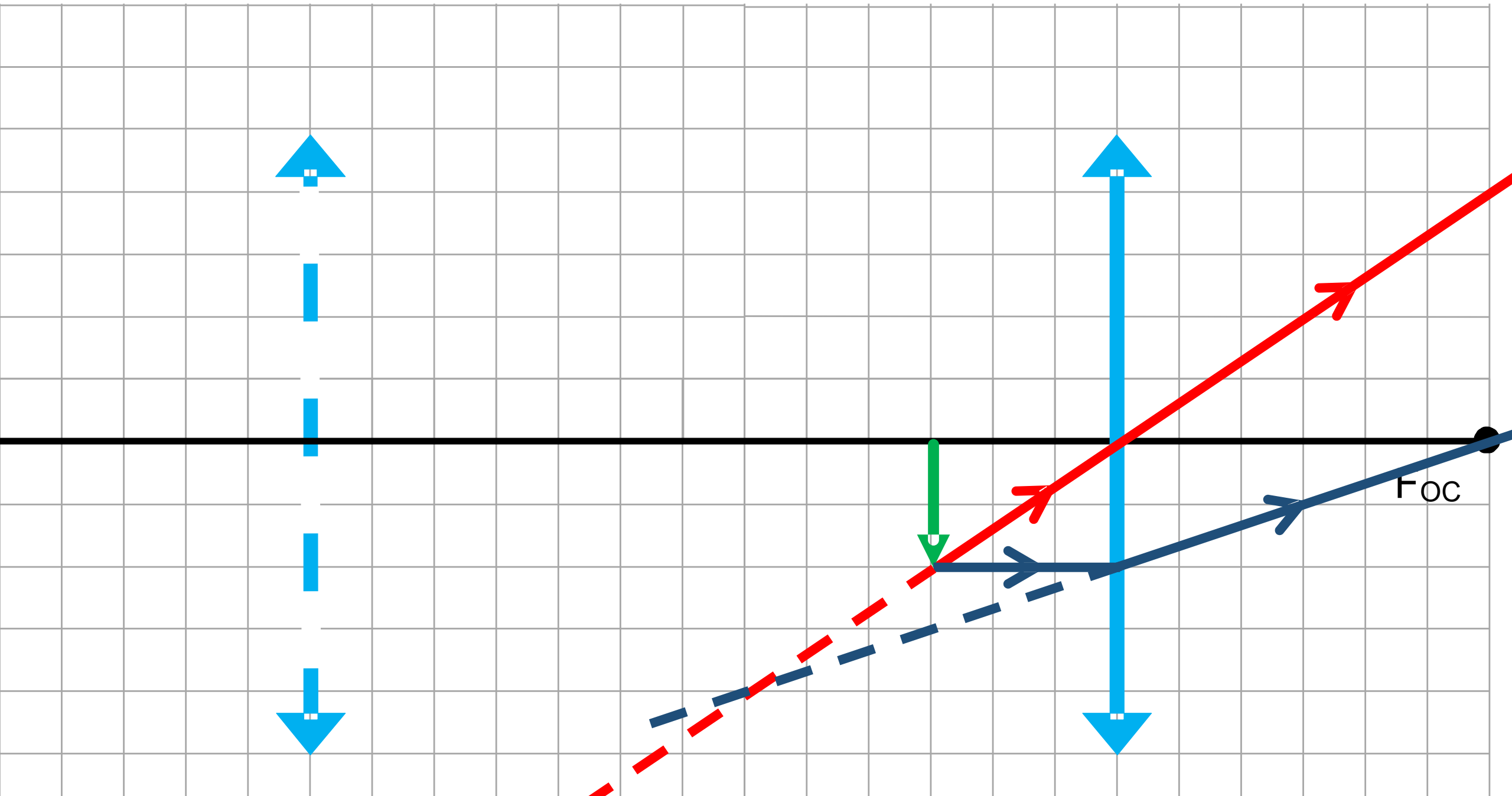


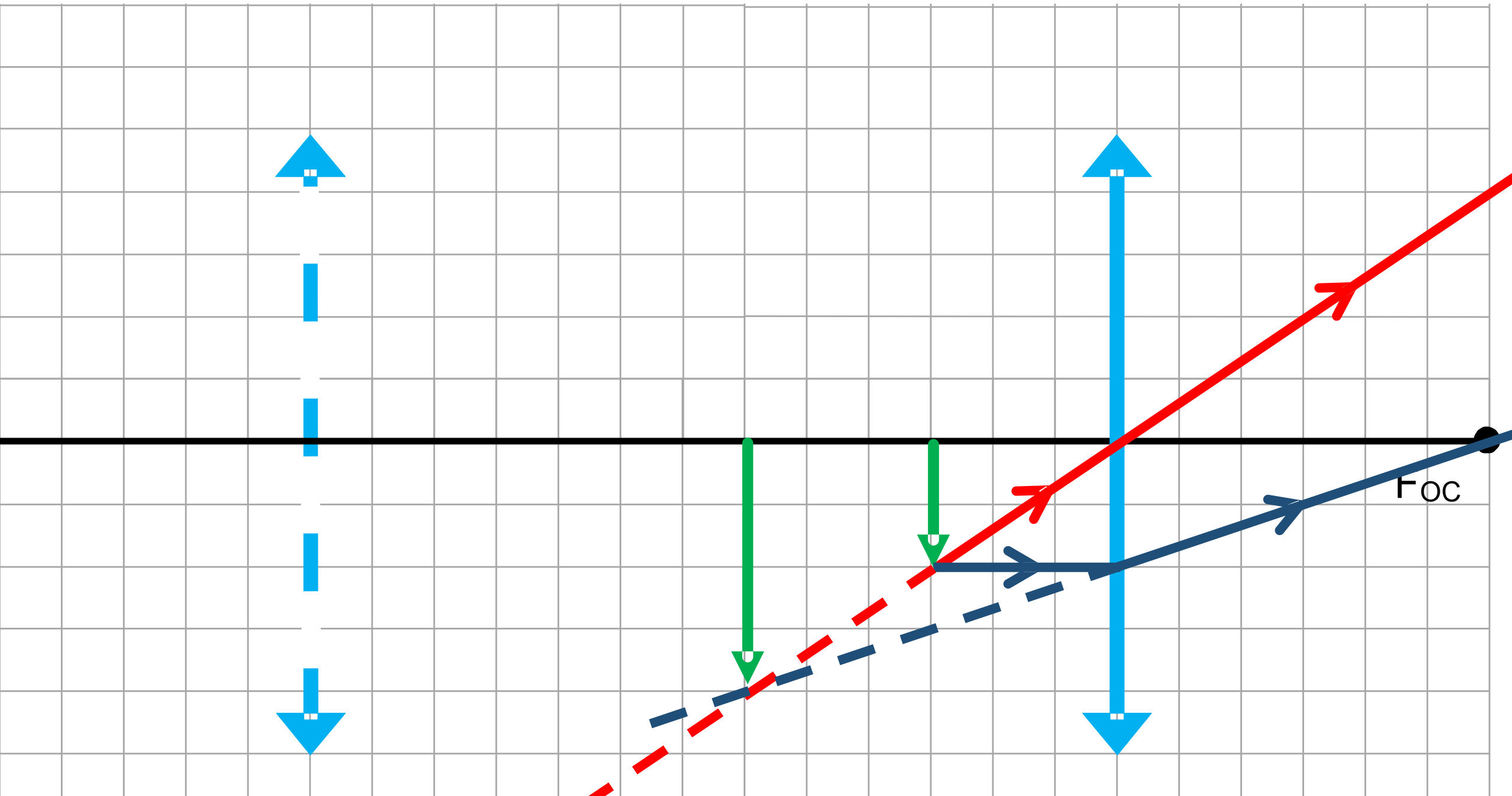


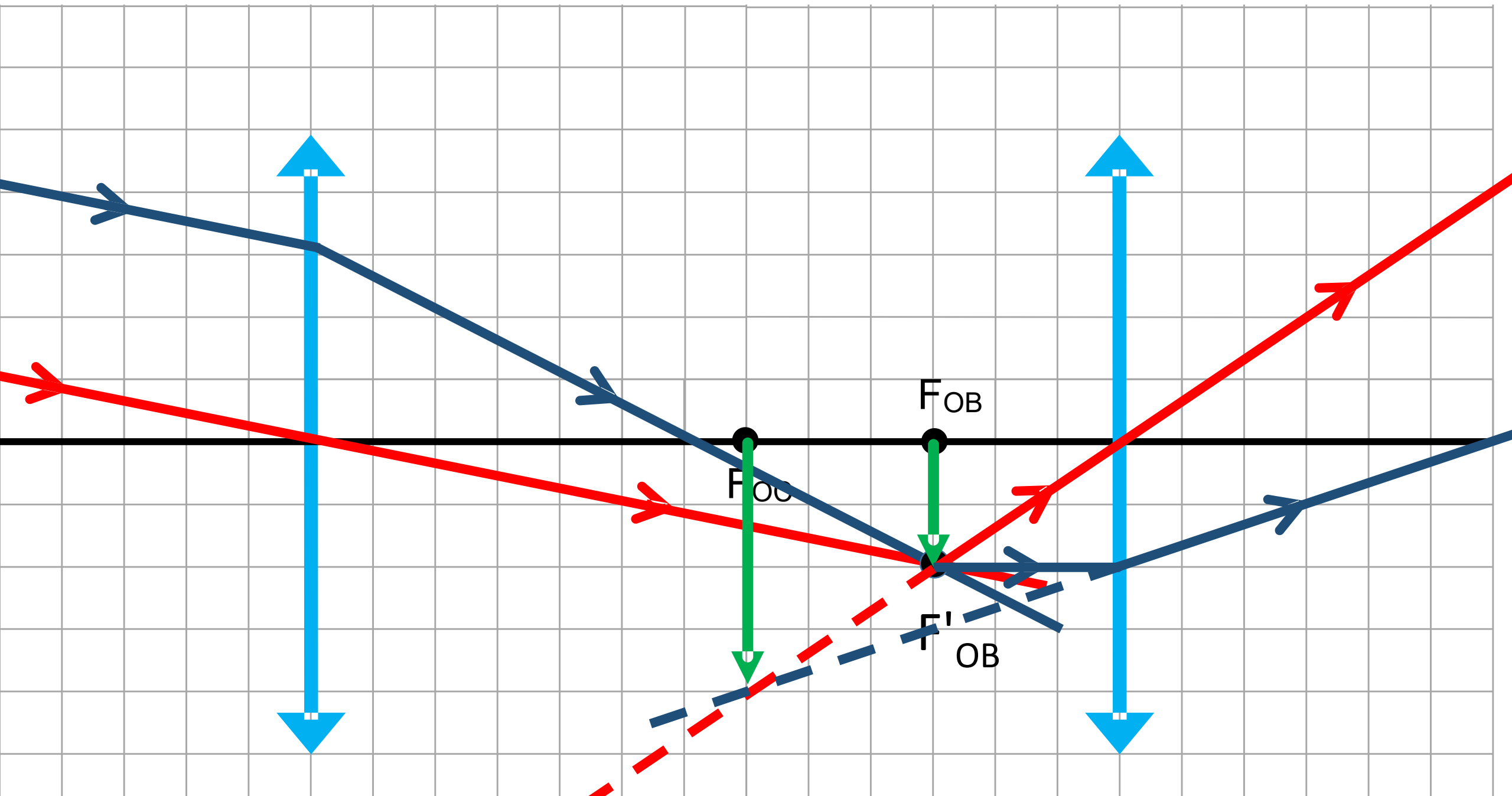












LUNETAS ASTRONÔMICAS

LUNETAS ASTRONÔMICAS

- IMAGEM:

LUNETAS ASTRONÔMICAS

- IMAGEM:
 - VIRTUAL

LUNETAS ASTRONÔMICAS

- IMAGEM:
 - VIRTUAL
 - INVERTIDA

LUNETAS ASTRONÔMICAS

- IMAGEM:

- VIRTUAL
- INVERTIDA
- QUANTO AO TAMANHO, DIZEMOS QUE HÁ UM AUMENTO ANGULAR (M), UMA VEZ QUE A IMAGEM FINAL É PONTUAL

Ocular

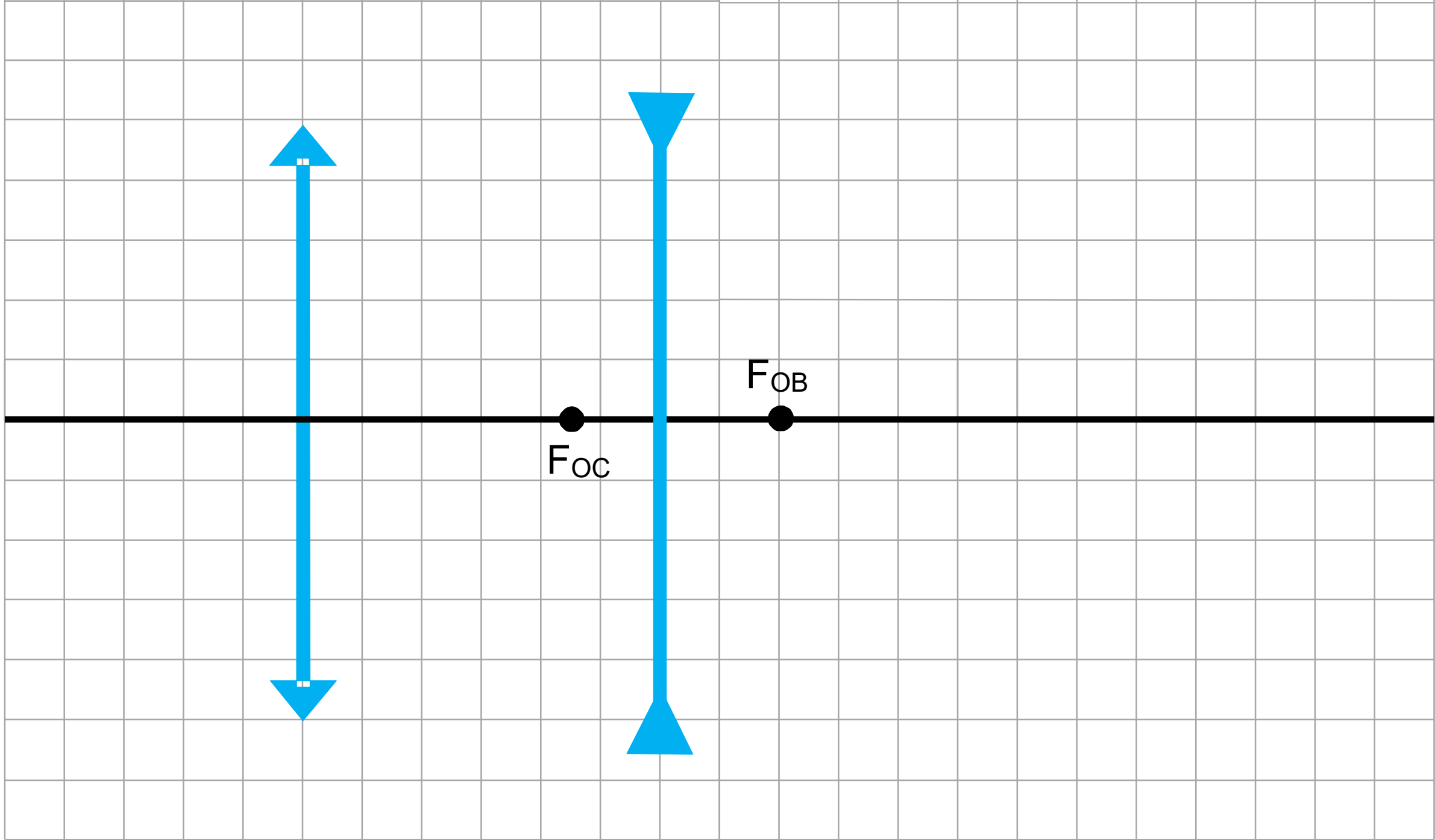


LUNETAS ASTRONÔMICA → TELESCÓPIO

- Os telescópios refratores são como lunetas, mas com um pequeno espelho plano próximo à ocular para ficar mais confortável a observação.



Q. 15 - LUNETTA TERRESTRE



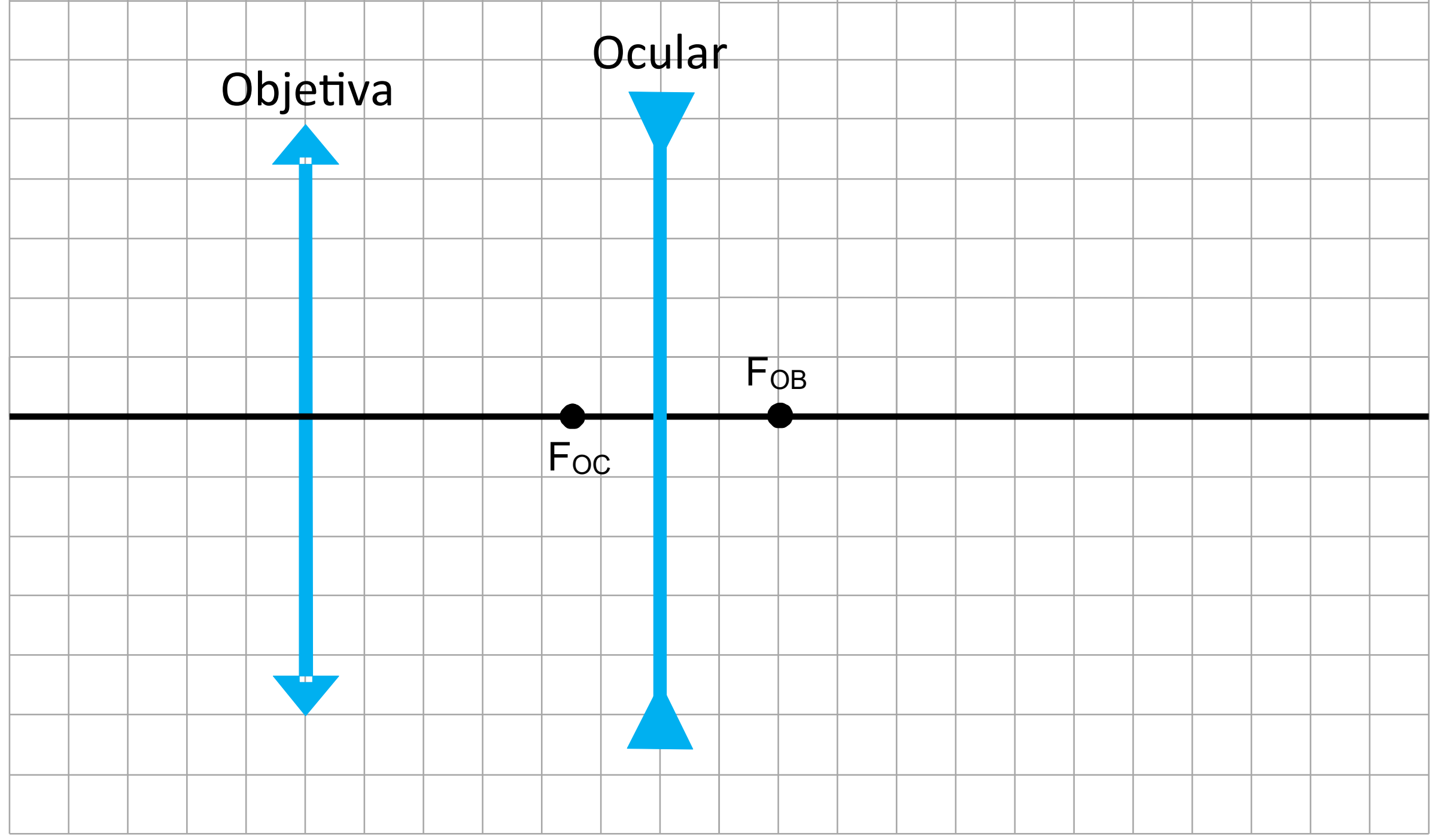
Objetiva

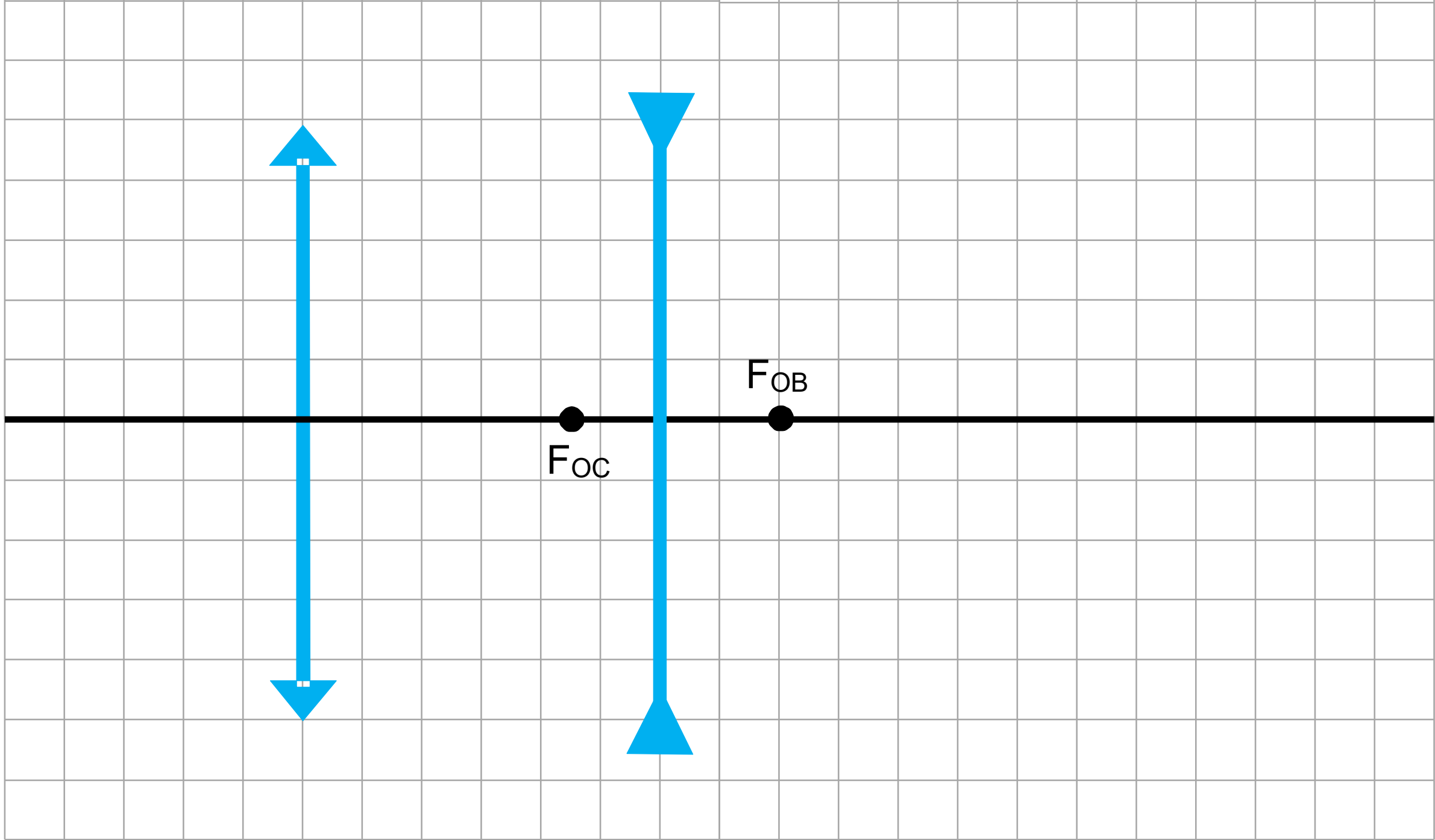
Ocular

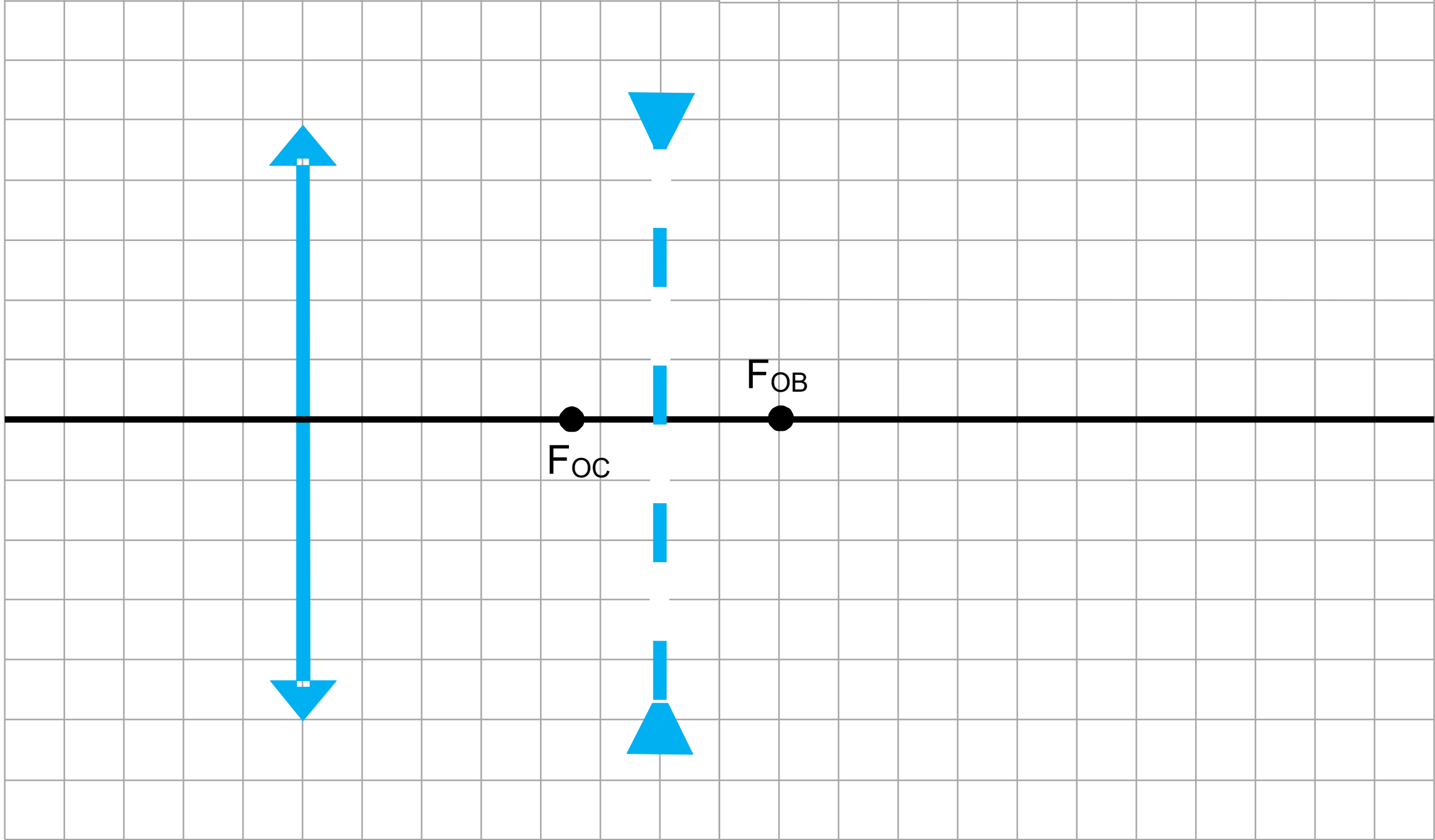


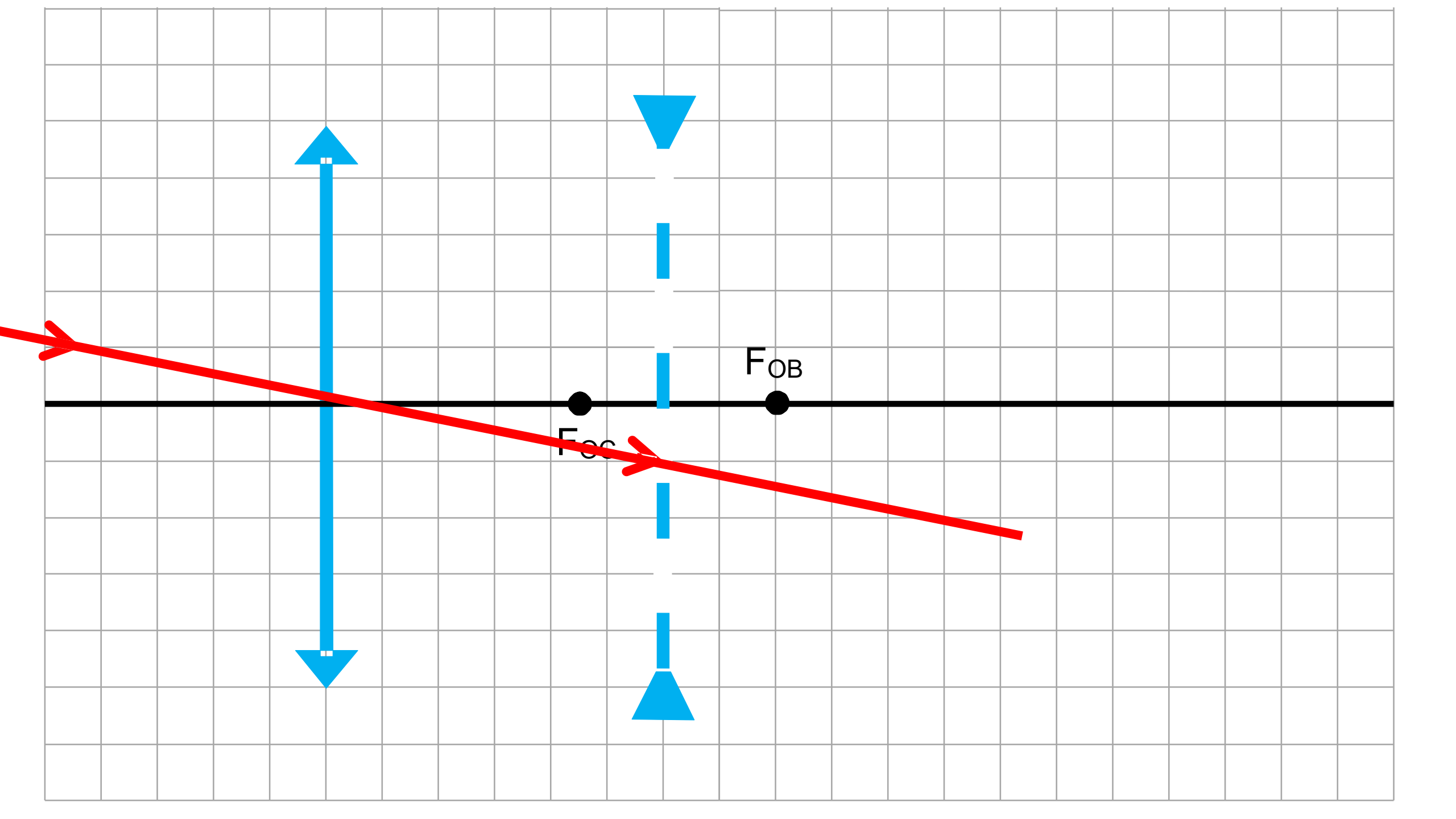
F_{OC}

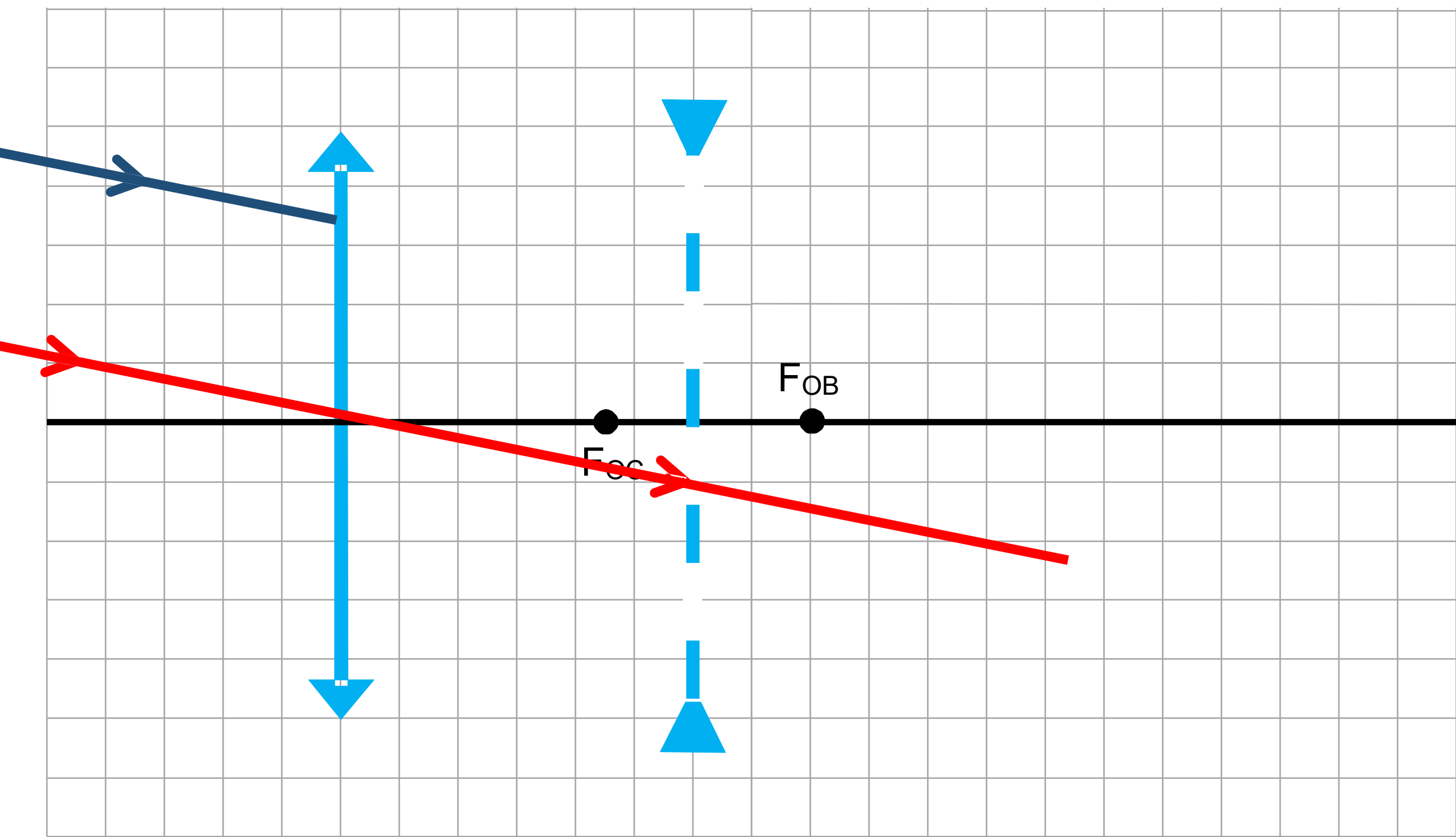
F_{OB}

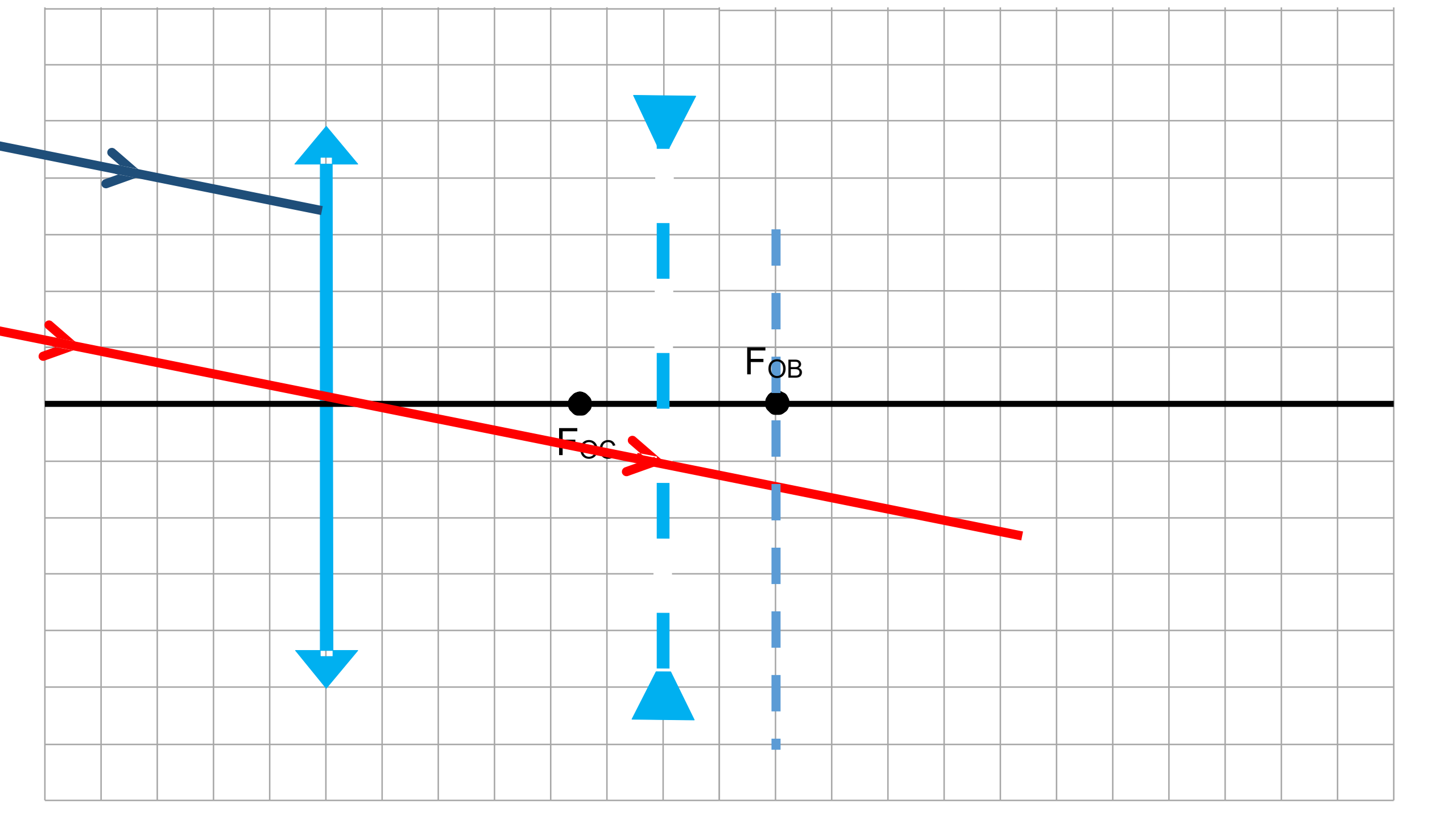


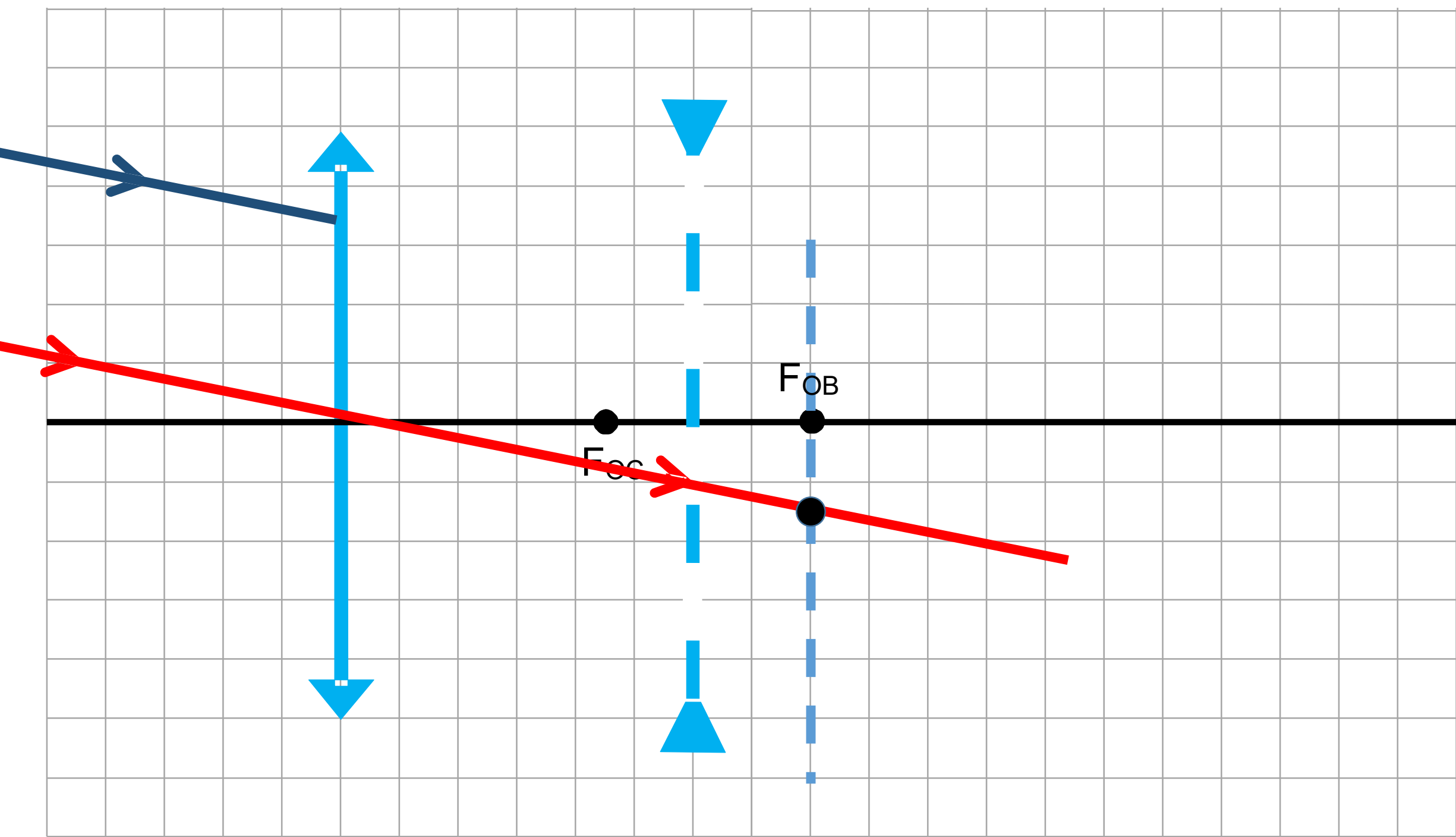


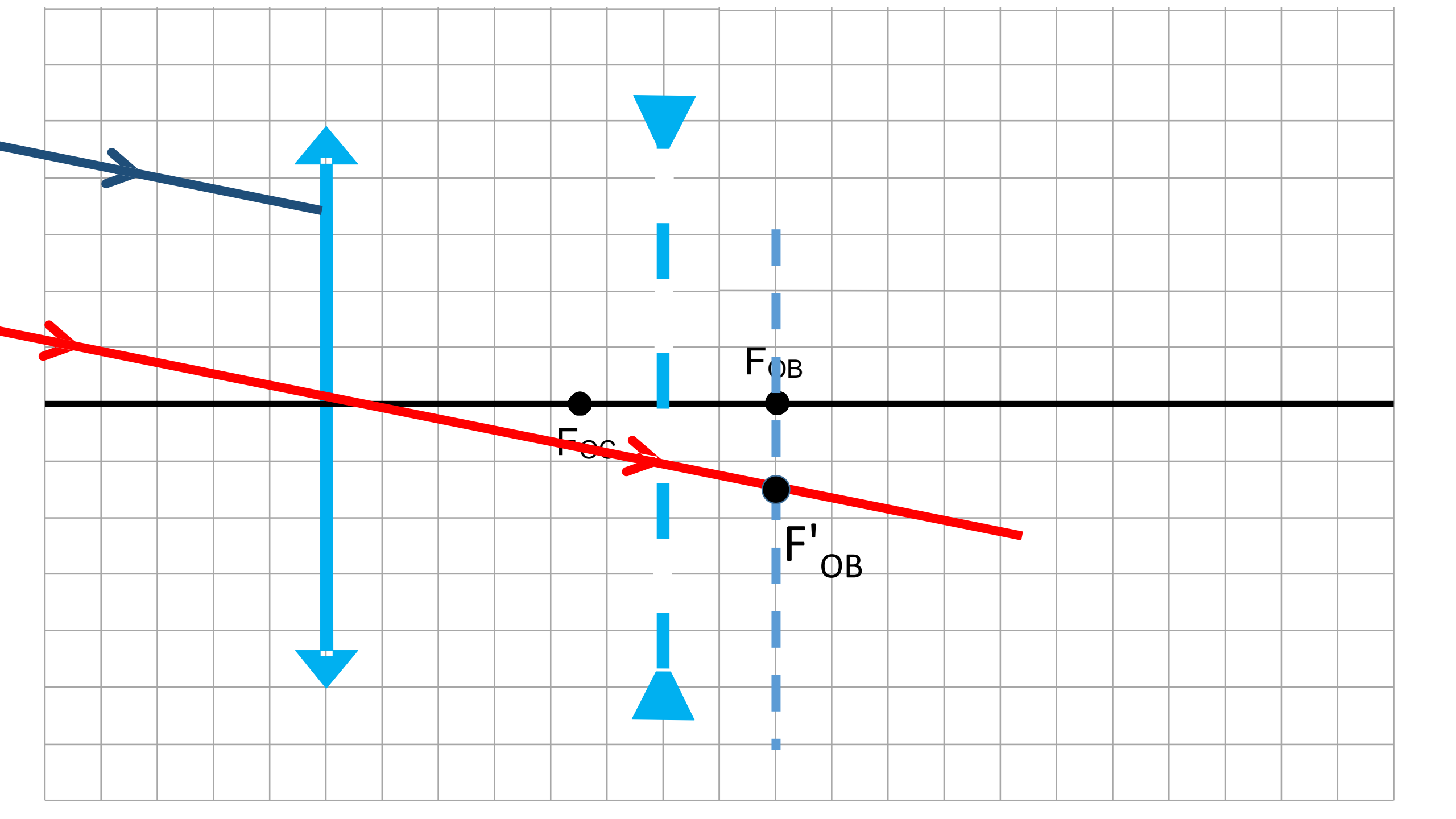


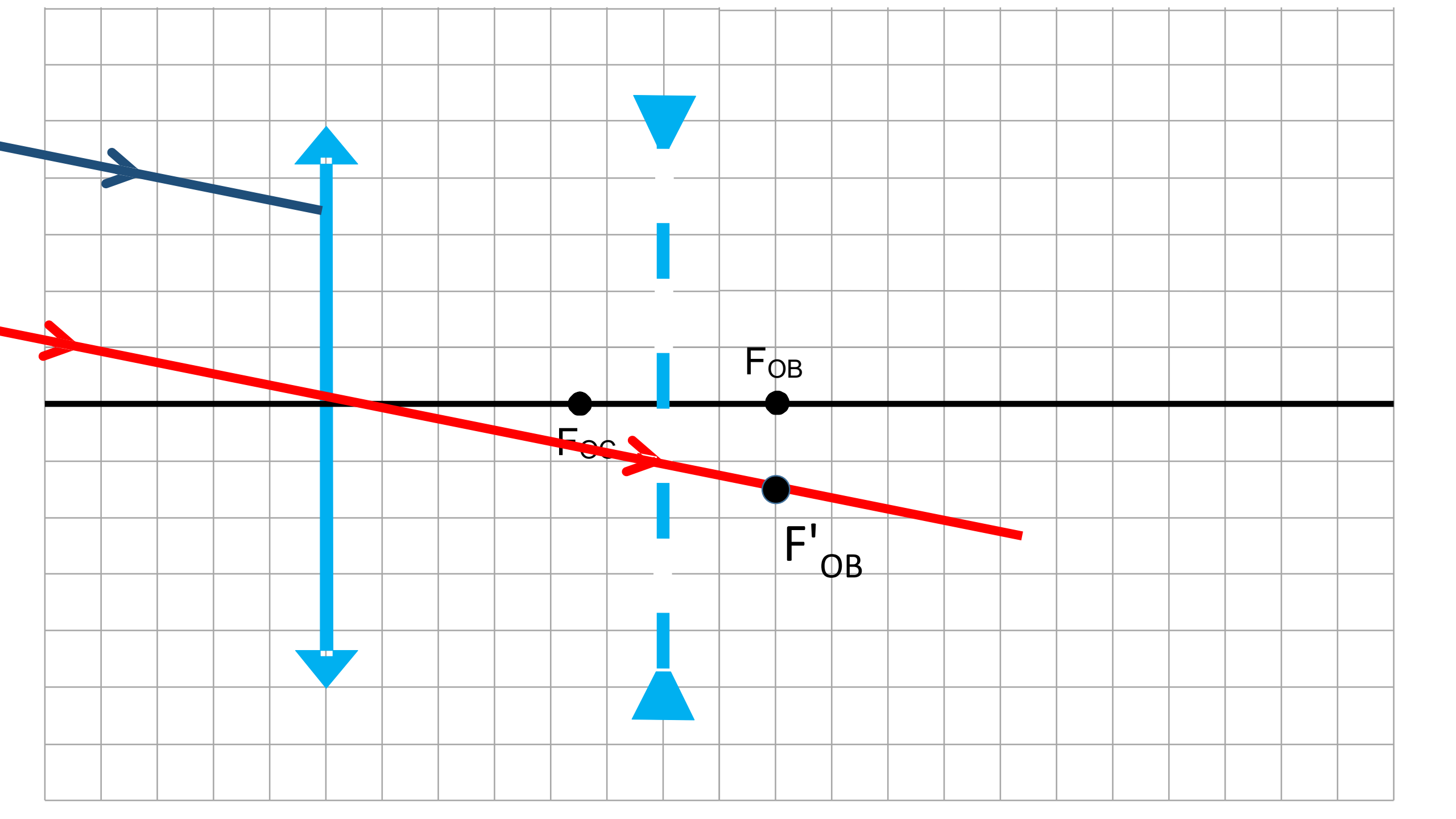


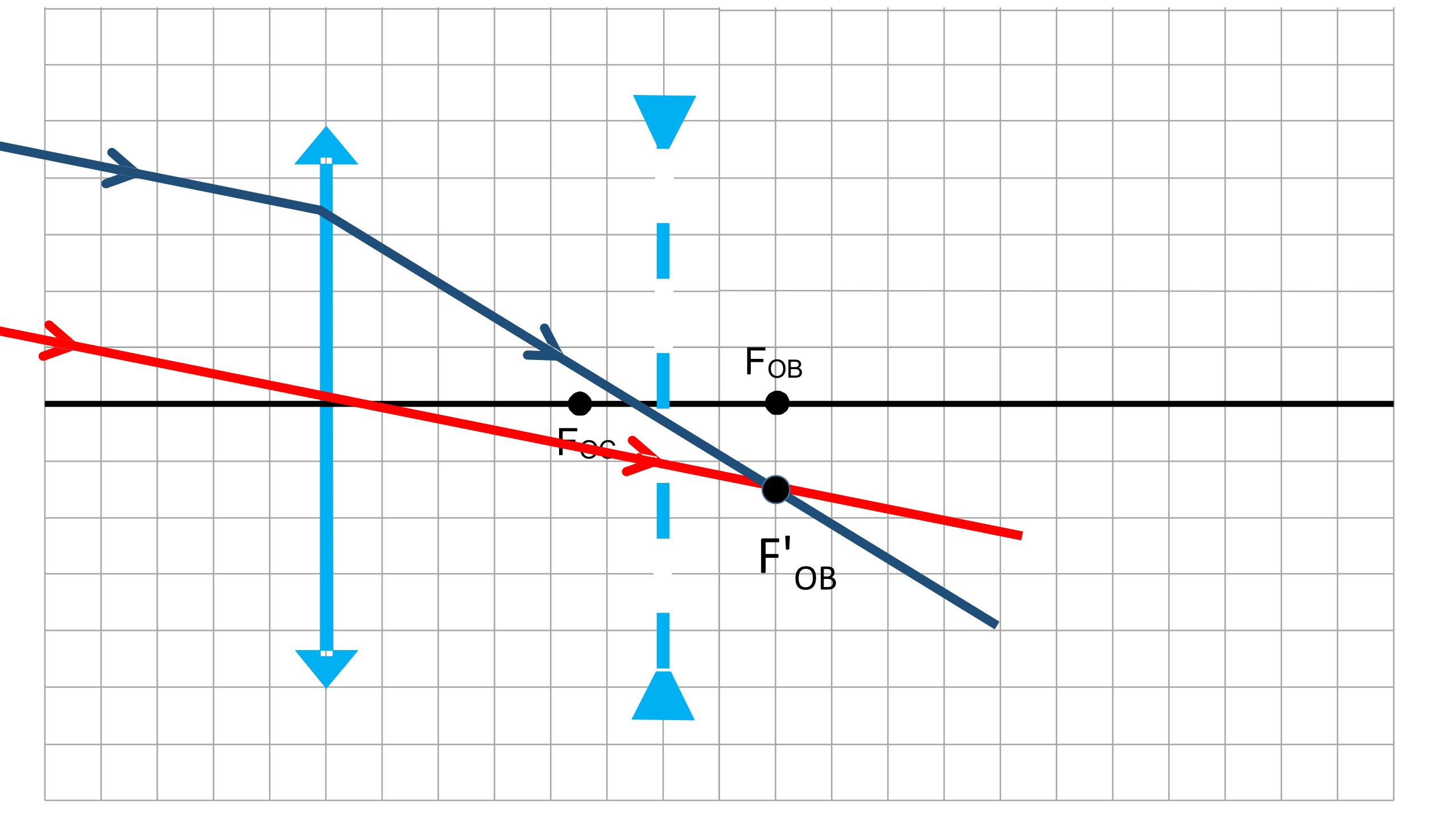


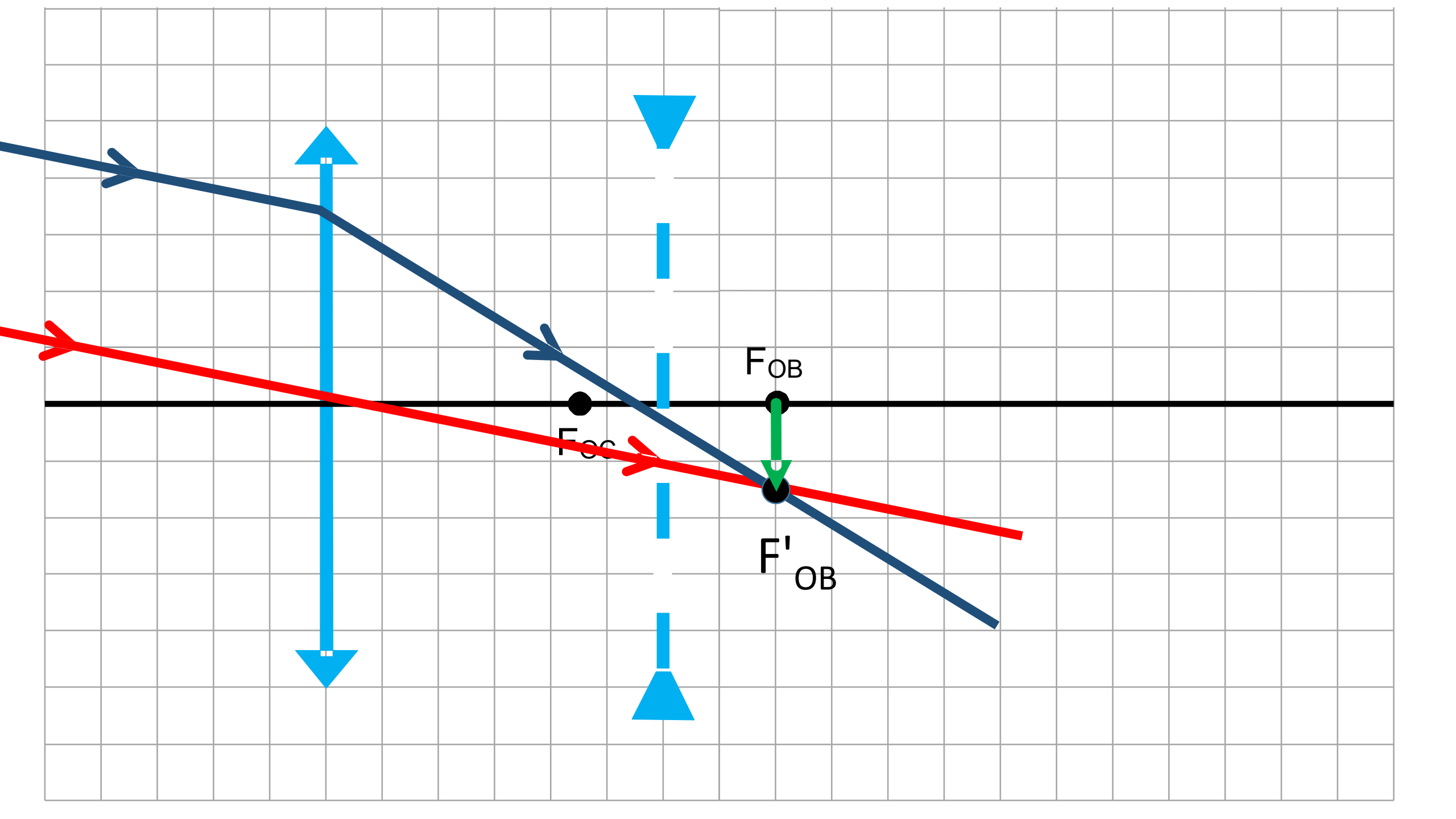


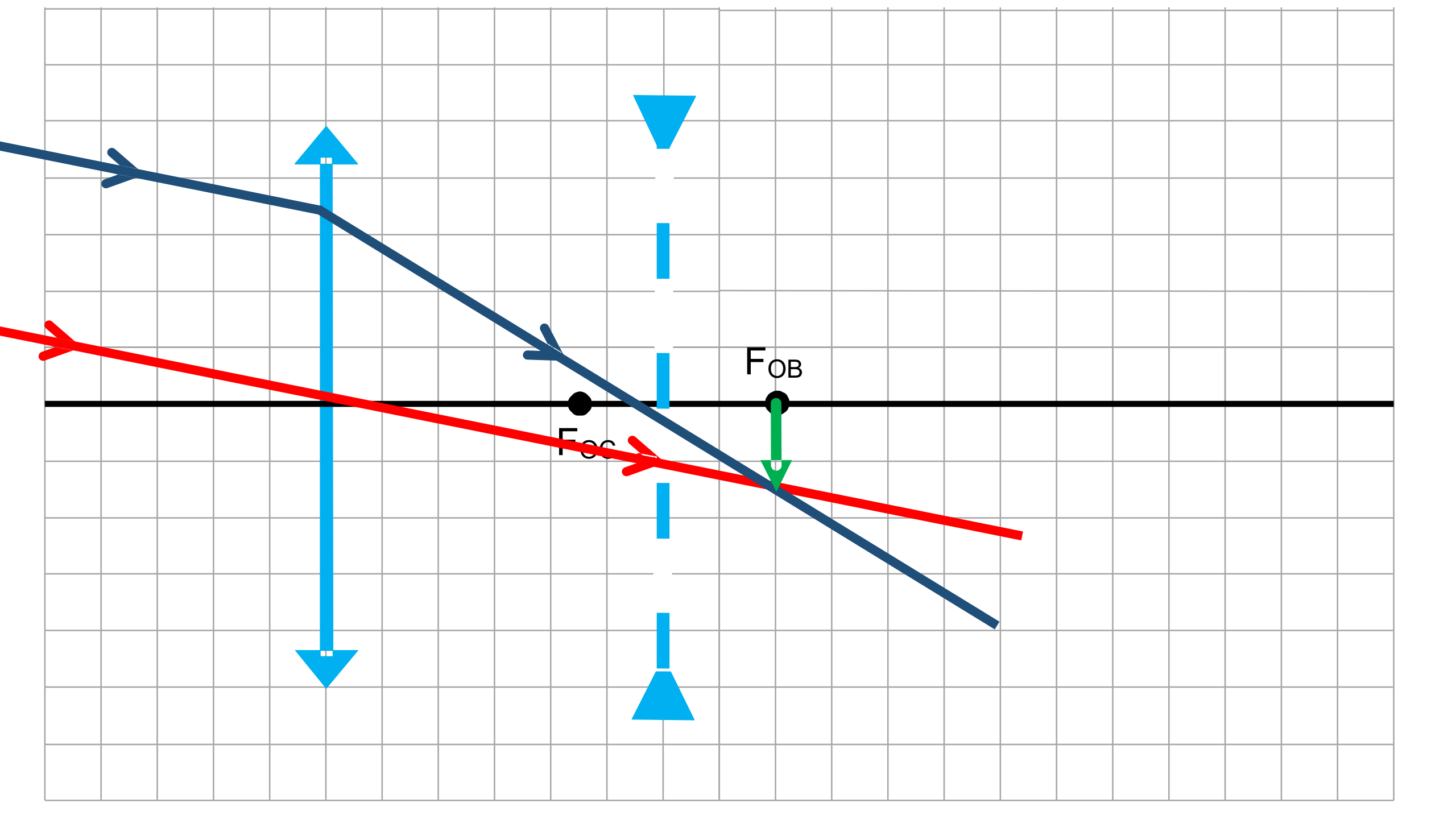


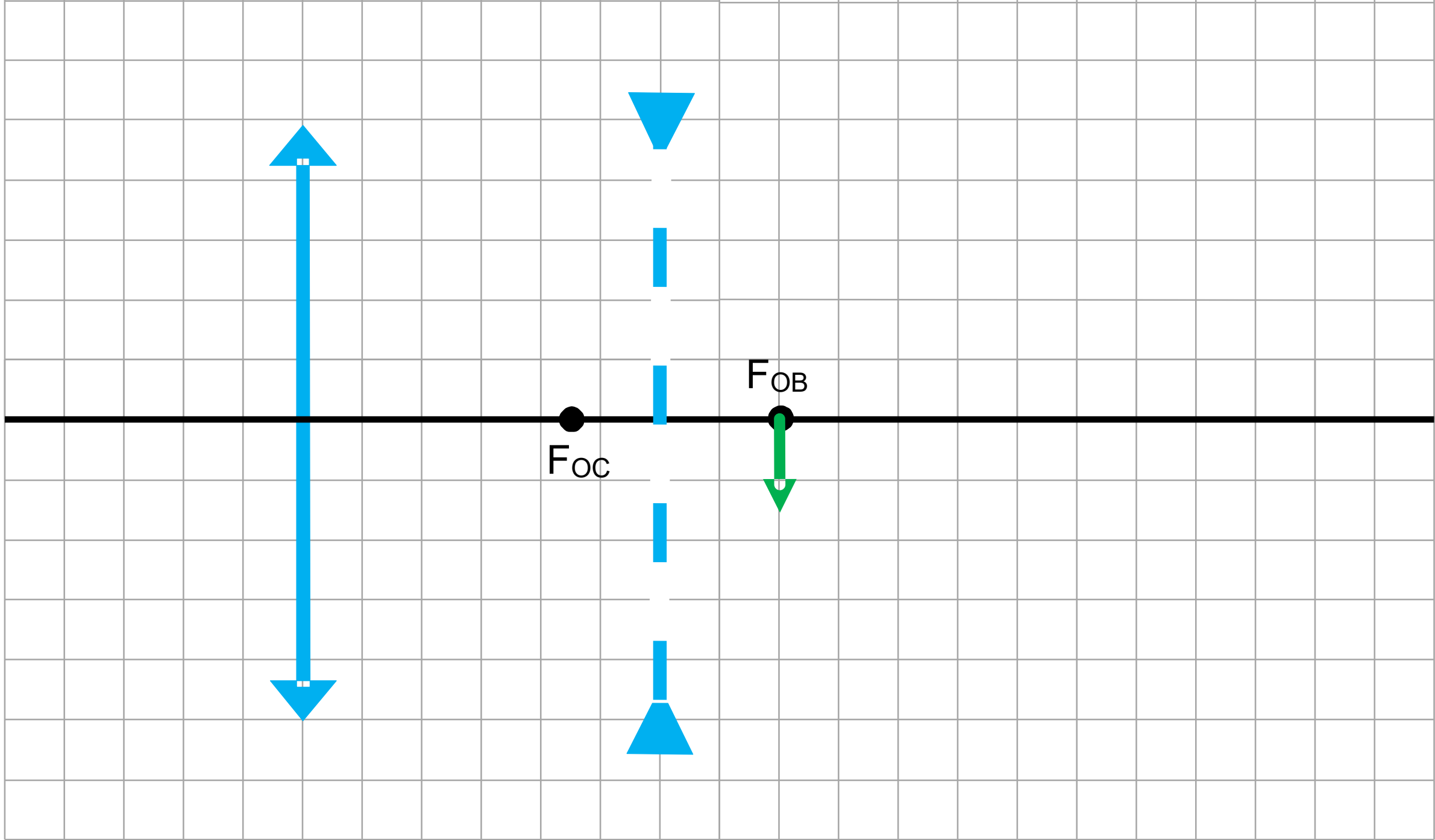


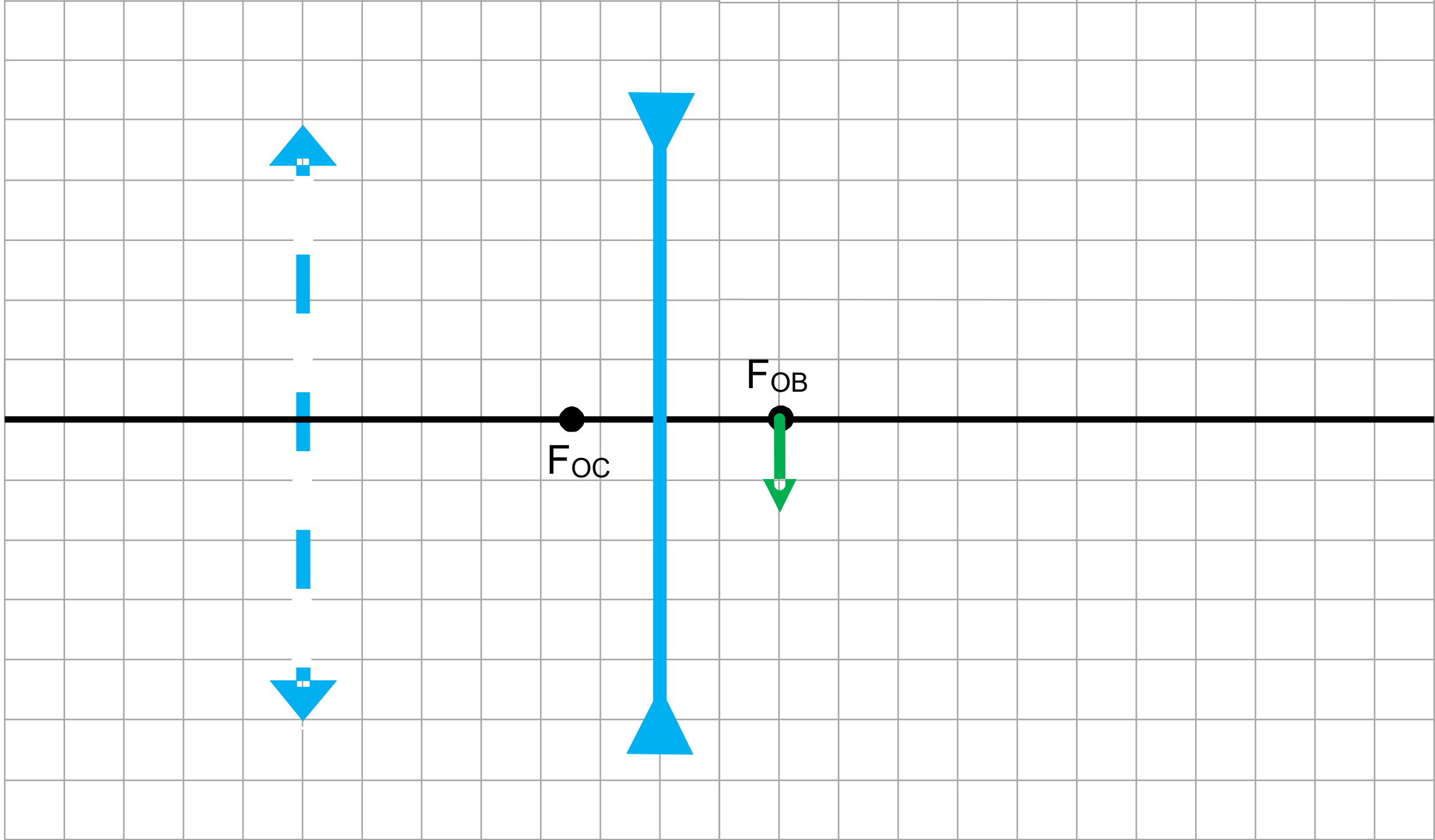


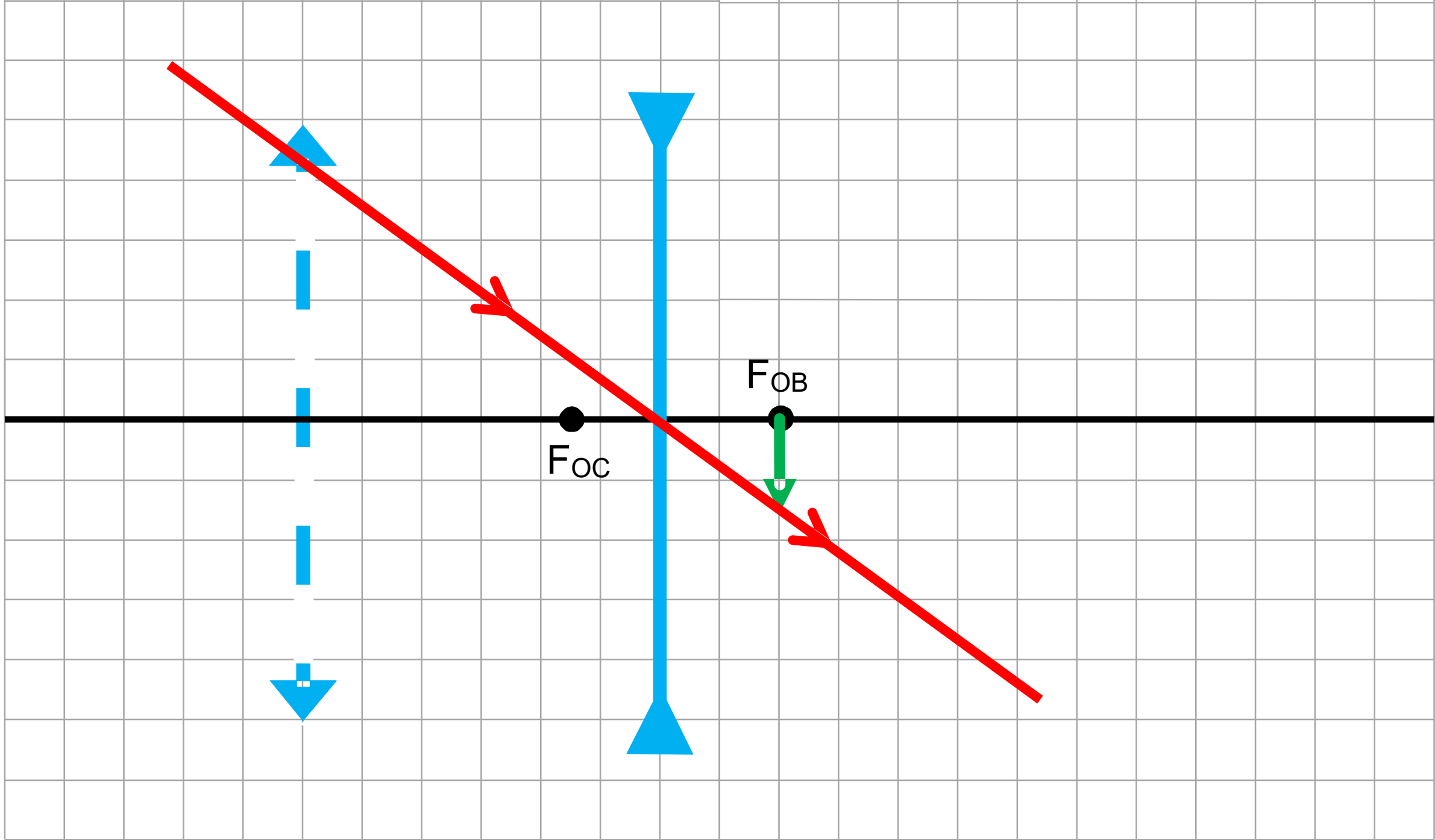


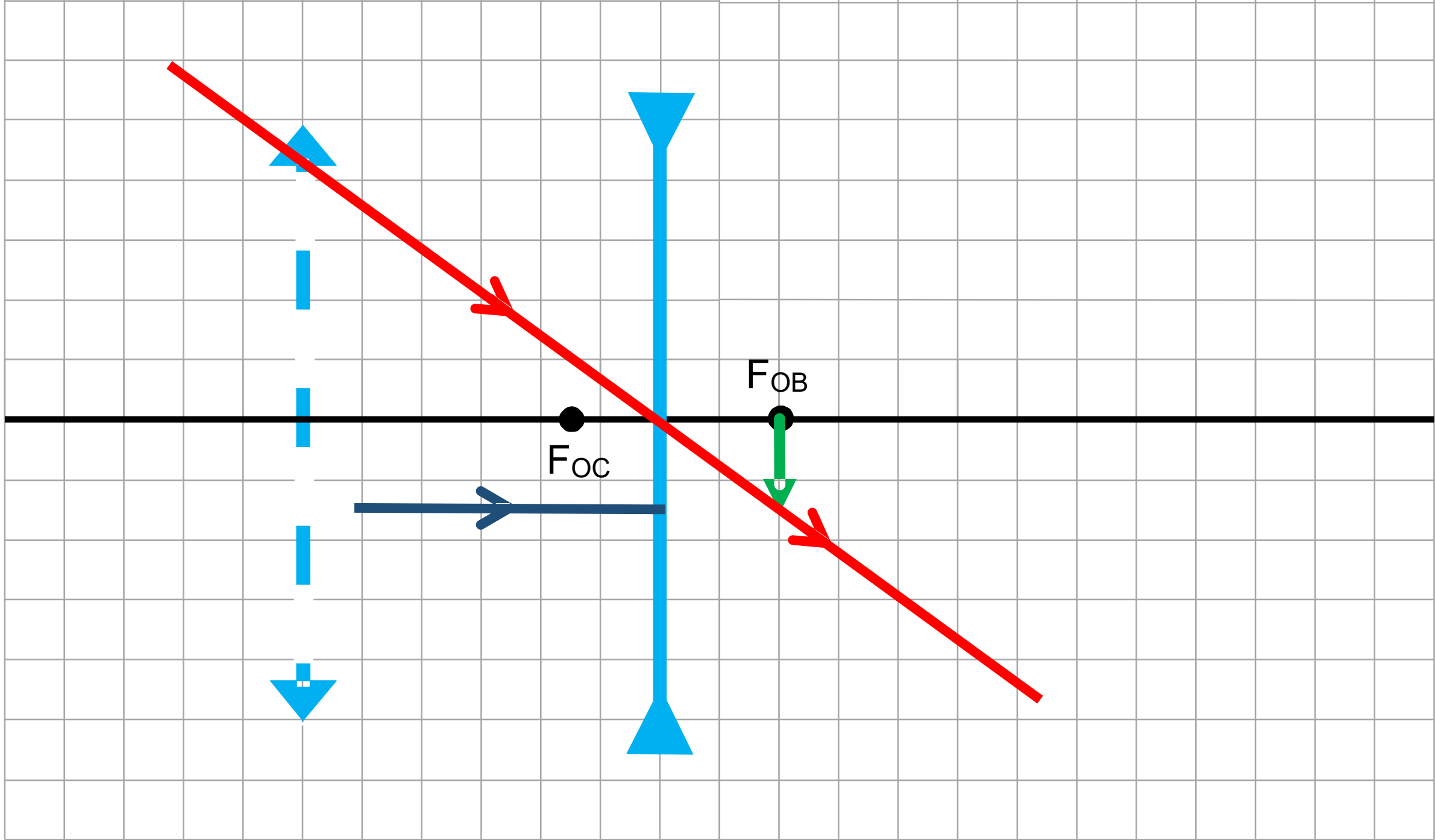


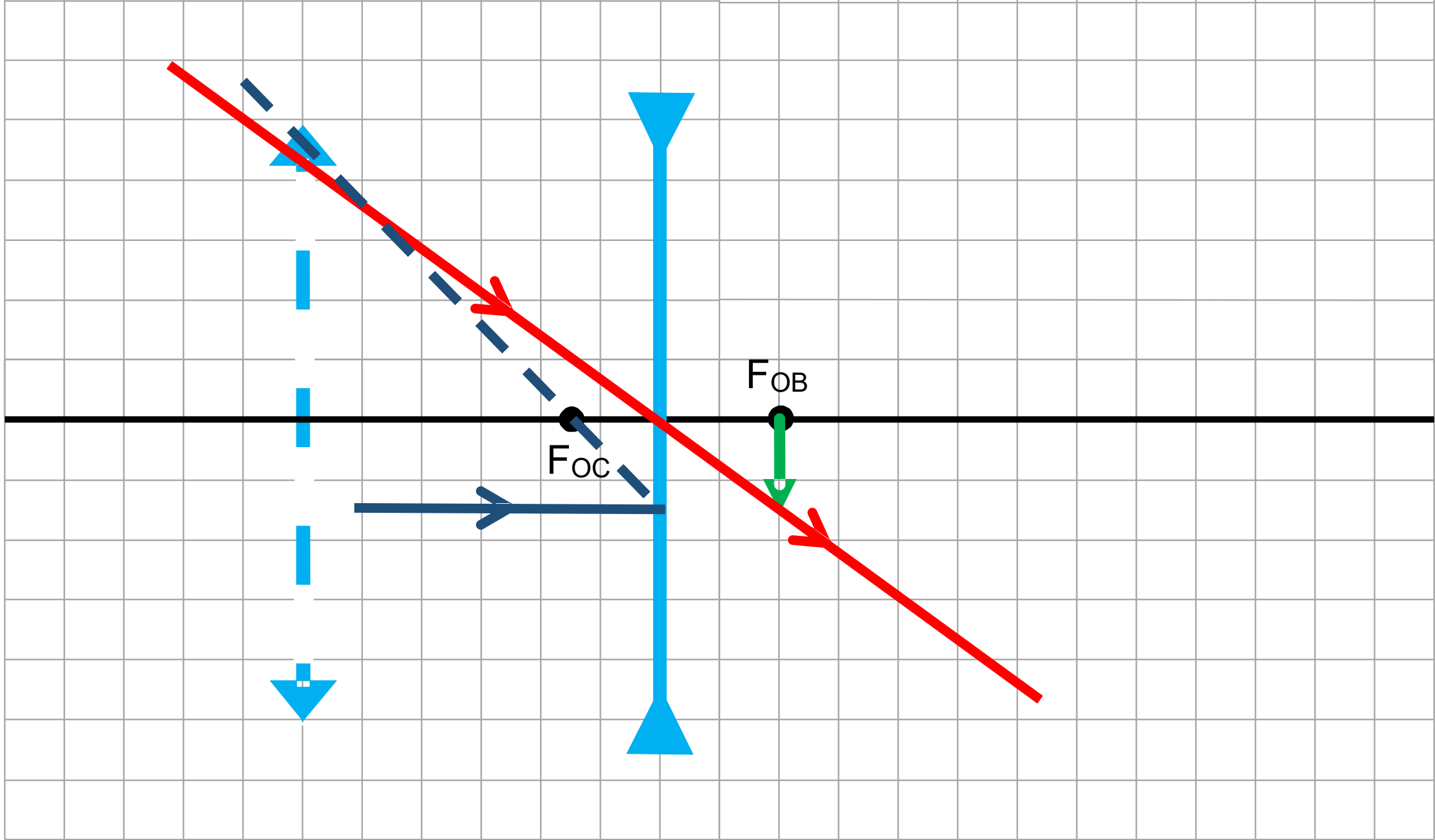


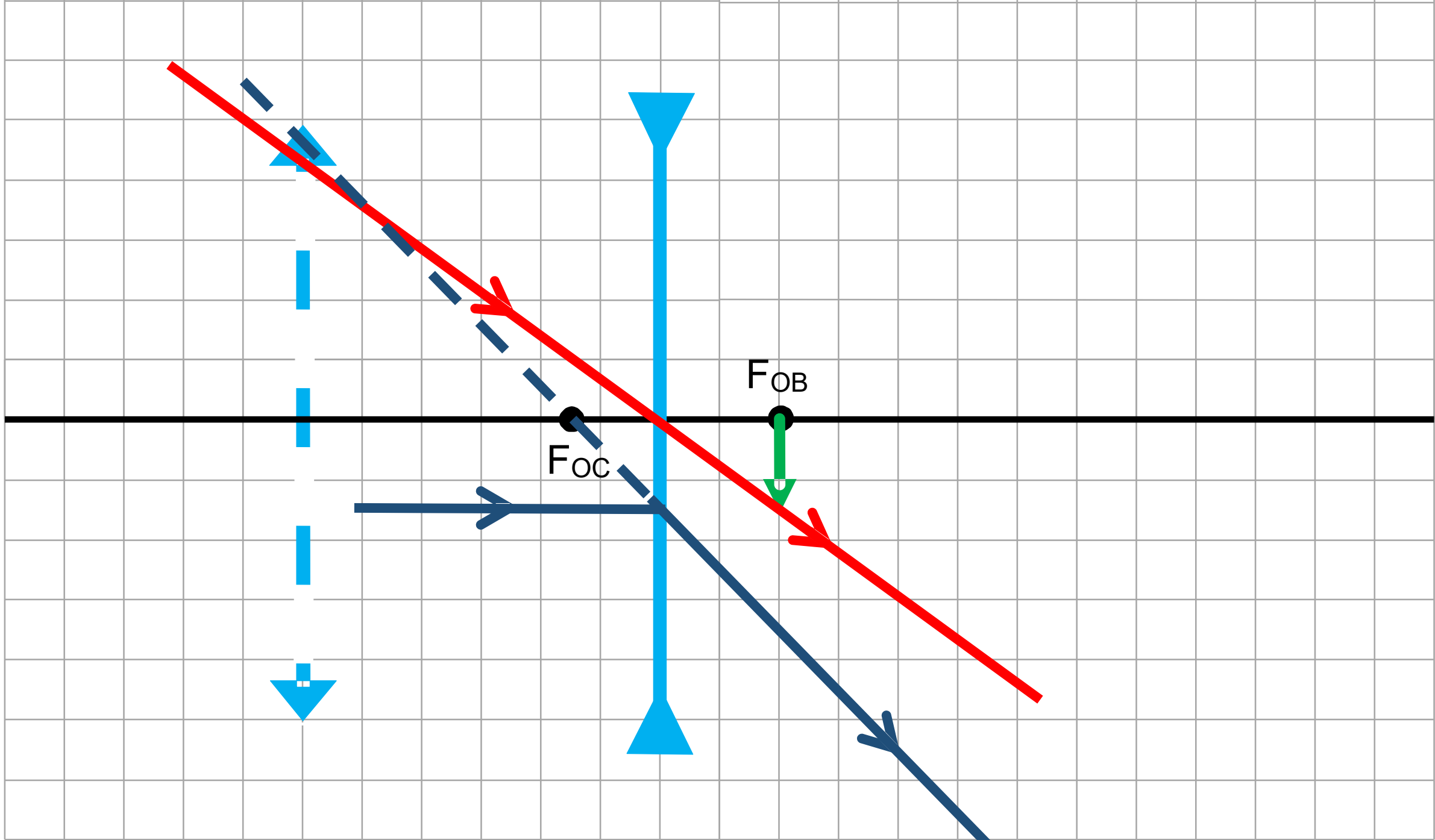


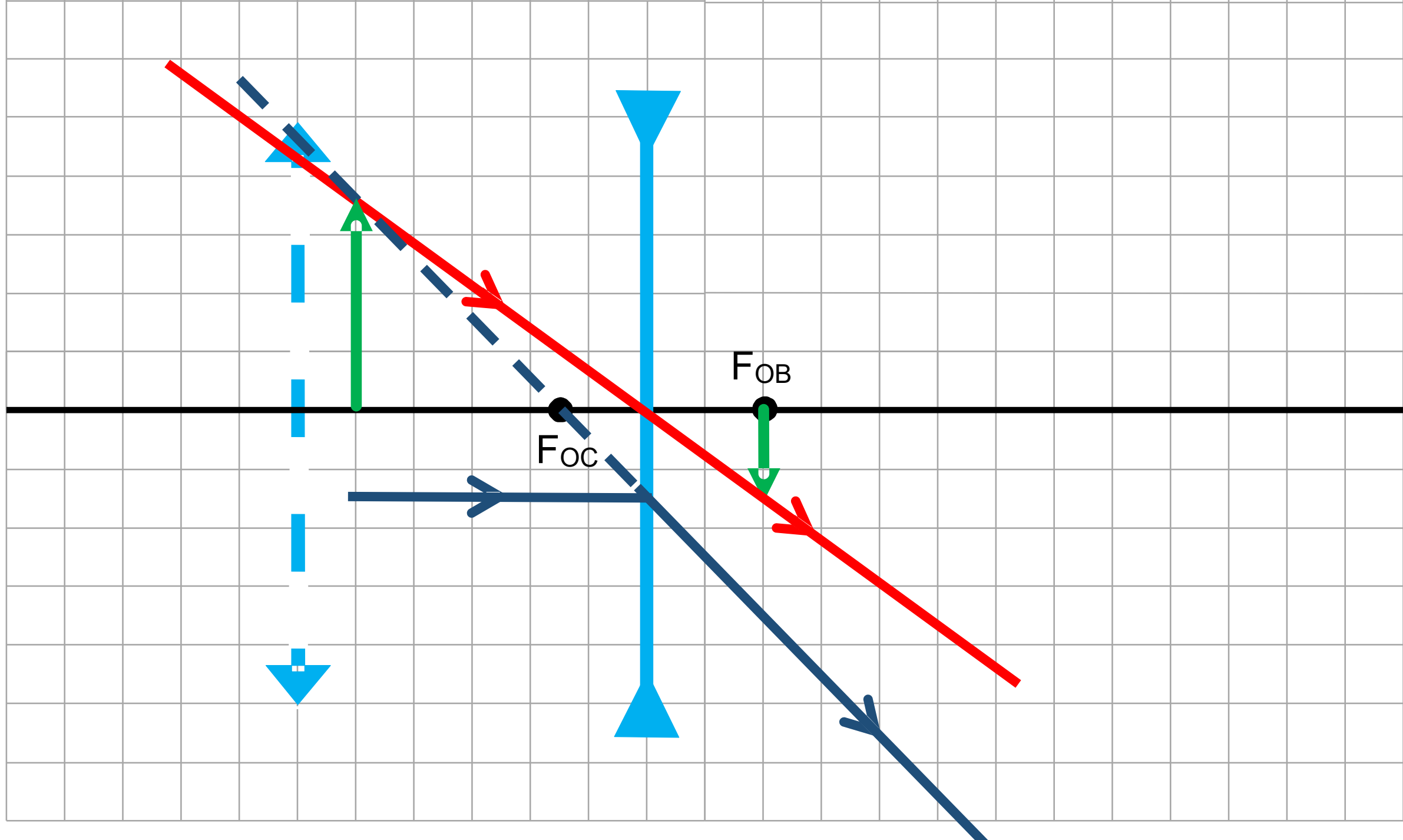


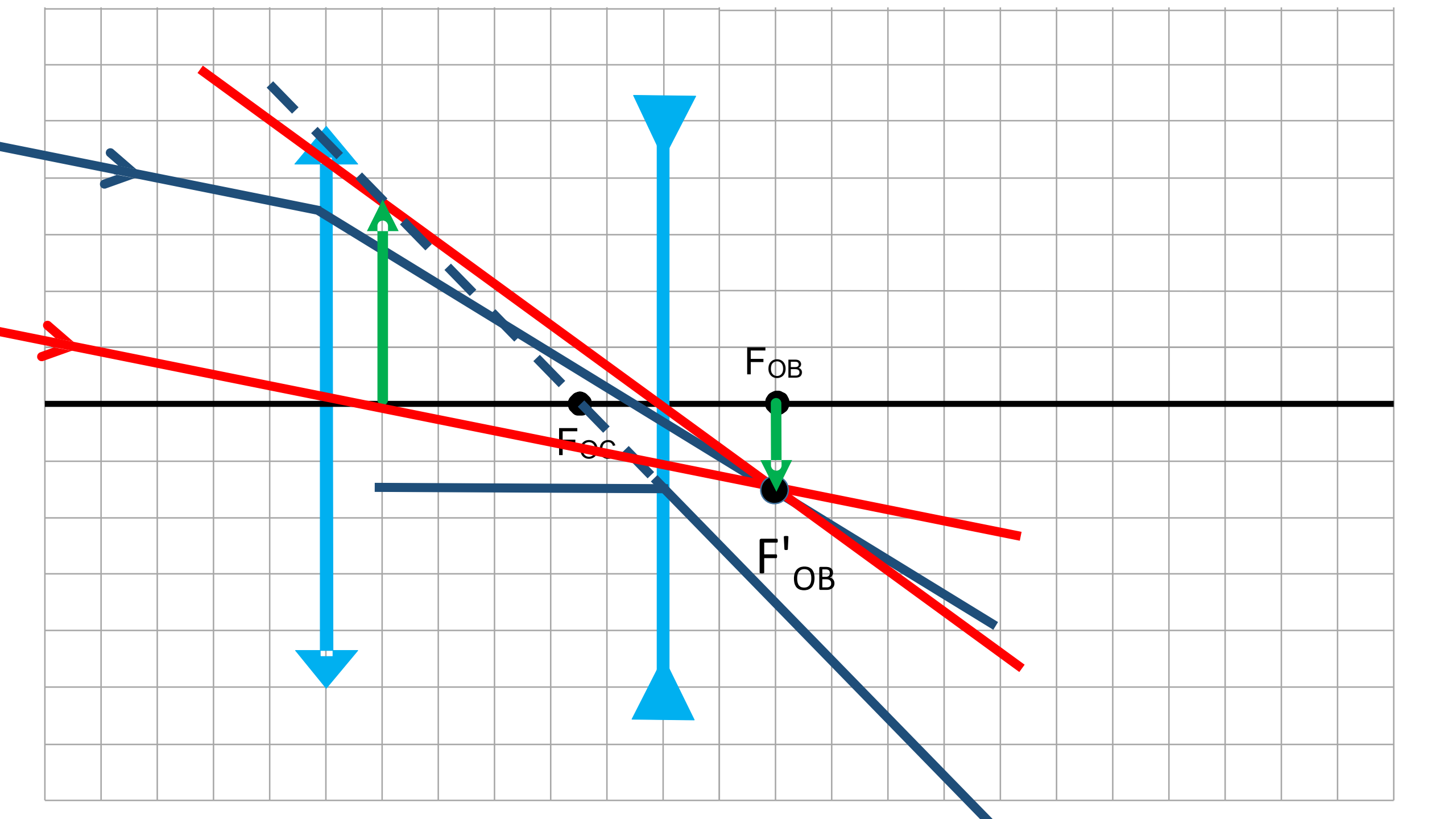












LUNETTA TERRESTRE

LUNETAS TERRESTRES

- IMAGEM:

LUNETAS TERRESTRES

- IMAGEM:
 - VIRTUAL

LUNETAS TERRESTRES

- IMAGEM:
 - VIRTUAL
 - DIREITA

LUNETAS TERRESTRES

- IMAGEM:

- VIRTUAL
- DIREITA
- QUANTO AO TAMANHO, DIZEMOS QUE HÁ UM AUMENTO ANGULAR, UMA VEZ QUE A IMAGEM FINAL É PONTUAL
- É CHAMADA DE LUNETAS TERRESTRES PELO FATO DE TER UMA IMAGEM FINAL DIREITA

Objetiva

Ocular



Curiosidade: a luneta de Galileu (luneta terrestre) foi criada em 1609, portanto durante as grandes navegações (por exemplo, no descobrimento do Brasil, em 1500) ainda não existiam, portanto é falso dizer que ela foi usada no descobrimento do Brasil.

AUMENTO ANGULAR

Para que o aumento de uma luneta ou telescópio seja o máximo possível, o foco imagem da objetiva deve coincidir com o foco objeto da ocular. Assim, o comprimento L da luneta será:

$$L = f_{OB} + f_{OC}$$

Note que se a ocular for uma lente divergente, o comprimento total será menor que o foco da objetiva uma vez que $f_{OC} < 0$.

Por fim, nesta condição, o aumento angular será:

$$M = \frac{f_{OB}}{f_{OC}}$$